

1.1

Lecture 1.1 Introduction

Maxwell-Boltzmann to Alder & Wainwright
via long-tailed probability distributions and
the scaling of atmospheric temperature and
winds.

Key References

Alder & Wainwright (1970), *Phys. Rev. A*, **1**, 18-21.

[emergence of fluid flow from molecular dynamics]

Schertzer & Lovejoy (1987), *J. Geophys. Res.*, **92**, 9693-9714.

[generalized scale invariance, statistical multifractals]

Tuck (2008), *Atmospheric Turbulence: A Molecular Dynamics Perspective*. Oxford University Press.

Tuck (2010, 2011), *Q. J. R. Meteorol. Soc.*, **136**, 1125-1144 & **137**, 275.

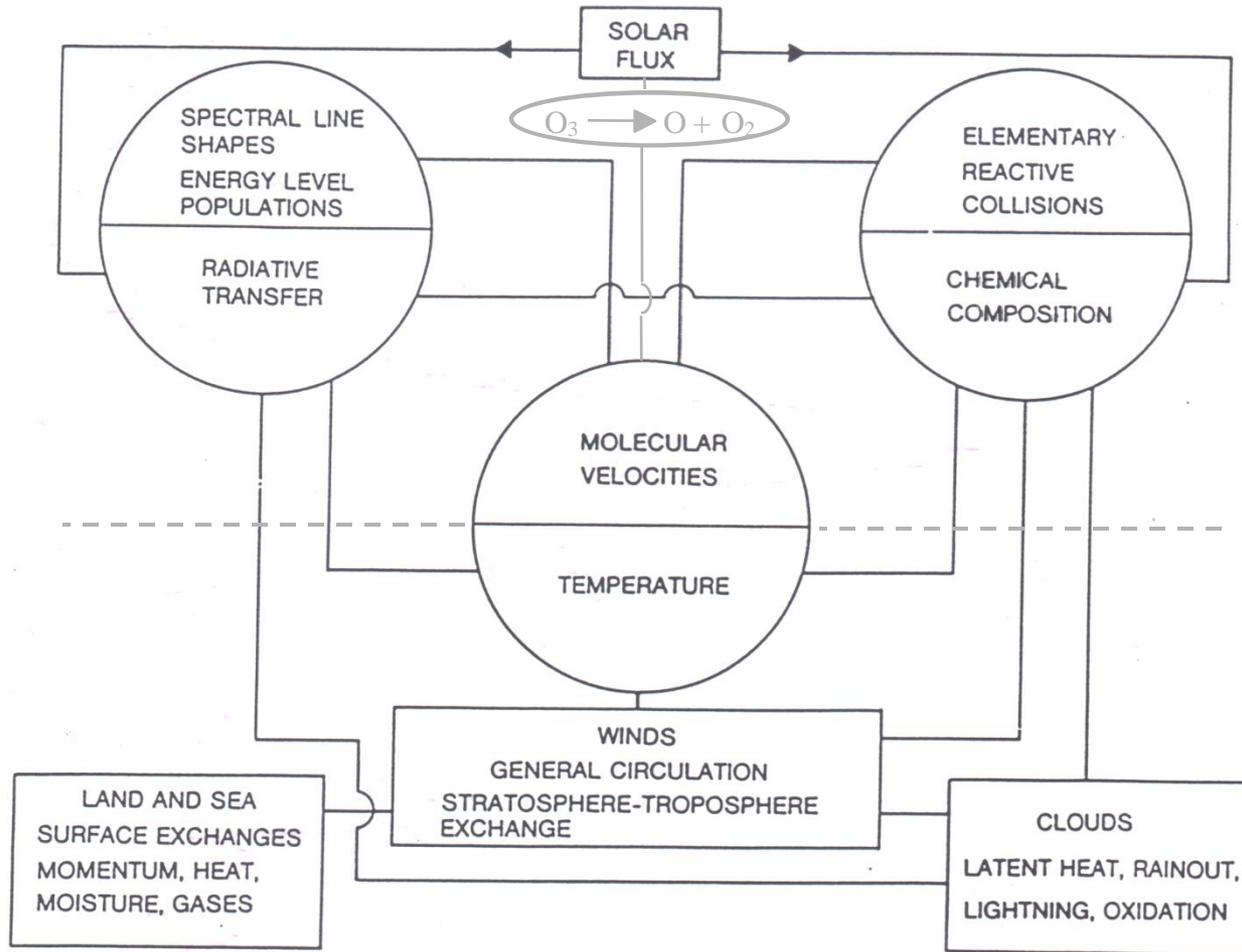
Schertzer & Lovejoy (2011), *Int. J. Bifurc. Chaos*, **21(12)**, 3417-3456.

[most recent review of generalized scale invariance]

Lovejoy & Schertzer (2012), *The Weather and the Climate: Emergent*

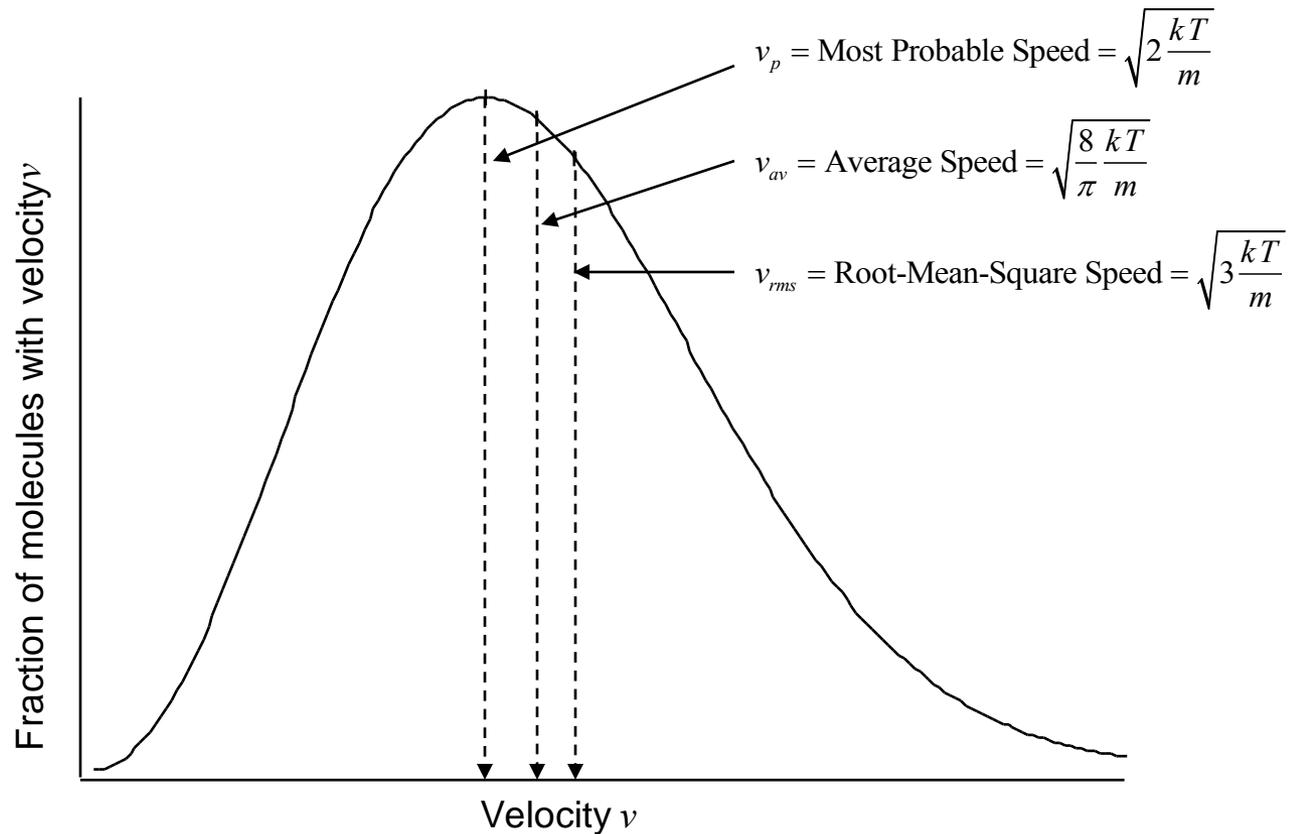
Laws and Multifractal Cascades. Cambridge University Press.

The correspondence and coupling of the microscopic and macroscopic processes in the atmosphere.

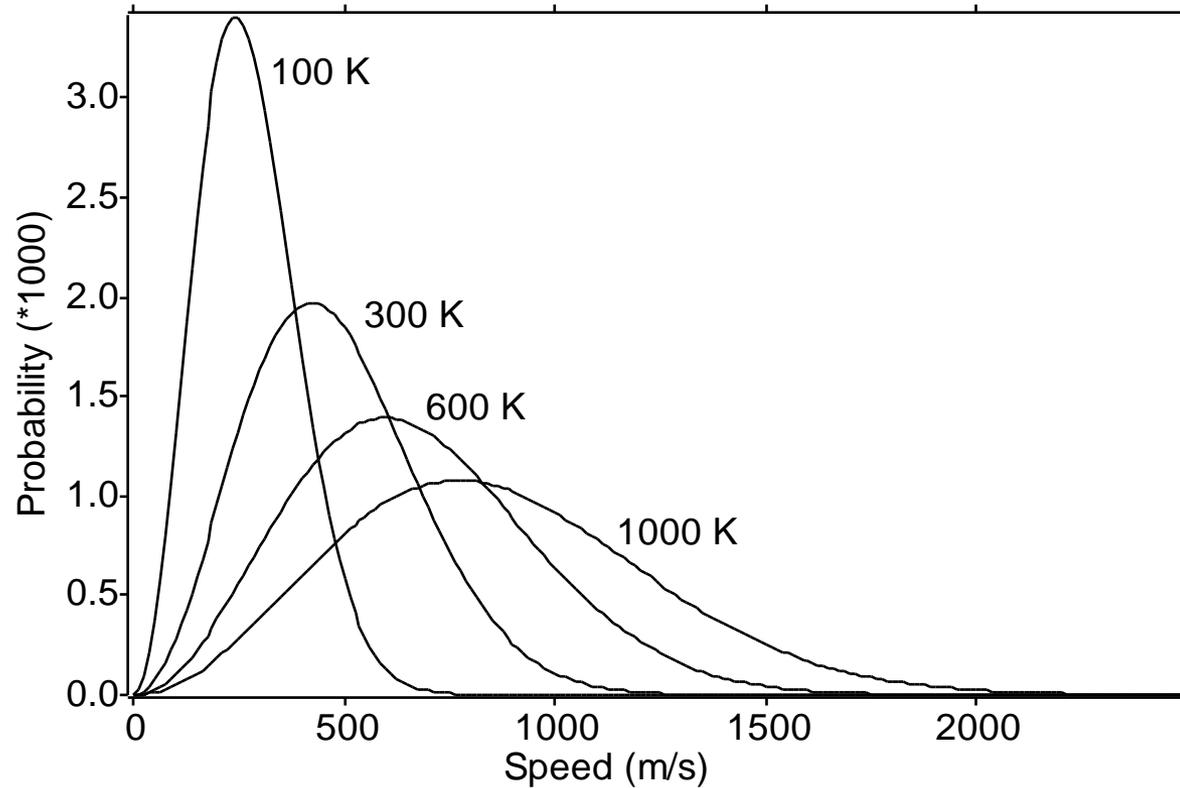


Maxwellian Velocity Distribution

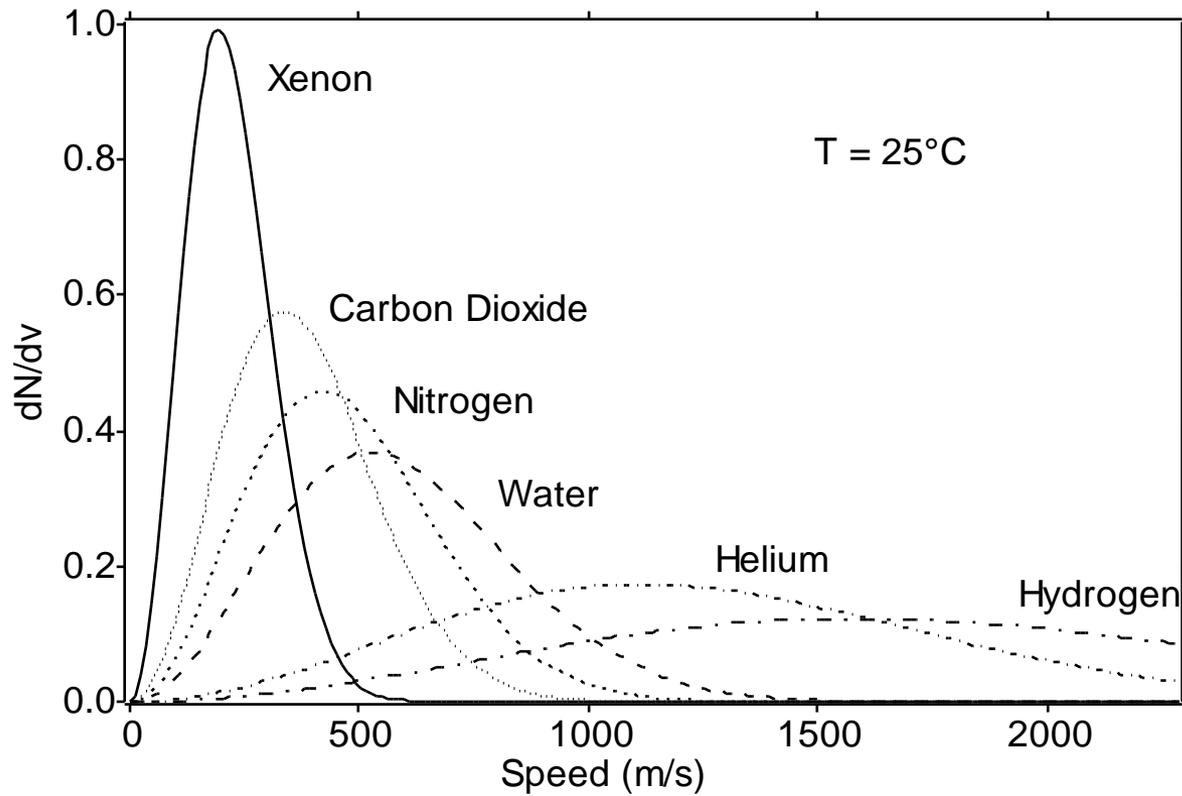
$$\frac{N_V(v)}{N} = 4\pi \left(\frac{m}{2\pi kT} \right)^{\frac{3}{2}} v^2 e^{-\frac{KE}{kT}}$$



Maxwellian speed PDFs, $m = 28$: T dependence



Maxwellian speed PDFs: mass dependence



DEFINITION

Atmospheric temperature is what is measured by your thermometer (meteorological).

The absolute temperature, T , of a system is the reciprocal of the derivative of the entropy, S , with respect to its energy, E :

$$\frac{dS}{dE} = \frac{1}{kT}$$

[Landau & Lifshitz, (1980), Statistical Physics, Course of Theoretical Physics, Vol. 5, 3rd ed., Chapters 1 - 3.]

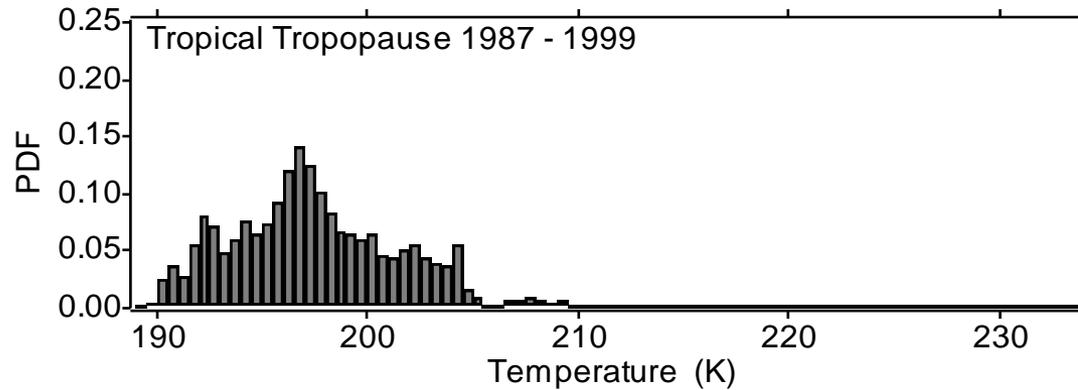
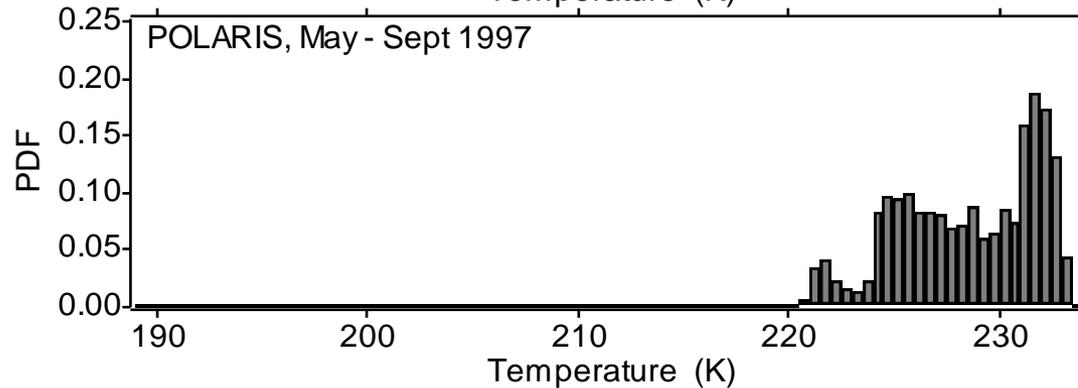
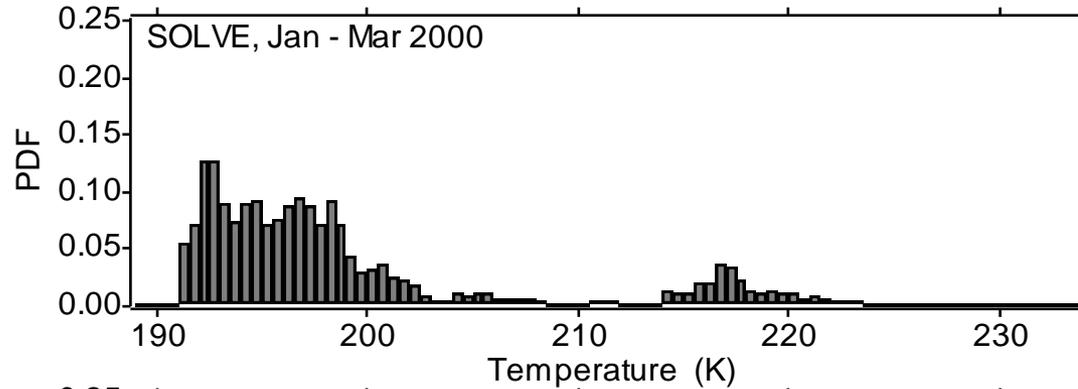
T is purely statistical, having strict meaning only for macroscopic bodies at equilibrium. One can of course observe with a calibrated thermometer. It averages over the velocity distribution of the molecules impinging upon it.

MICROSCOPIC VIEW OF TEMPERATURE

- The total kinetic energy of N classical particles is $3NkT/2$.
- In terms of distributions of molecular velocity v

$$\overline{v_x^2} = \sqrt{\frac{m}{2\pi kT}} \int_{-\infty}^{\infty} v_x^2 e^{-mv_x^2/2kT} dv_x = \frac{kT}{m}$$

Long-tailed PDFs of temperature: millions of 5 Hz points from scores of ER-2 flight segments, Arctic summer 1997 & winter 2000, ER-2 & WB57F tropical tropopause 1987-1999.



Scale Invariance and Statistical Multifractality

$$S_q(r; \Psi) = \langle |\Psi(x+r) - \Psi(x)|^q \rangle$$

↑ q^{th} order structure function S of variable $\Psi(x)$

If a plot of $\log S_q$ vs. $\log(r)$ is linear with slope $\zeta(q)$, then $\zeta(q)$ is a scaling exponent for $\Psi(x)$, which therefore has scale invariance and power law PDFs.

Define

$$H(q) = \zeta(q)/q$$

Further define

$$H = H_q + K(q)/q$$

We will be interested in $K(q)$ and H

To obtain $K(q)$, consider $\Psi(x)$ to have been observed at finite intervals $x = 1, 2, 3, \dots, x_{\max}$ and define:-

$$\varepsilon(1, x) = \{ |\Psi(x+1) - \Psi(x)| \} / \langle |\Psi(x+1) - \Psi(x)| \rangle$$

for $x = 1, 2, 3, \dots, x_{\max}$

$$\varepsilon(r, x) = (1/r) \sum_{j=x}^{x+r-1} \varepsilon(1, j)$$

for $x = 1, 2, 3, \dots, x_{\max} - r$

then a plot of $\log \langle \varepsilon(r, x) \rangle^q$ vs $\log r$ has slope $-K(q)$ and a plot of $-K(q)$ vs q shows a convex function with $K(0) = K(1) = 0$.

We can now note equivalences between scale invariance and statistical thermodynamics.

Formal equivalences between **scale invariant** (r.h.s.)
and **statistical thermodynamic** (l.h.s.) variables

$$T = 1/qk_{\text{Boltzmann}}$$

temperature

$$f = e^{-K(q)}$$

partition function

$$G = -K(q)/q$$

Gibbs free energy

This offers possible links:

molecular scale



statistical thermodynamics



macroscopic scale invariant observables

Basic scaling relations for atmospheric turbulence

Consider horizontal velocity v with fluctuations Δv over horizontal length interval Δx and vertical interval Δz :

$$\Delta v = \varnothing(\text{horiz}).(\Delta x)^{H(\text{horiz})}$$

$$\Delta v = \varnothing(\text{vert}).(\Delta z)^{H(\text{vert})}$$

where $\varnothing(\text{horiz})$ is the turbulent energy flux ε in the vertical and $\varnothing(\text{vert})$ is the turbulent energy flux; $H(\text{horiz})$ and $H(\text{vert})$ are the associated power law exponents. The Kolmogorov law for isotropic turbulence is obtained by setting

$$\varnothing(\text{horiz}) = \varnothing(\text{vert}) = \varepsilon^{1/3} \quad \text{and} \quad H(\text{horiz}) = H(\text{vert}) = 1/3$$

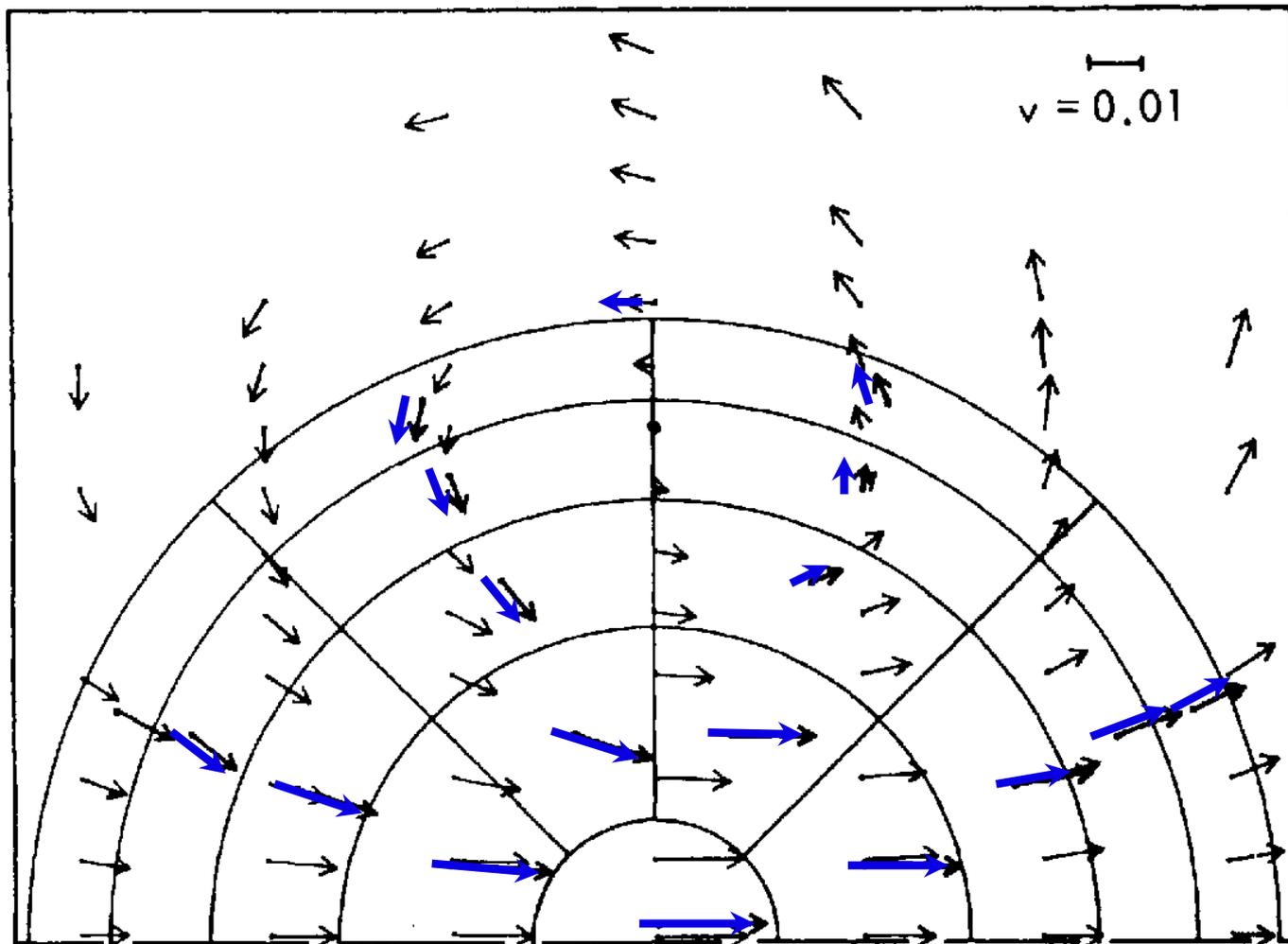
However, as we shall see later, this is at odds with observations. We therefore **assume** via **dimensional analysis** relations that preserve scaling while permitting anisotropy:

$$\varnothing(\text{horiz}) = \varepsilon^{1/3} \quad \text{implying} \quad H(\text{horiz}) = 1/3$$

$$\varnothing(\text{vert}) = \eta^{1/5} \quad \text{implying} \quad H(\text{vert}) = 3/5$$

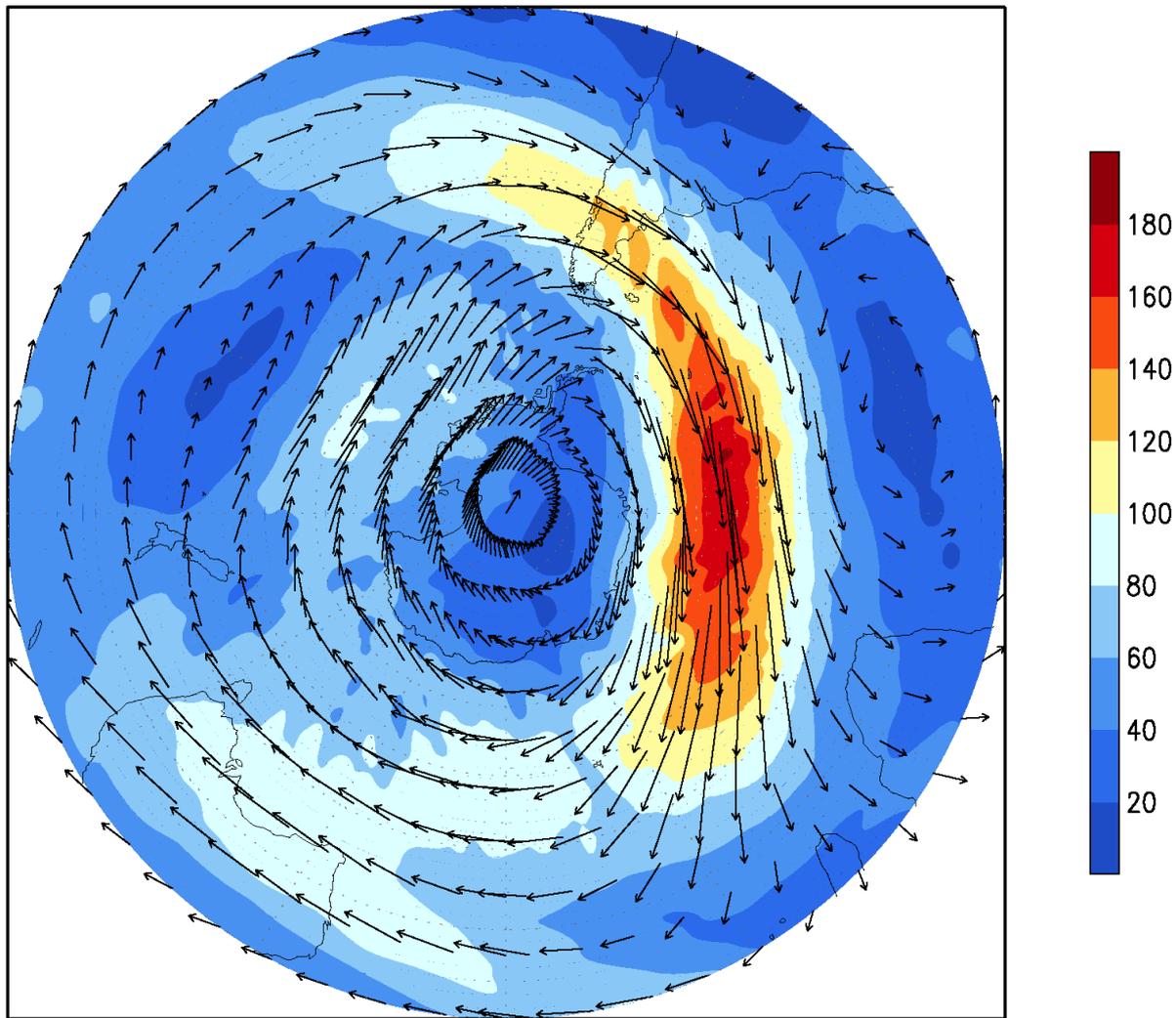
where energy flux ε (m^2s^{-3}) dominates in the horizontal and buoyancy variance flux η (m^2s^{-5}) dominates in the vertical.

Alder & Wainwright (1970): **molecular dynamics** simulation of a flux applied to an equilibrated Maxwellian population results in the emergence of vortices on scales of 10^{-12} seconds & 10^{-8} metres.



NOAA-NCEP GFS 0.5° RESOLUTION ANALYSIS, FLECHES & ISOTACHS

isotachs (m/s) 2 mb Fri 12Z 06jul2007



100

1.1 Summary

- Organized flow emerges from a randomized (thermal) population of Maxwellian 'billiard balls' subject to an anisotropic flux simulated by molecular dynamics, on very short time and space scales.
- Observed temperature distributions in the atmosphere are not of Maxwell-Boltzmann character.
- To accommodate the observed non-M-B PDFs, we must lift the assumption of isotropy in the atmosphere.
- The atmosphere has permanent anisotropies: gravity, planetary rotation, solar beam, planetary surface and, for any one scale except the great circle, larger scale winds.
- Abandoning isotropic motions on any scale yields predictions:

$$H(\text{horiz}) = 1/3$$

$$H(\text{vert}) = 3/5$$

and so $H(\text{horiz})/H(\text{vert}) = 5/9$

*How do these work out in practice? See next lecture