



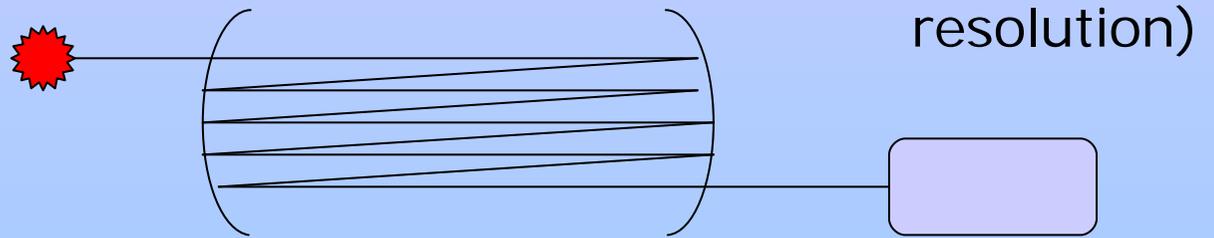
**VALIDATION OF SCIAMACHY WATER VAPOR AND METHANE  
PROFILES BY BALLON-BORNE IN-SITU MEASUREMENTS WITH THE  
“CHILD” SPECTROMETER ONBOARD TRIPLE**

Wolfgang Gurlit, Carsten Gieseemann, Volker Ebert,  
Rainer Zimmermann, and John P. Burrows  
supported by TRIPLE (A. Engel, C. Schiller, F. Stroh)

# CHILD key components & function

Compact High-altitude In-situ Laser Diode spectrometer

monochromatic  
tunable  
light source



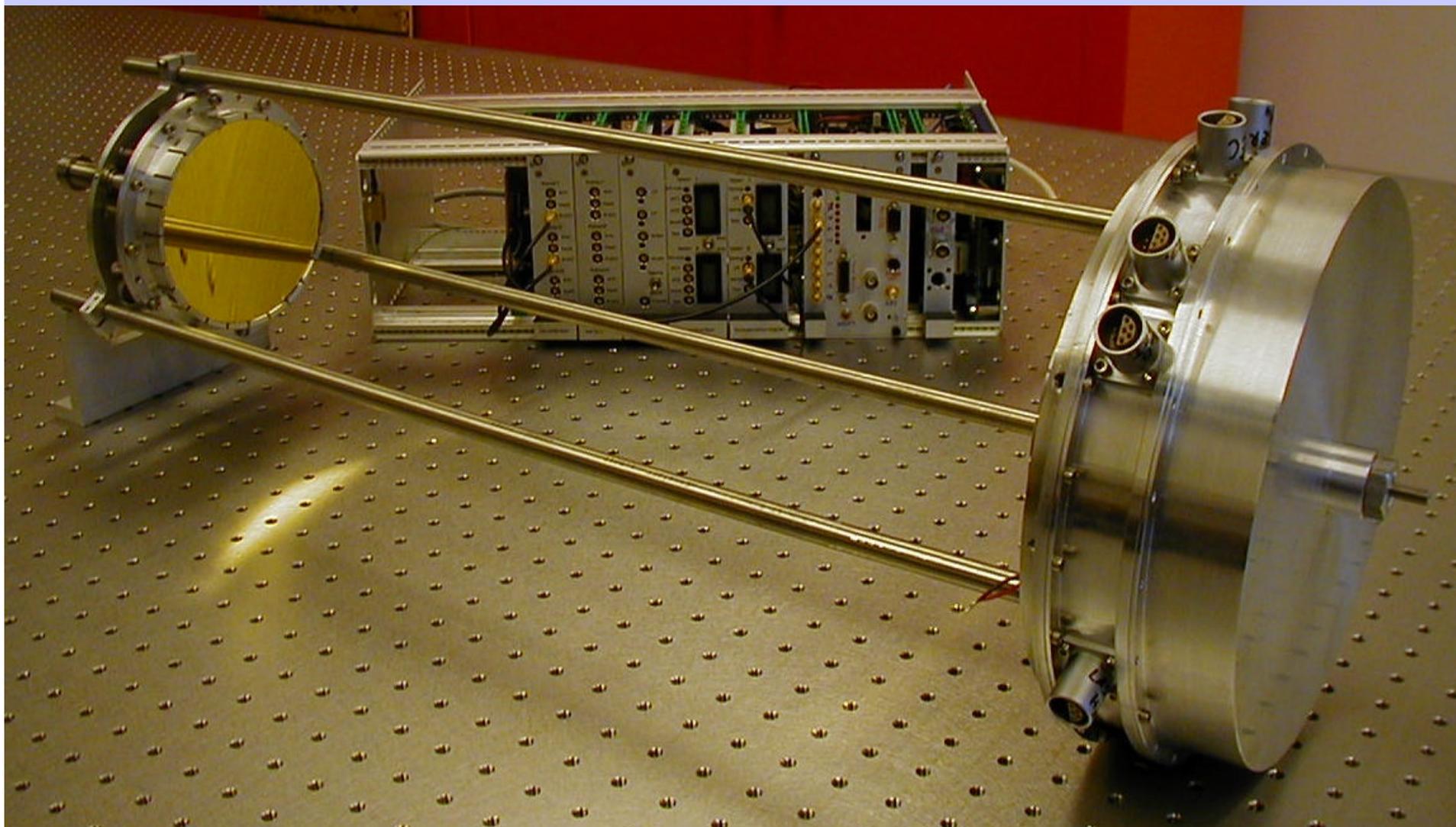
detector  
(no spectral  
resolution)

absorption path  
(multireflection)  
open to the  
atmosphere

## Key features:

- in-situ measurement delivers altitude resolution of a few meters
- fast response time (seconds)
- absorption cell open to the atmosphere avoids sampling errors (no deposition by adsorption, condensation)
- same NIR absorption features as SCIAMACHY, uncertainty of spectroscopic parameters removed from validation
- high spectral resolution (laser linewidth < 30 MHz)
- low weight, low power consumption, small size
- combined TRIPLE payload enables sensor intercomparison

## CHILD Herriott-cell and control unit



CHILD specification	CH4	H2O
wavelength	1648.2 nm	1392.53 nm
absorption pathlength	74.0 m	35.9 m
scan-range	+/- 8.5 GHz	+/-7.5 GHz
MDA	4 x 10 E-5	4 x 10 E-4

weight	25 kg with batteries
power consumption	40 watts
telemetry	RS-232 9600 bps

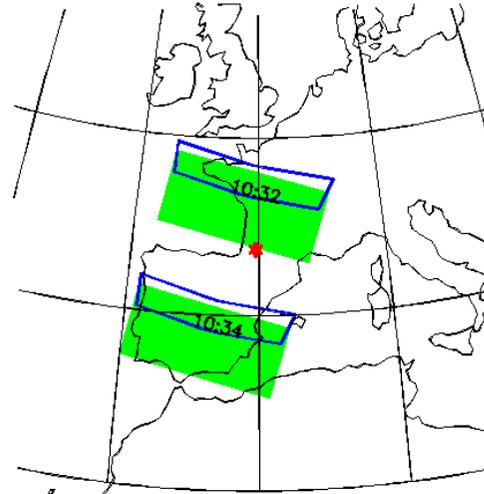
Table 2: Campaigns with CHILD onboard the TRIPLE gondola

Aire sour `l Adour fall 2001 (France, mid-latitude)	technical test
Aire sour `l Adour fall 2002 (France, mid-latitude)	ENVISAT validation, Flight 24.09.2002
Kiruna spring 2003 (Sweden, high-latitude)	ENVISAT validation Flight 06.03.2003
Kiruna summer 2003 (Sweden, high-latitude)	ENVISAT validation Flight 09.06.2003

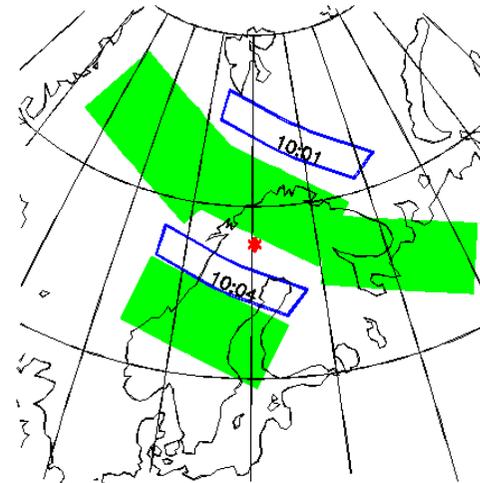
# Validation flights with TRIPLE & SCIAMACHY overpasses



SCIAMACHY overpasses on 24-SEP-2002

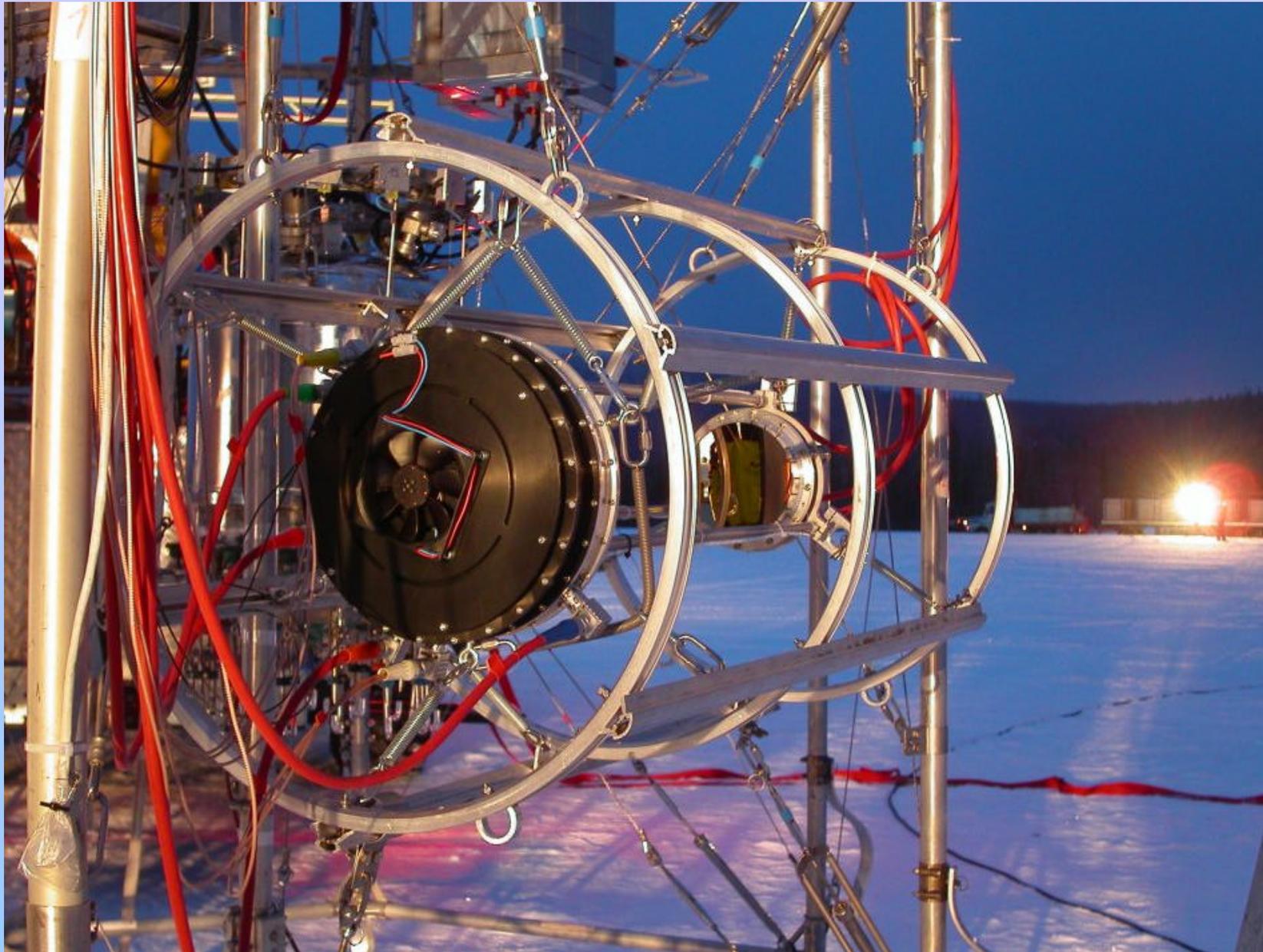


SCIAMACHY overpasses on 06-MAR-2003



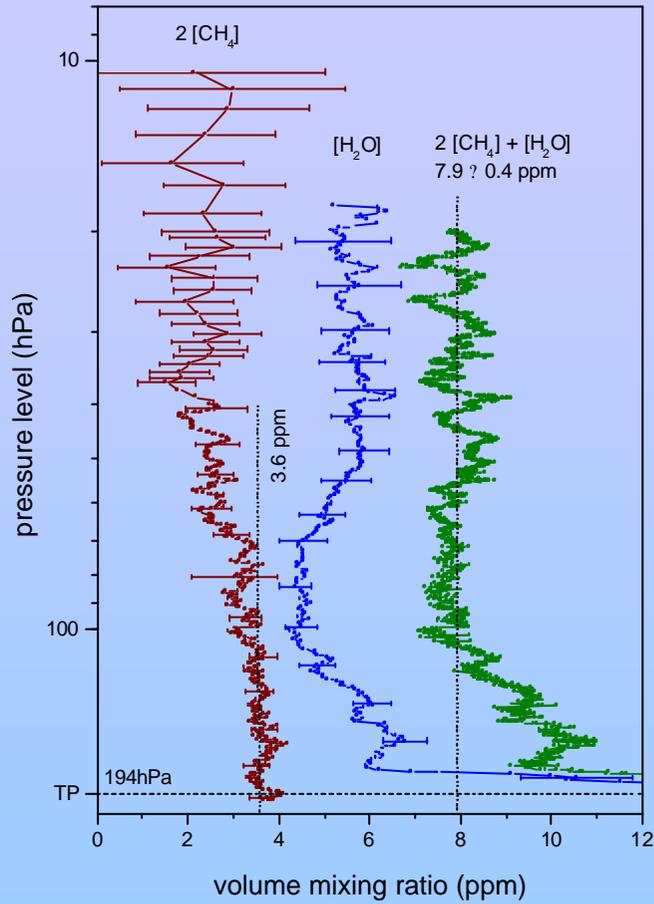


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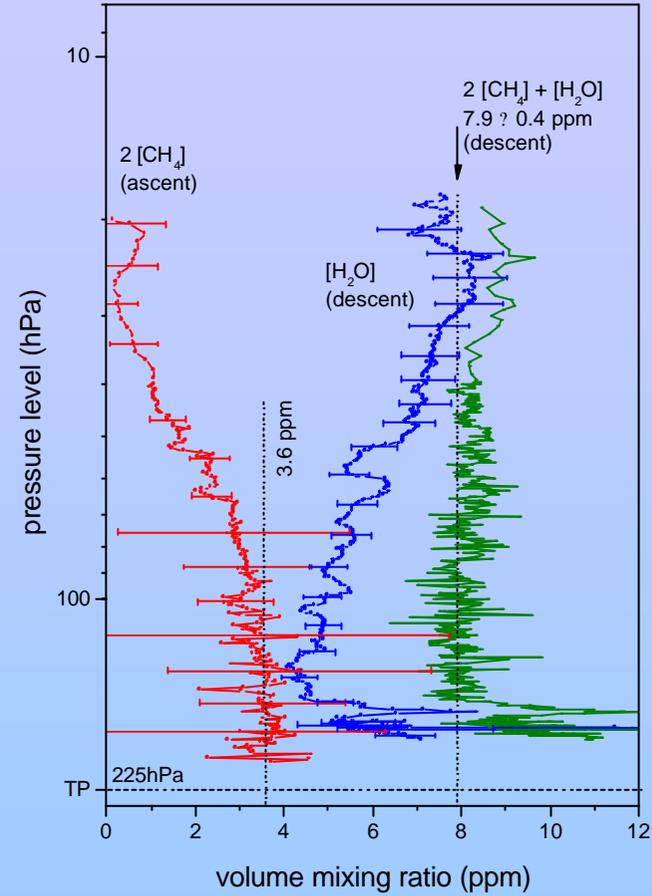


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# Flight Results



ASA 24 September 2002



Kiruna 06 March 2003

## Conclusion

- three successful validation flights with TRIPLE so far
- SCIAMACHY level-2 data (to be validated) not yet available
- intercomparisons of validation sensors are possible and useful

**SCIAMACHY SOLAR IRRADIANCE VALIDATION USING  
RADIOMETRIC  
CALIBRATION OF BALLOONBORNE SPECTROMETERS**

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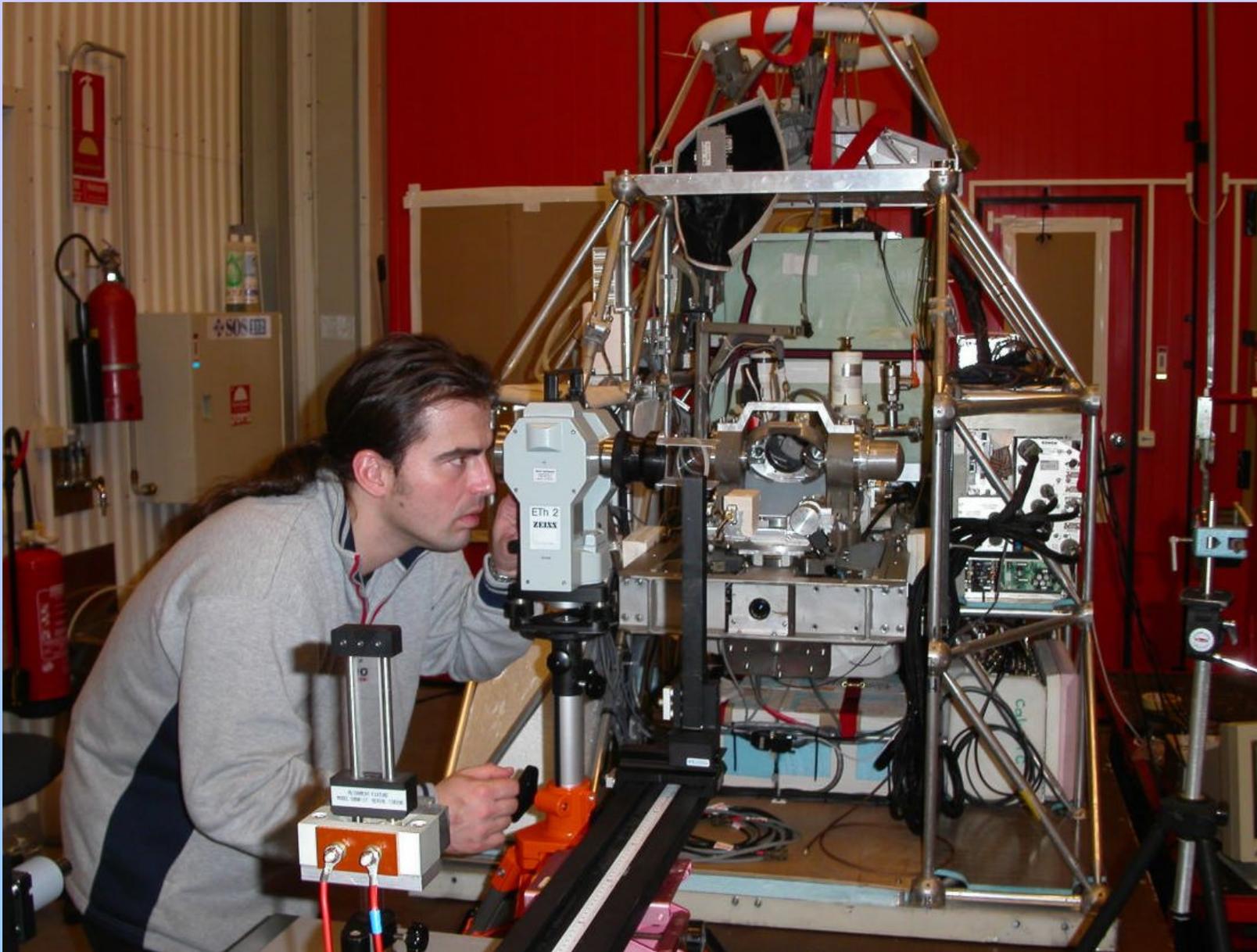
*Partners : Claude Camy-Peyret, CNRS-LPMA,  
Hartmut Boesch, Klaus Pfeilsticker, University of Heidelberg*

## General Concept

- use existing balloonborne instruments for additional measurements
- LPMA and DOAS = combined payload for tracegas measurements in occultation geometry, covering SCIAMACHY channels
- LPMA : Fourier Transform IR spectrometer
- DOAS : two channel UV-VIS grating spectrometer
- use same entrance optics to track the solar disc during sunrise or sunset from stratospheric float position
- solar irradiance measurements require radiometric calibration of LPMA-DOAS
- after correction for atmospheric effects (absorptions + scattering), top-of-atmosphere solar irradiance for validation of the corresponding SCIAMACHY level-1 product

## Calibration Sources & Procedures

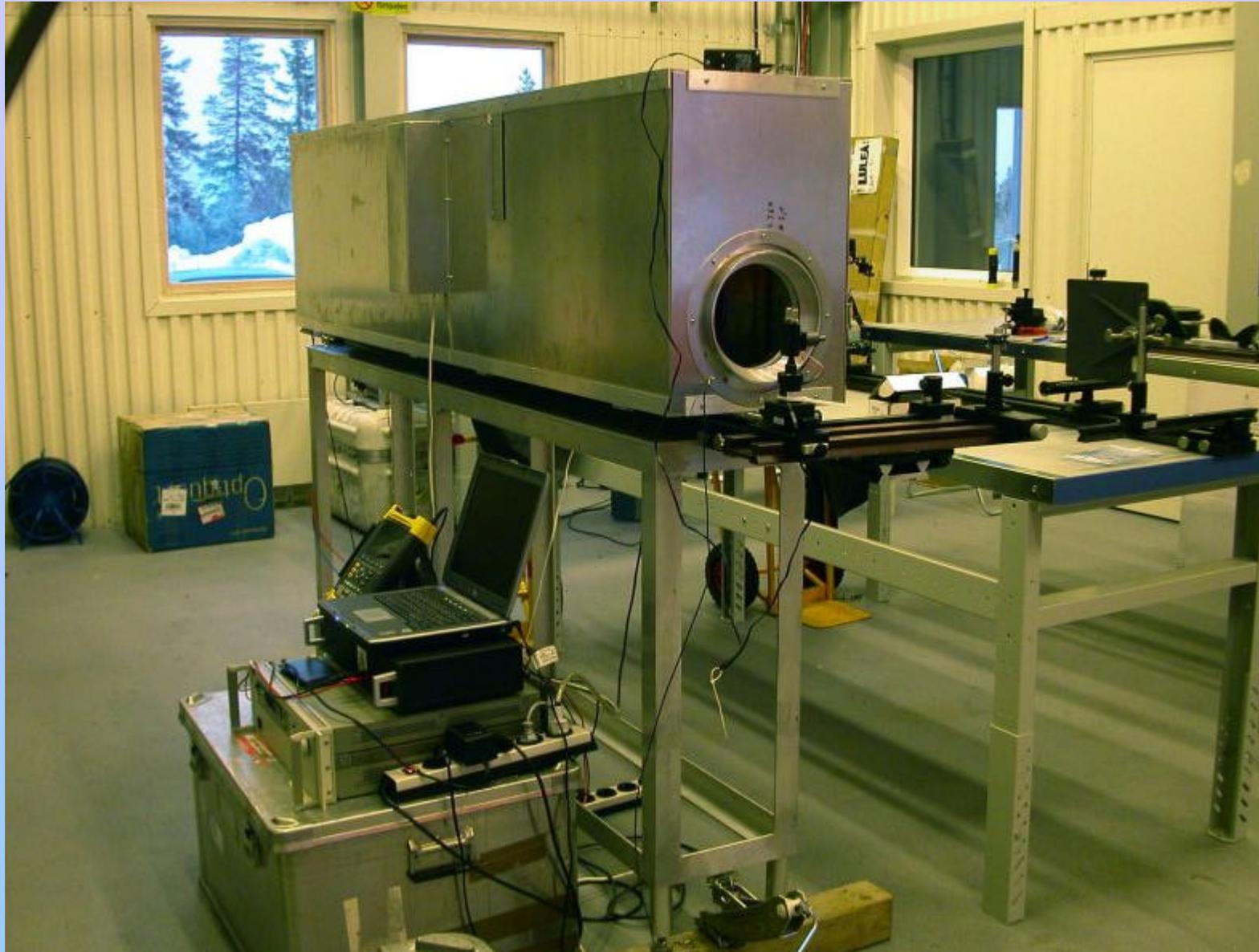
- absolute calibration of LPMA-DOAS before each flight
- calibration has to be done on the campaign site because integration of optical components (LPMA, DOAS and suntracker) takes place in the field
- setup of a mobile optical calibration lab in the field
- first campaign using this approach:  
Kiruna spring 2003 ENVISAT validation campaign  
with two LPMA-DOAS flights
- different calibration procedures for the LPMA  
and the DOAS spectrometer



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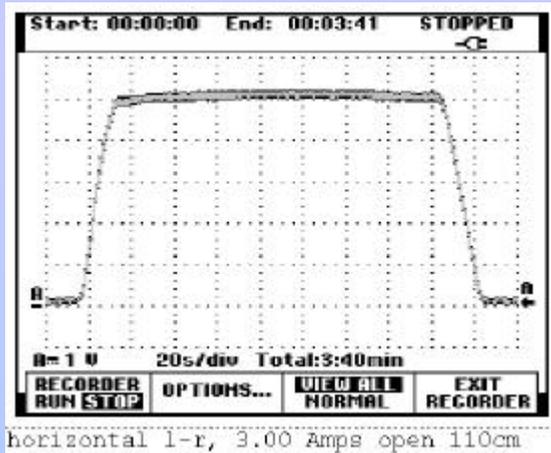
## Calibration of LPMA (Fourier Transform IR spectrometer)

- requires a collimated optical input
- advantage of collimated calibration source:  
produces same results when used at different distances  
(taken aside absorptions along the lightpath)
- source with 1000 watts initial optical power and beam forming  
with mirror optics specially developed for this purpose
- selected lamps, special delivery by Phillips
- solar simulation, solar disc appears at a viewing angle of  
 $0.54^\circ$
- highly uniform light beam, diameter approx. 16 cm



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## Results of LPMA radiometric calibration



beam cross-section

- usable radiation is produced from UV (300 nm) to IR (beyond 3000 nm)
- stability and reproducibility better than 0.5 %
- distance to target not critical (measured variation of 1% when distance to target was varied from 1.2 to 2.2 meters)
- optical output sufficient to operate LPMA in flight gain
- for characterisation, SCIAMACHY optical stimuli are available at IUP University of Bremen, calibrated results expected 08/2003

## Calibration of the DOAS spectrometer

- no collimated source required
- NIST standard lamp (1000 watts QTH quartz tungsten halogen) used as the calibration source



## Setup & procedure for DOAS calibration

- optical axis of DOAS aligned with NIST lamp
- NIST lamp mounted on optical rail (3 meters) with center of filament aligned with DOAS center of field of view
- for absolute calibration distance between source and target must be known
- this distance could not be measured directly using e.g. a ruler
- lamp could be moved on the optical axis to produce different intensities according to inverse square law
- intensity per unit area varies in inverse proportion to the square of distance  $X$  from source to target

## Calculation of unknown distance X between source and target

$$X = [d_1 (E_1/E_2)^{1/2} - d_2] / [1 - (E_1/E_2)^{1/2}]$$

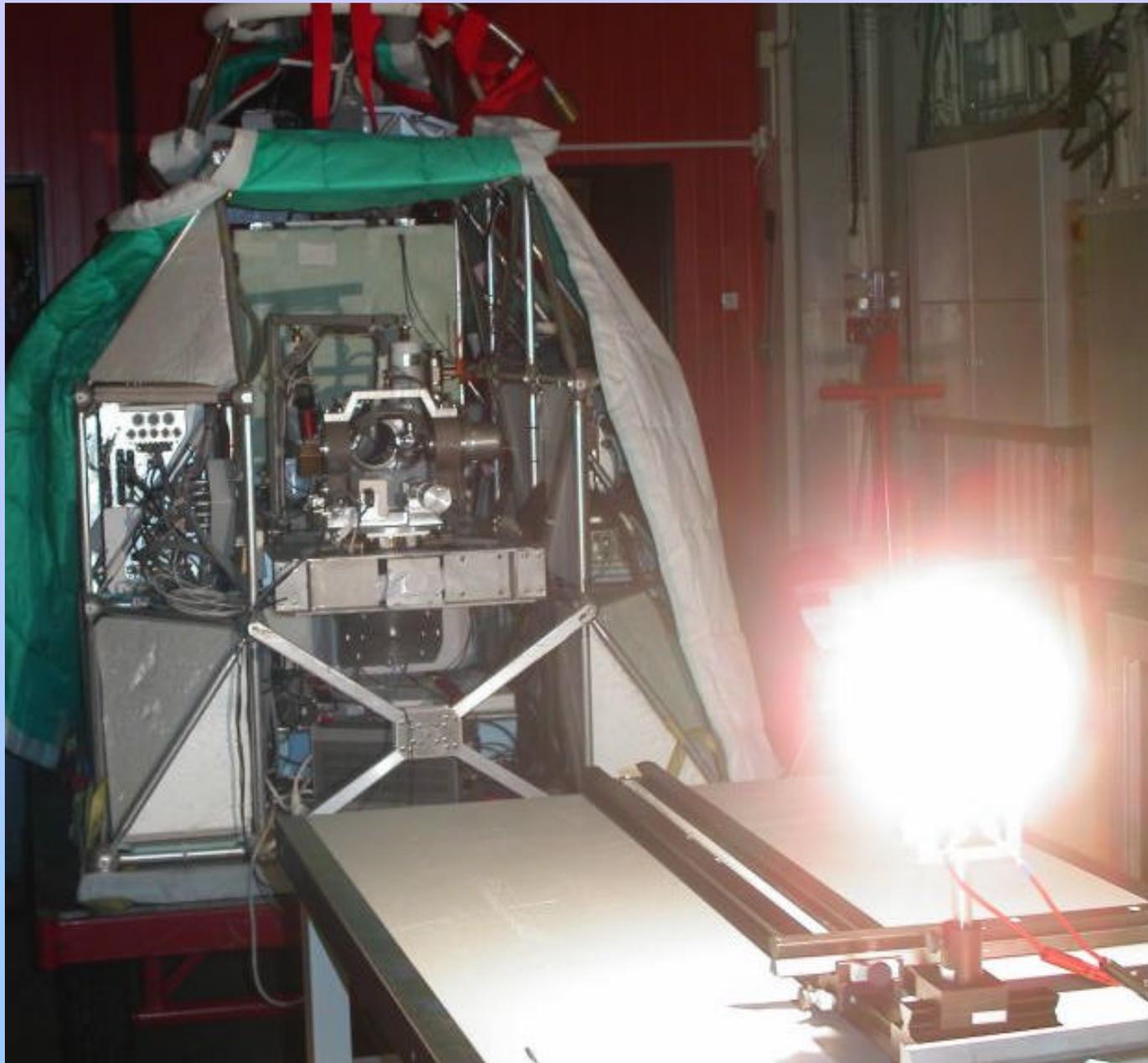
$d_1$  and  $d_2$  are two known variations of the distance measured relatively to point X

$E_1$  and  $E_2$  are the corresponding intensities (instrument readings)

- with known distance X spectral irradiance of NIST lamp on DOAS can be calculated
- instrument's spectral response then is known on absolute scale

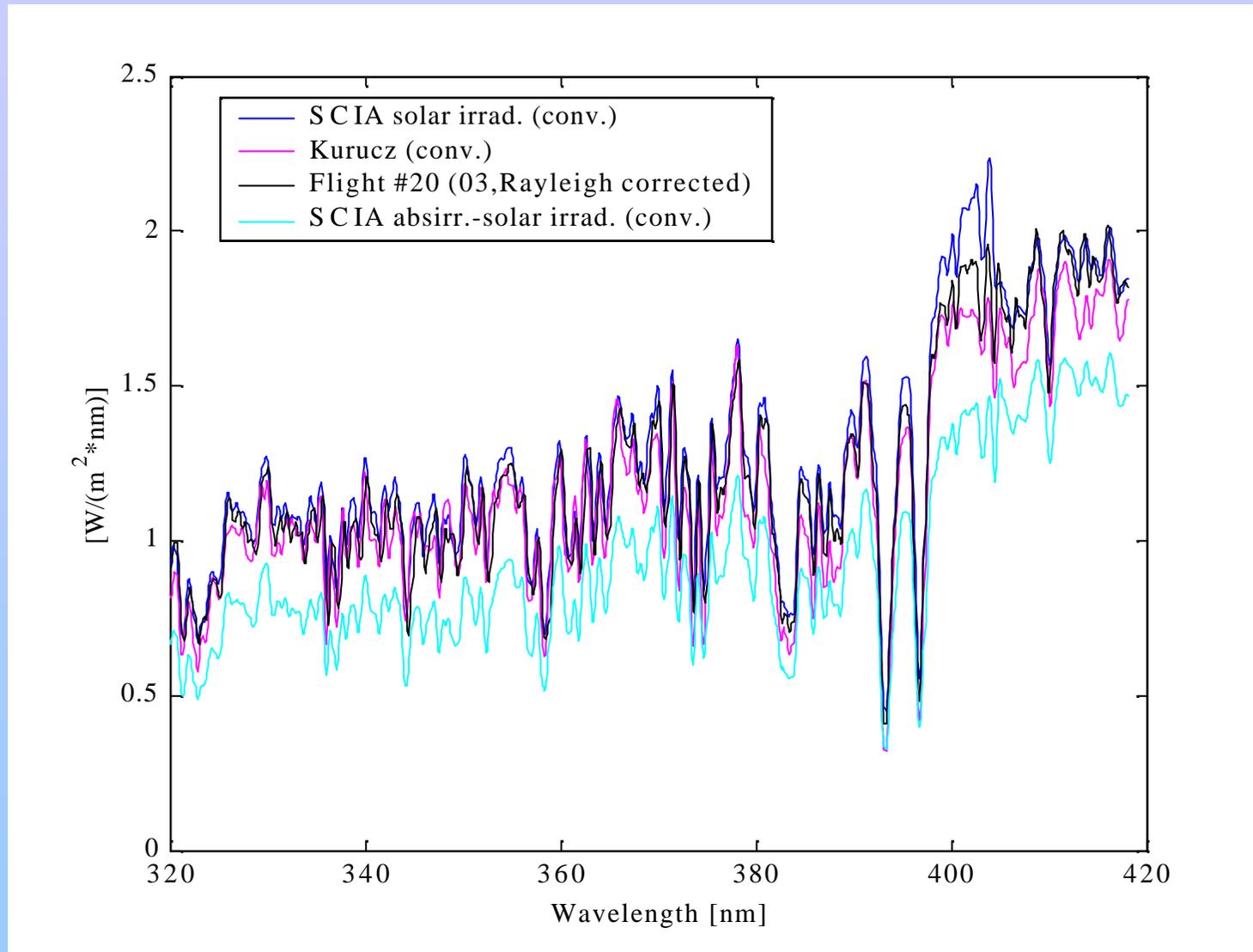
## ...a closer look at the inverse square law

- 1) source must be point source, not an area source  
(source must be completely inside DOAS FOV, 3° for UV)  
long initial distance of 2 meters needed, long integration times
- 2) optical axis must be known exactly. Suntracker could not be used due to limited dynamic range, manual optimisation.  
DOAS optical telescopes were not perfectly aligned with the suntracker. Quantitative measurement of this misalignment with collimated source
- 3) linearity of the instrument to be calibrated. DOAS showed some linearity problems for short integration times (used during flight)  
clarification required additional measurements



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## Solar irradiance results with calibrated DOAS measurements



## Error budget of calibrated irradiance spectra from DOAS

- 1) uncertainty of irradiance standards used (~1%)
- 2) standard drift and ageing errors (0.5% per 24 h operation)
- 3) positioning errors in optomechanical setup, measured by reproduceability (1%)
- 4) errors by not completely satisfying inverse square law (DOAS linearity errors 3%, definition / reproduceability of optical axis 2%)

errors are in part systematic (linearity error, integration time error) and can be reduced by further investigation

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