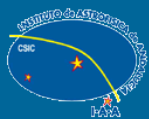
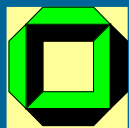
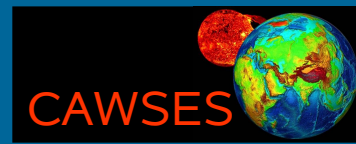

Modelling disturbed stratospheric Chemistry during solar induced NO_x Intrusions observed by MIPAS-ENVISAT

Thomas Reddmann, Thomas v. Clarmann, Wolfgang Kouker, Roland Ruhnke, Gabriele Stiller, Stefan Versick, and Roland Uhl
Institut für Meteorologie und Klimaforschung, Universität Karlsruhe und
Forschungszentrum und Universität Karlsruhe, Germany

Bernd Funke and Manuel Lopez-Puertas,
Instituto de Astrofisica de Andalucia, Granada



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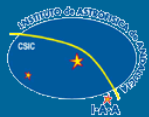
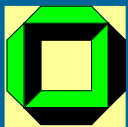


Motivation

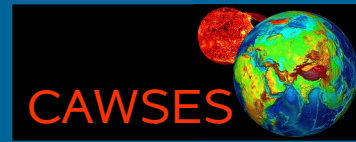
- Possible solar influence on climate via amplification mechanism within the stratosphere
- Interference through radiative active gases, especially ozone
- **Direct effects through heating**
- **Indirect effects through change in chemistry**
 - Direct by radiation and particle interaction
 - Indirect by mesospheric NO_x intrusions
- Change in dynamics to be translated to the troposphere

Evaluation through combination and comparison of

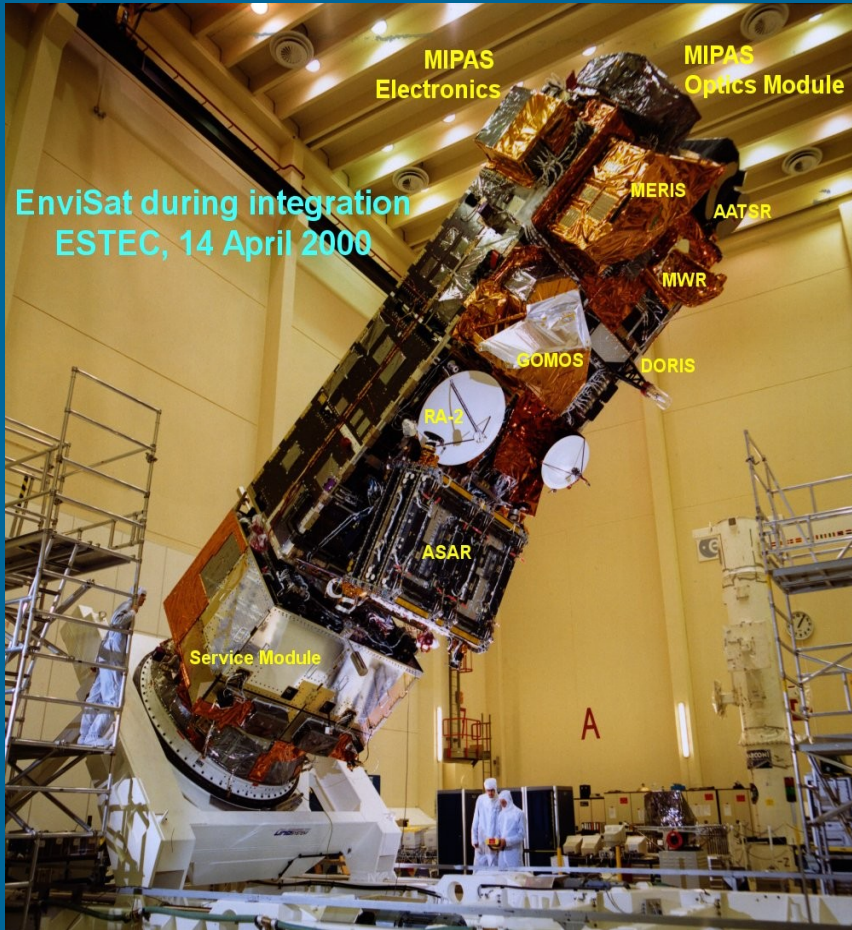
- **MIPAS-
ENVISAT
observations**
- **CTM-modelling**



MANOXUVA

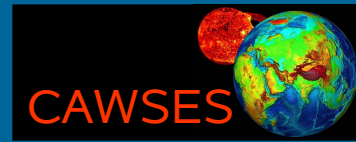
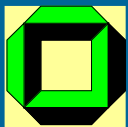


MIPAS instrument



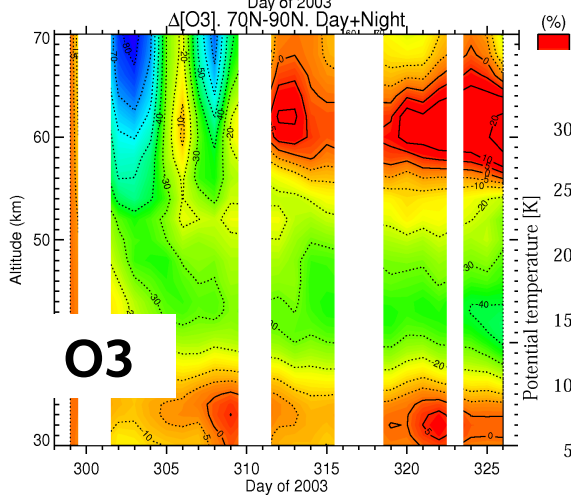
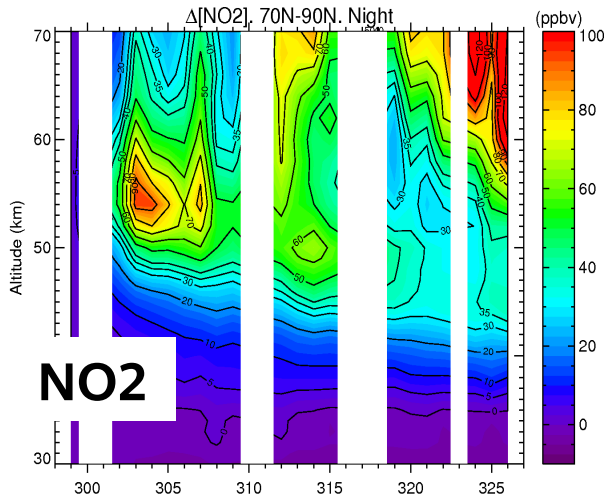
EnviSat during integration
ESTEC, 14 April 2000

- Michelson interferometer launched on 1 March 2002 on ENVISAT (sun-synchronous polar orbit, 98.5°).
- Spectra in 4.15-14.6 μm (685- 2410 cm^{-1}) at 0.035 cm^{-1} in 4.5 s.
- Sensitivity: 4-30 $\text{nW}/(\text{cm}^2 \text{ sr cm}^{-1})$.
- Radiometric accuracy: 1-3 % (4-15 μm).
- Limb, altitude coverage: 6-68 km at 3 km steps in 85 s. (6-170 km)
- Global latitude (Day/Night) coverage
- 72 profiles/orbit; ~1000 profiles/day
- Continuous measurements Sep 2002-Mar 2004.
- Campaigns mode in 2005 (RR) (30-40%)
- Nominal obs. 2006 on ... (40-80%)

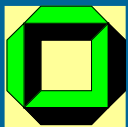
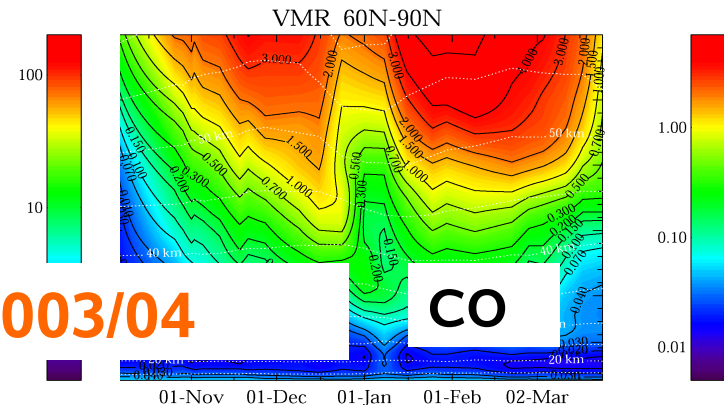
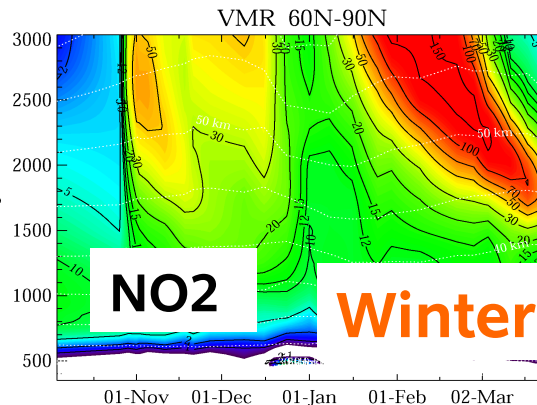


MIPAS sees particle effects in the strato-/mesosphere

SPE Oct/Nov 2003

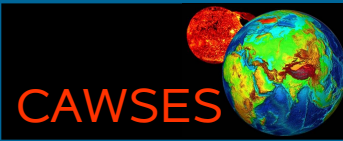


- MIPAS observes effects related to solar proton events (SPEs) (in-situ particle effects) and NO_x intrusions transported downward during polar winter
- Good spatial and time coverage with **standard product of ESA**
- More species and higher accuracy with **IMK/IAA retrieval**



IMK

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Description of the KASIMA model

- Euler and spectral, diagnostic and prognostic
- vertical range 7 - 120 km, vertical resolution variable with height
- horizontal resolution T21 ($5.6^\circ \times 5.6^\circ$)
- ECMWF analyses in combination with trend of prognostic part (T, Div, Vort nudged to analyses; below 17.5 km CTM); operational analyses after Sep 2003, ERA-40 before
- prognostic part uses physical parameterisation of net heating rates and breaking gravity waves
- Chemistry (108 reactions, 38 substances, heterogeneous reactions) up to 90 km
- On-line fast-j2 (vectorized) + Lyman- α
- Mixing ratios of source gases according WMO

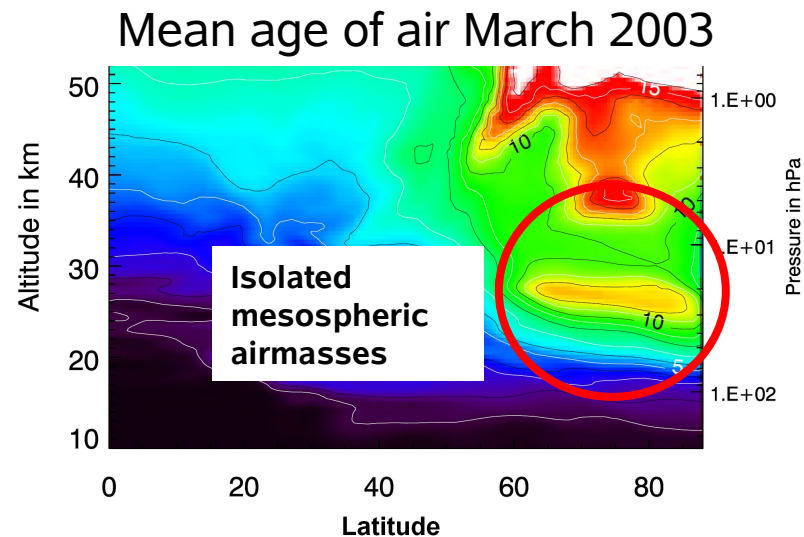
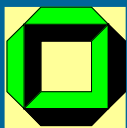


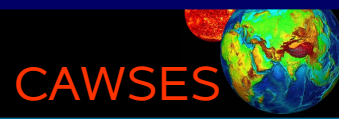
Figure 4: Meridional plot for $l = 20$, day 64

Applications

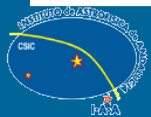
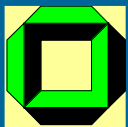
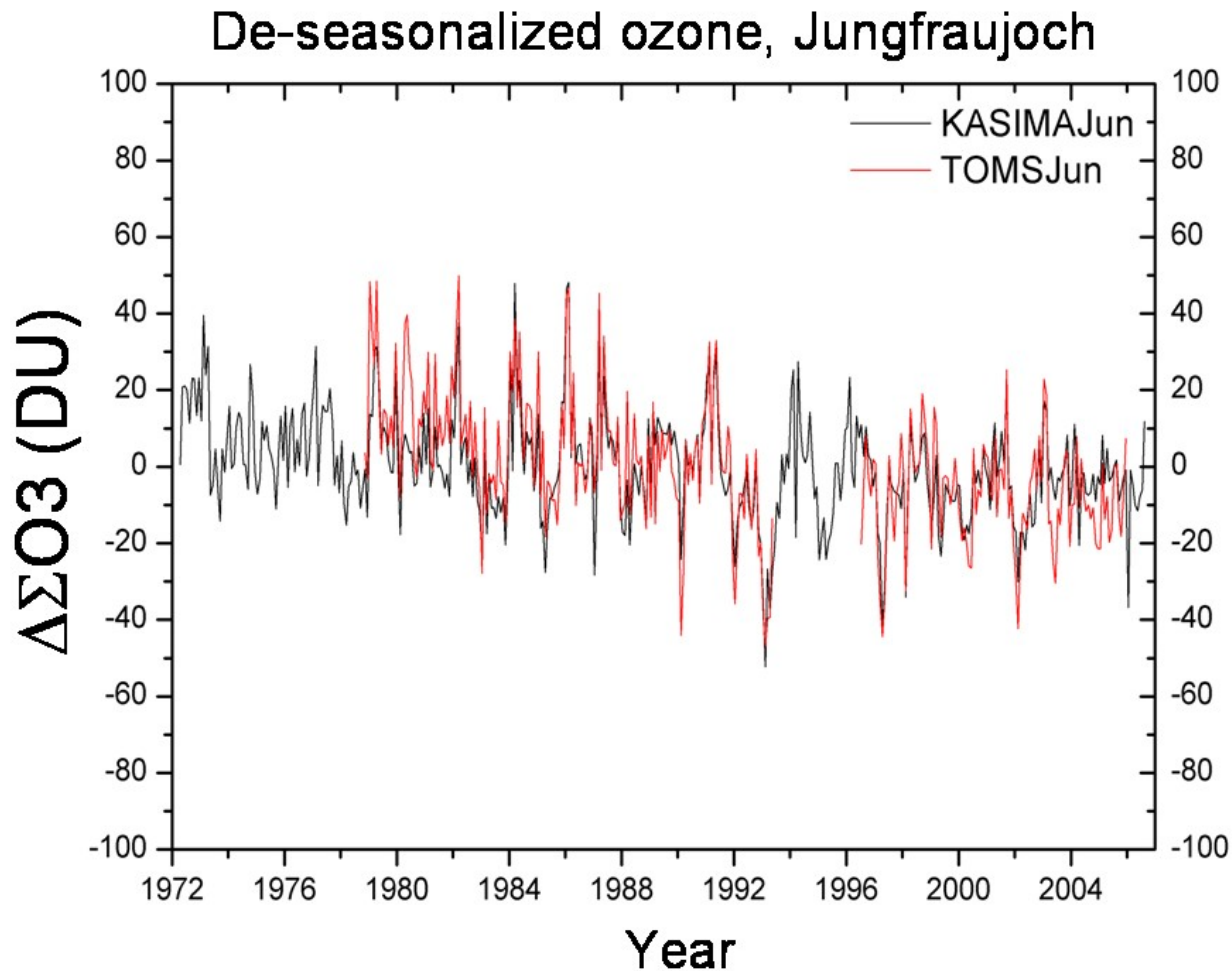
- Seasonal studies of transport and comparison with observations
- Polar chemistry,
- NO_y budget in the lower stratosphere
- Multi-annual integrations with long-lived tracers
- Studies of SF₆ with mesospheric chemistry



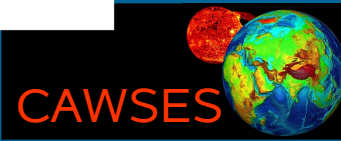
MANOXUVA



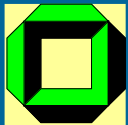
Performance of the model: long term simulations of ozone



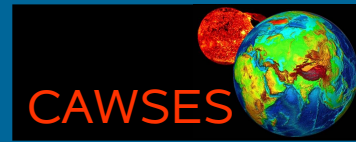
MANOXUVA



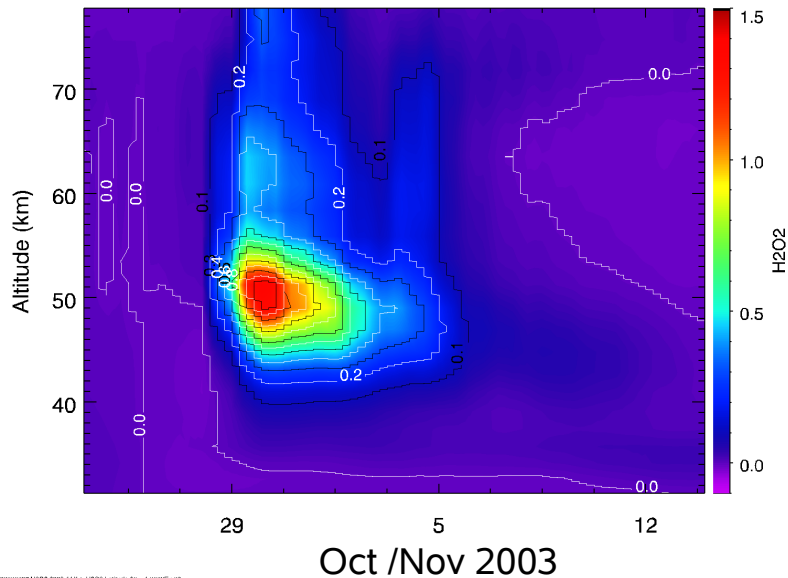
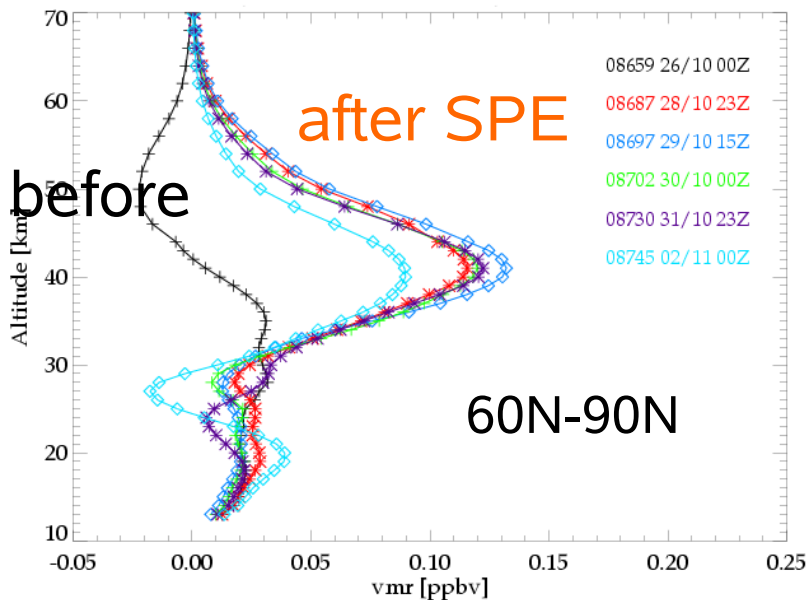
Solar Proton Event(s)



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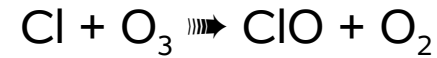
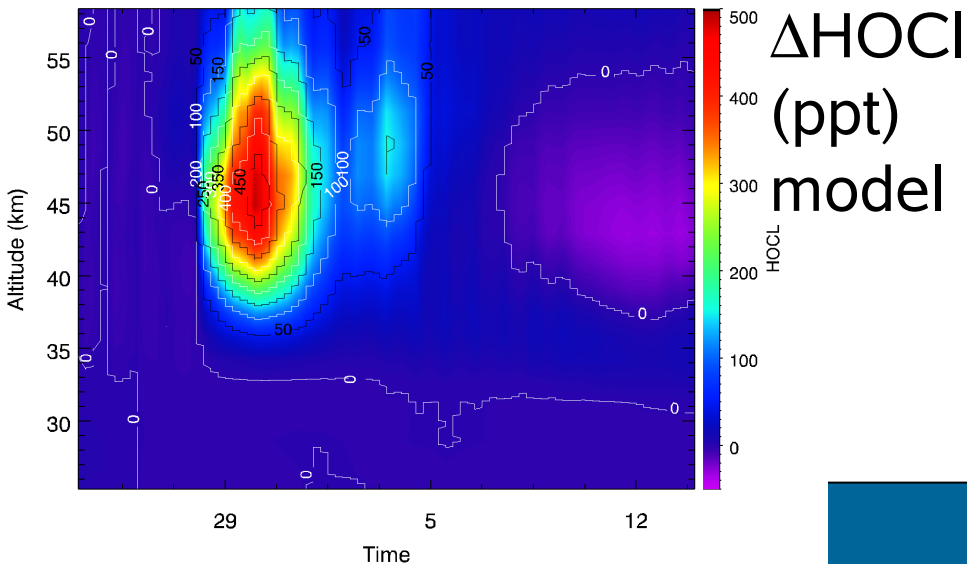
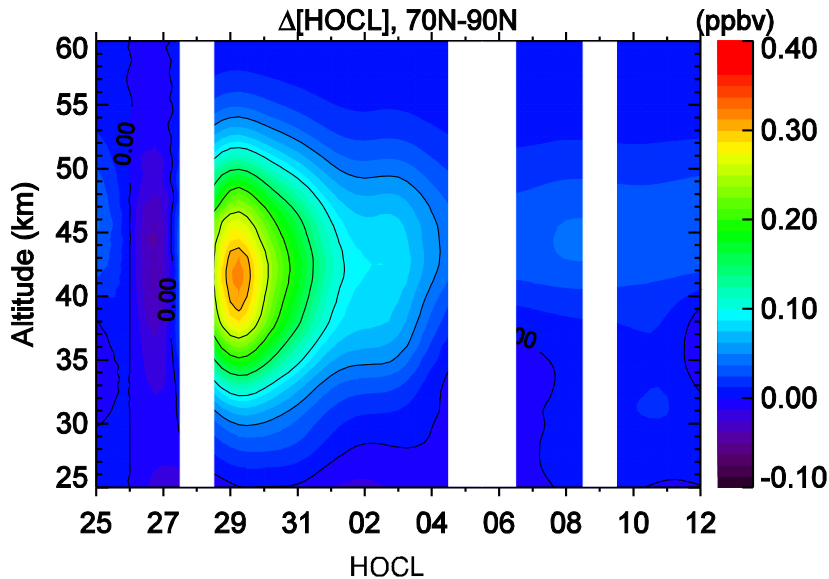


HOx related chemistry: observation of H₂O₂



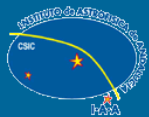
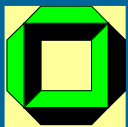
- Retrieval of H₂O₂ for better estimation of HOx related chemistry effects (preliminary results)
HOx cycle. Photochemical equilibrium with OH:
$$\text{OH} + \text{OH} + \text{M} \rightarrow \text{H}_2\text{O}_2 + \text{M}$$
$$\text{H}_2\text{O}_2 + h\nu (\text{longwave}) \rightarrow 2 \text{OH}$$
- Comparison with model: production of HOx via ionisation rates calculated on basis of GOES extrapolated flux, Bethe-Bloch and standard efficiencies for HOx/ionis.

HOCl: comparison with MIPAS

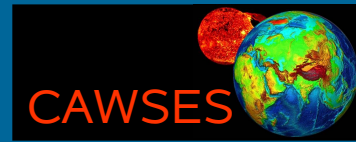


- possible estimation of HOx
- good agreement with HOCl
- problems with HNO₃

NOx intrusions and long term effects

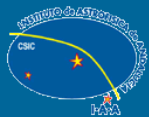
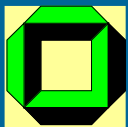
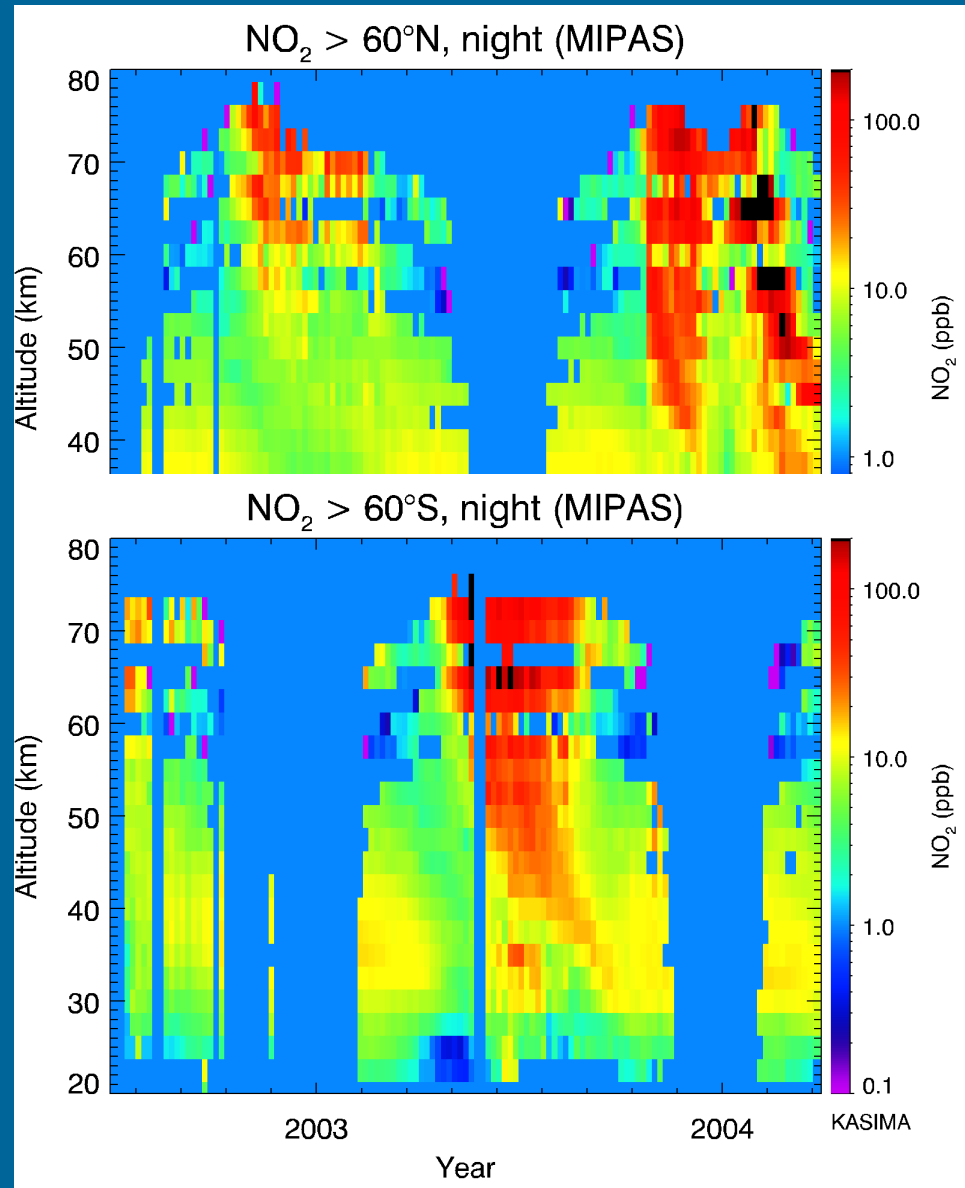


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NO_x intrusions and long term effects

- Two types of NO₂ enhancements are seen in data:
 - thermospheric origin with enhancements in the mesosphere only in the arctic night
 - global enhancements in the mesosphere and stratosphere as a aftermath of solar proton events (SPE)
- ESA data in the mesosphere only available during phases of enhanced NO_x
- Use MIPAS NO₂ night to prescribe NO_x
- Only above ~55 km as a disturbance
- Compare with reference run
- Fixed heating rates (chemistry effects)



Using data in the model

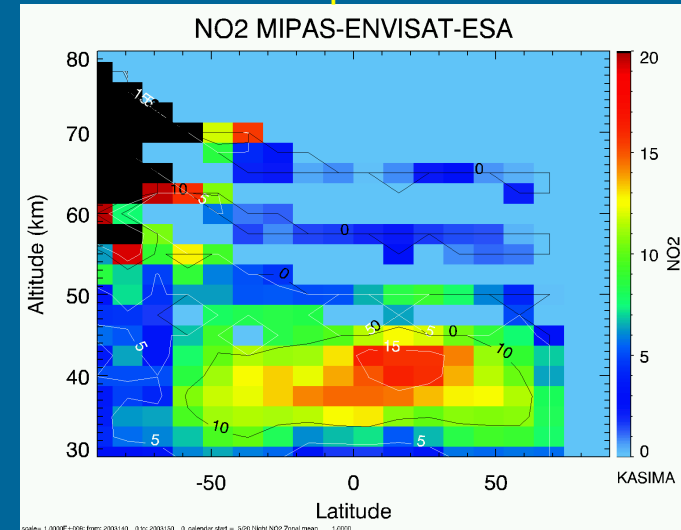
- We use NO₂ observation during night as a proxy for NO by sampling several days
- Because of the low frequency of MIPAS observations, as a compromise 5 days averages are used; not zonal means
- Data are interpolated to a grid
- Interpolation is done in a form of modified inverse distance weighting:

$g_j = \sum w_{ij} y_i$ g : gridded data at gridpoint x_j ; y_i observation at x_i

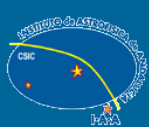
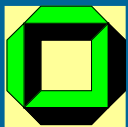
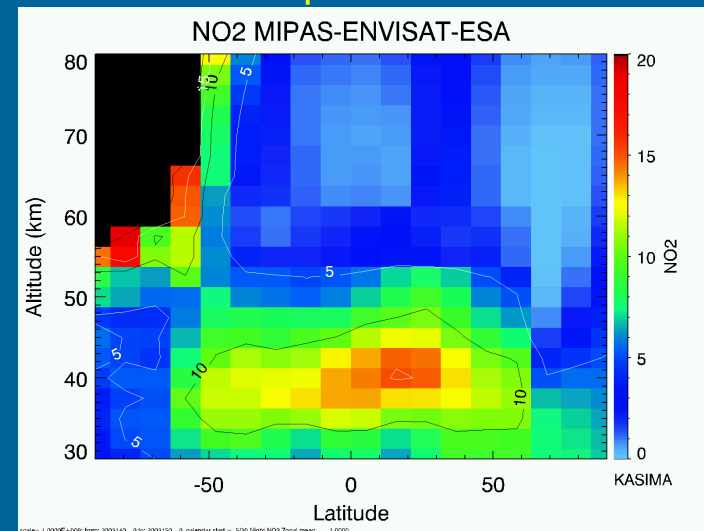
$$w_{ij} \sim 1 / 1 + (d_{ij} / l)^p$$

$d_{ij} = d(x_j, x_i)$: distance on sphere of gridpoint and observation, aspect ratio vertical to horizontal about 100, p some exponent (here $p=8$), l some characteristic length (about 200 km)
parameters used result from tests and somewhat subjective

Sampled data

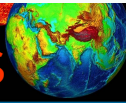


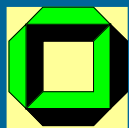
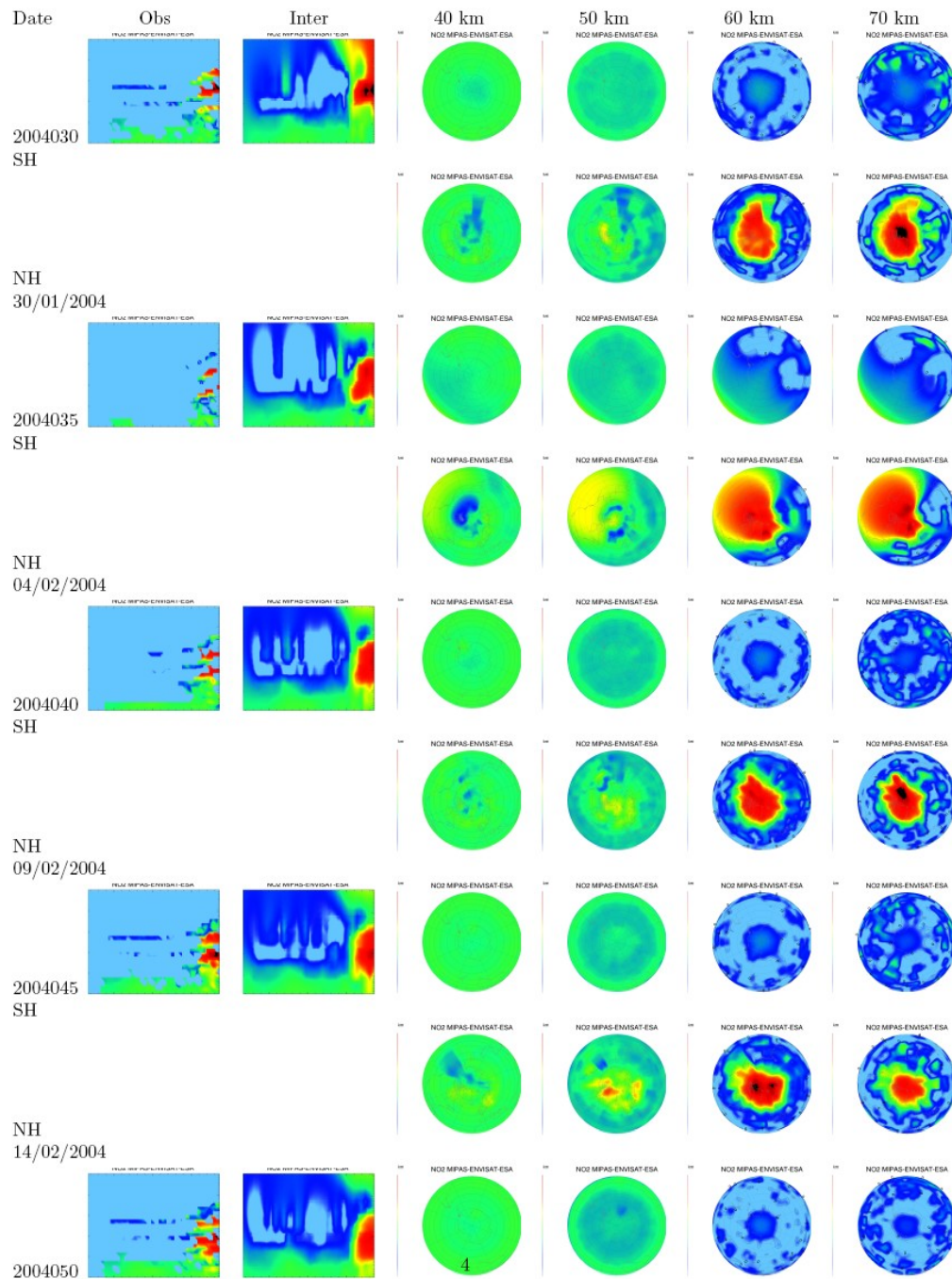
Interpolated data



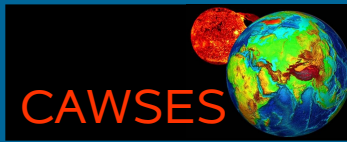
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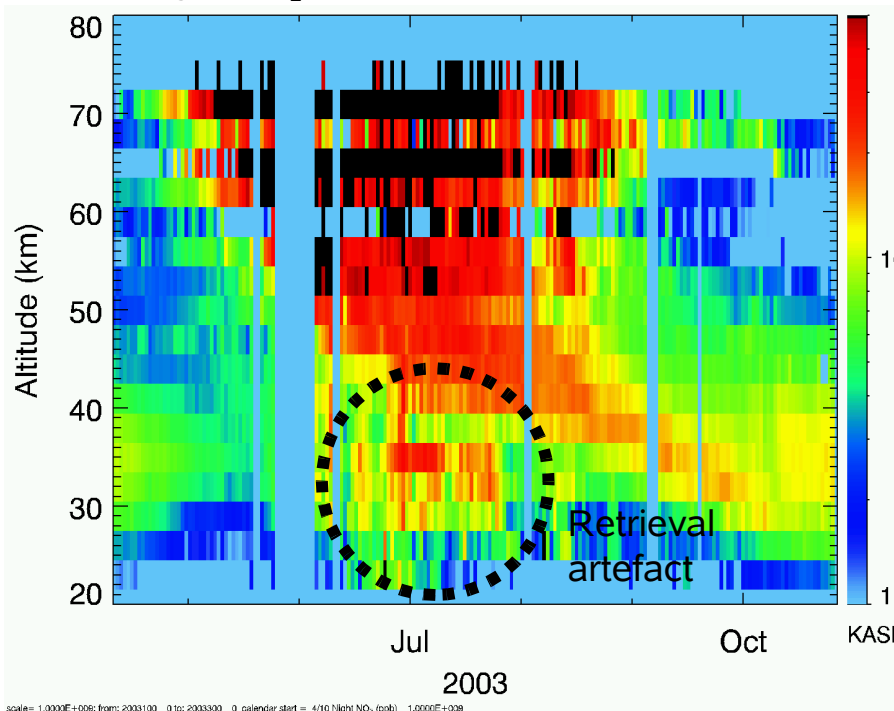


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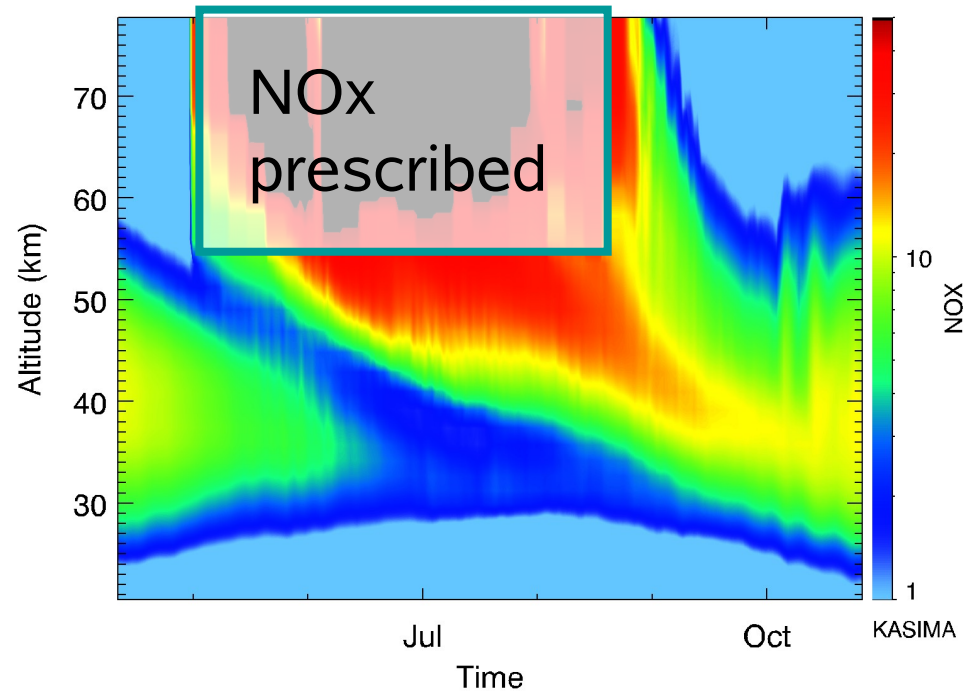


NO_x in the model and in observations

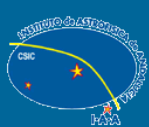
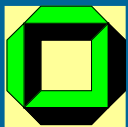
Observations
night NO₂ MIPAS-ENVISAT > 60°S



Model NO_x > 60°S

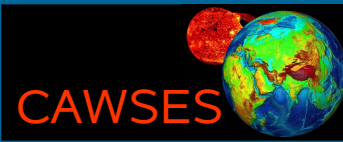


Antarctic winter 2003



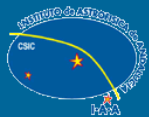
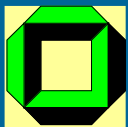
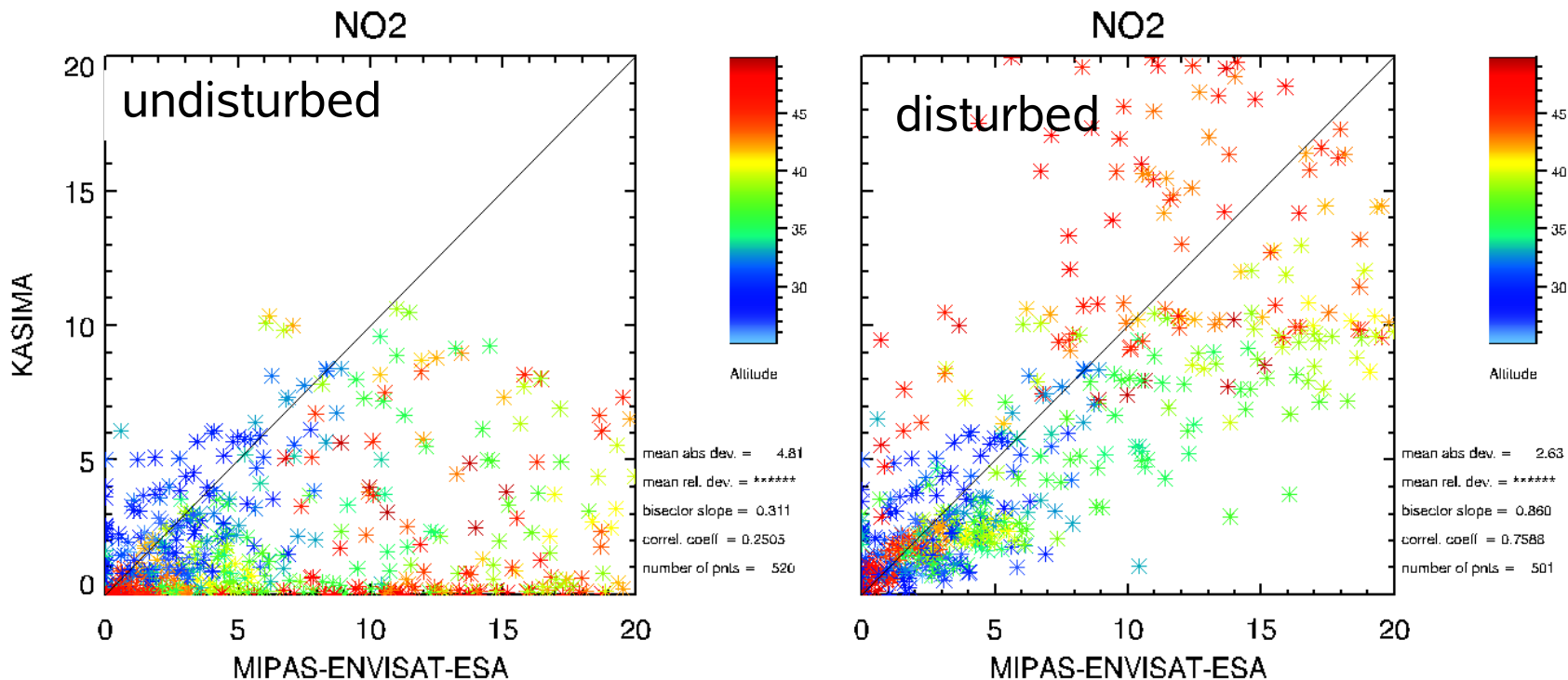
FIMK

MANOXUVA

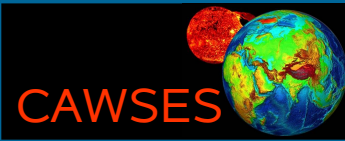


Correlations Observations vs. Model

Effect of NO_x intrusion on correlation for stratosphere 50 days after initialization 2003 DOY 180. Only coincident data are used ($|\Delta t| < 30$ min)

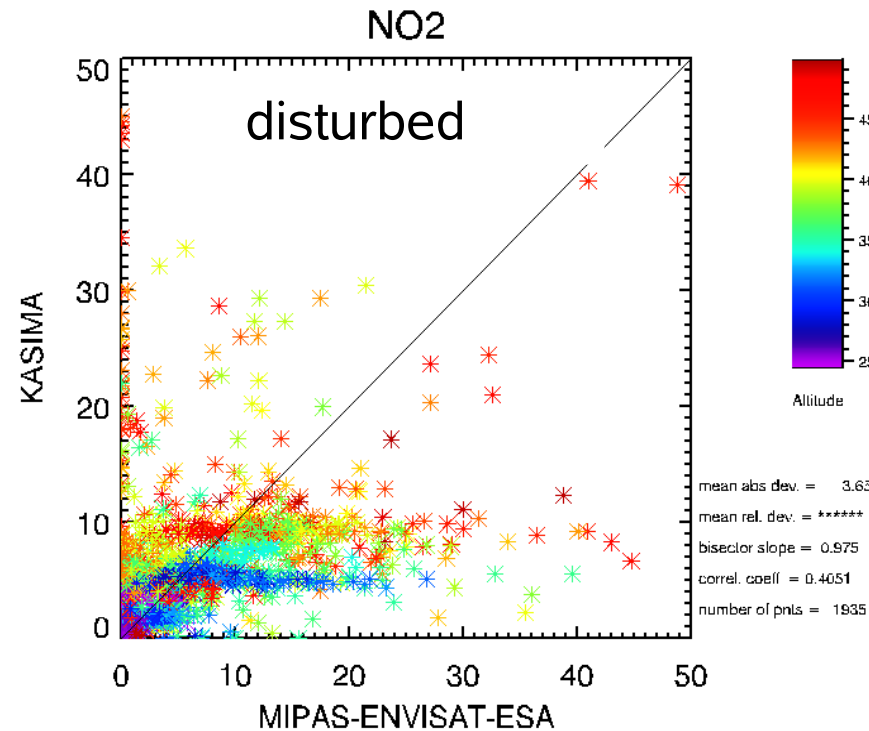
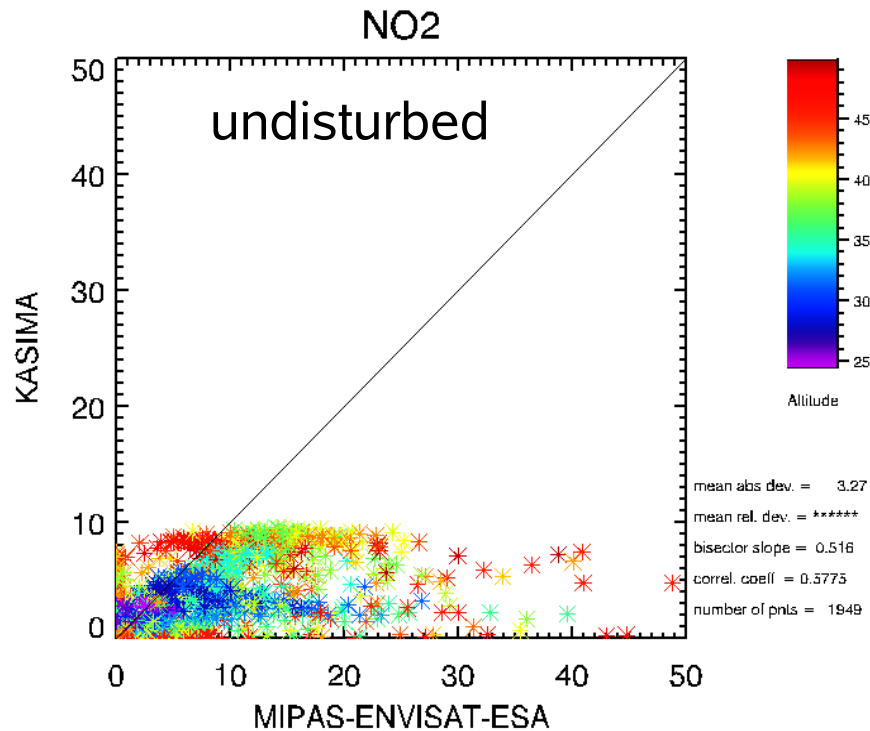


MANOXUVA

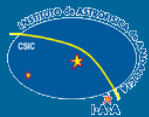
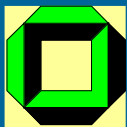


Correlations Observations vs. Model

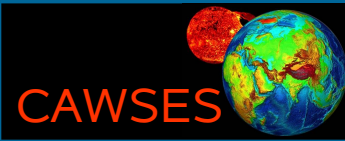
Effect of NO_x intrusion on correlation for stratosphere winter 2003/2004
DOY 30 -80. Only coincident data are used ($|\Delta t| < 30$ min)



Problems with dynamically very active winter 2003/2004

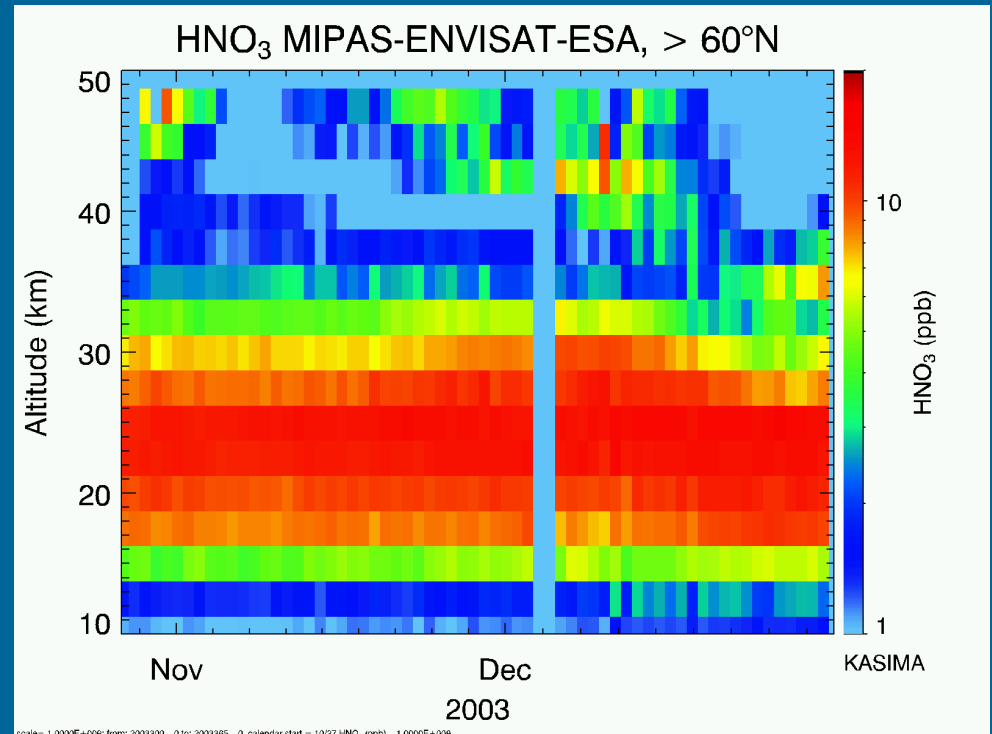


MANOXUVA

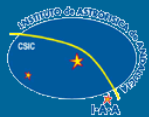
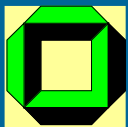


HNO_3 buildup through ion cluster chemistry

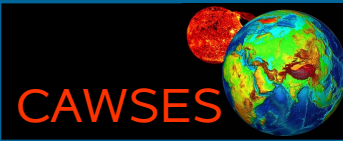
- MIPAS data show strong HNO_3 enhancement in Antarctic and Arctic winters (see analysis of Orsolini et al. 2004 for winter 2003/2004)
- Has been also observed by CLAES
- Even in zonal mean present till January
- Missing in model
- Ion mechanism (Böhringer et al)?



- Simplified approach of de Zafra et al.
- Ion reaction $\text{N}_2\text{O}_5 + \text{H}^+(\text{H}_2\text{O})_n \rightarrow \text{H}^+(\text{H}_2\text{O})_{n-1} + 2\text{HNO}_3$ (only for $z > 40\text{km}$)

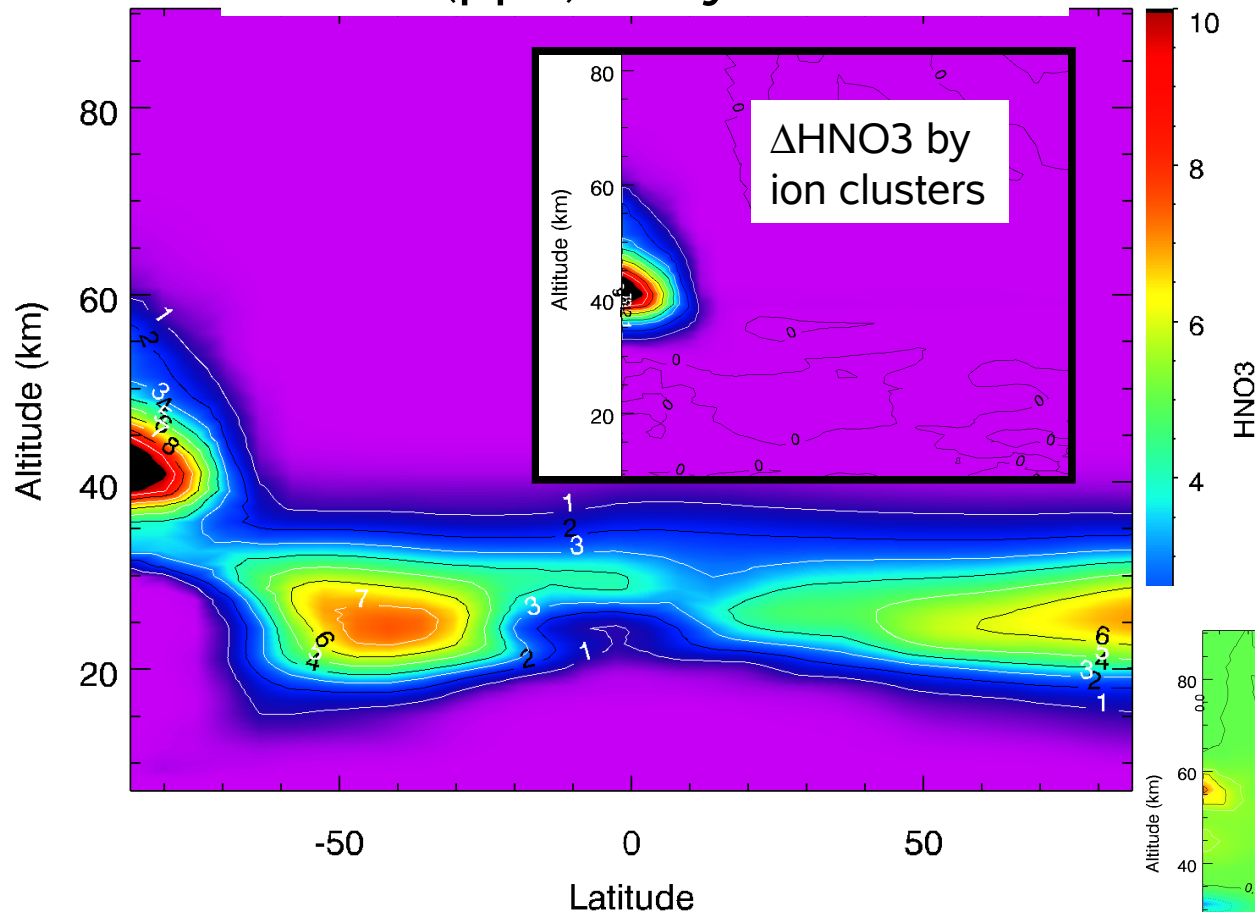


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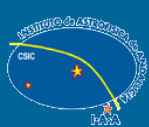
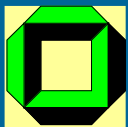
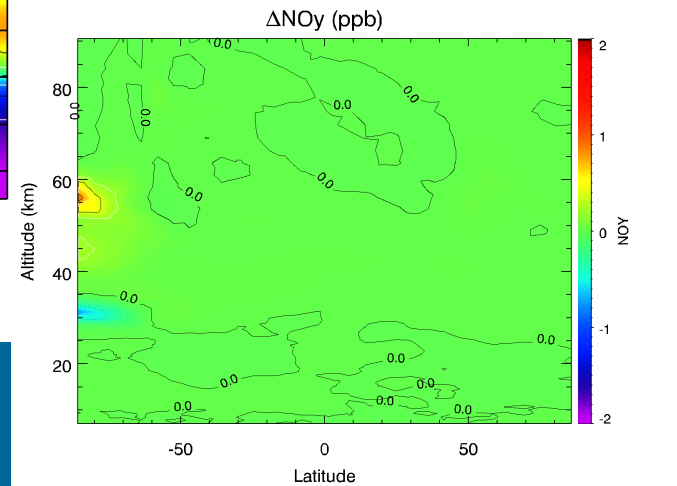


Effects of water ion clusters on NOy

HNO₃ (ppb), Day 216 2003

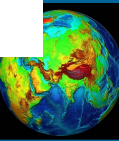


- Ion clusters can provide HNO₃ secondary maximum
- Small effects on total NO_y and ozone

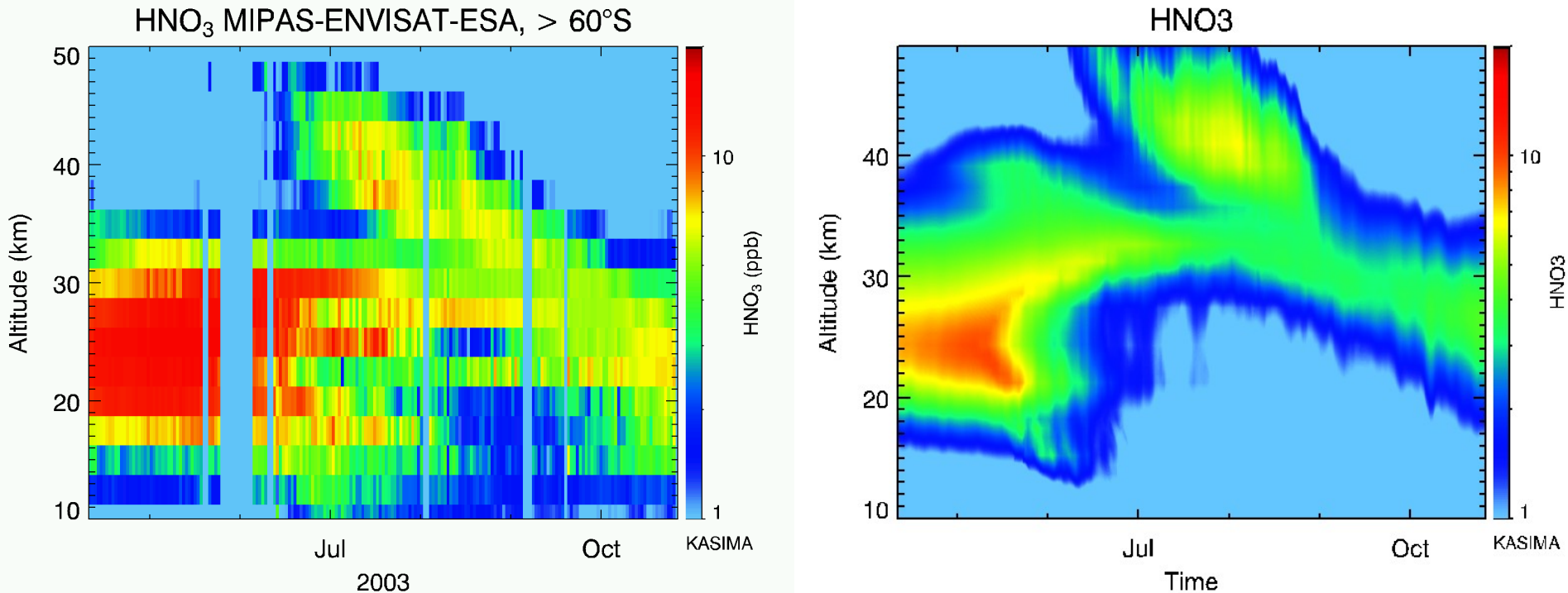


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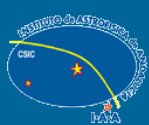
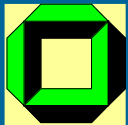
CAWSES



HNO₃ in the model and in observations

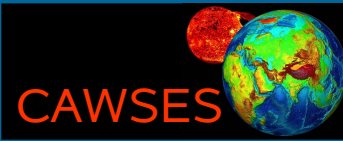


Antarctic winter 2003 (see also analysis of Funke et al. and Stiller et al.)

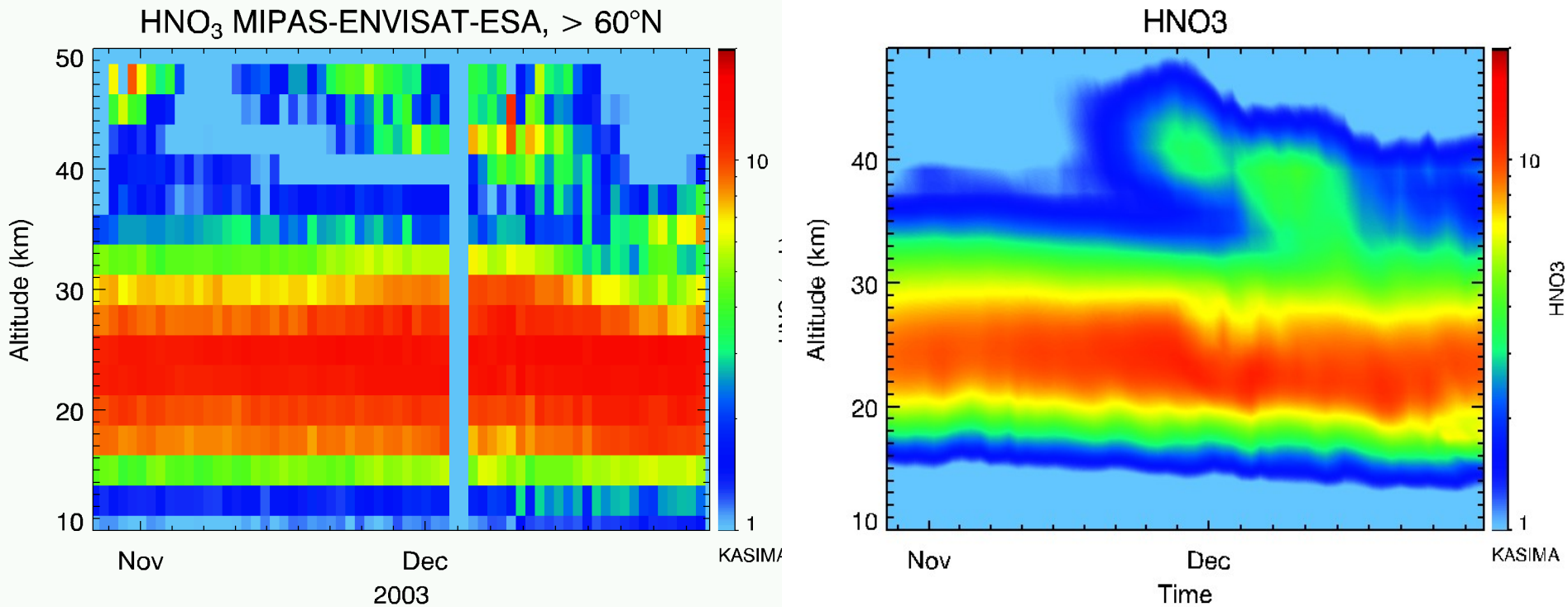


FIMK

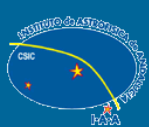
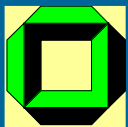
MANOXUVA



HNO₃ in the model and in observations

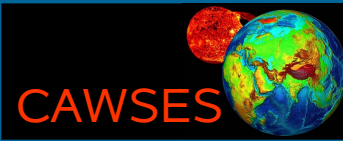


Early Arctic winter 2003/2004

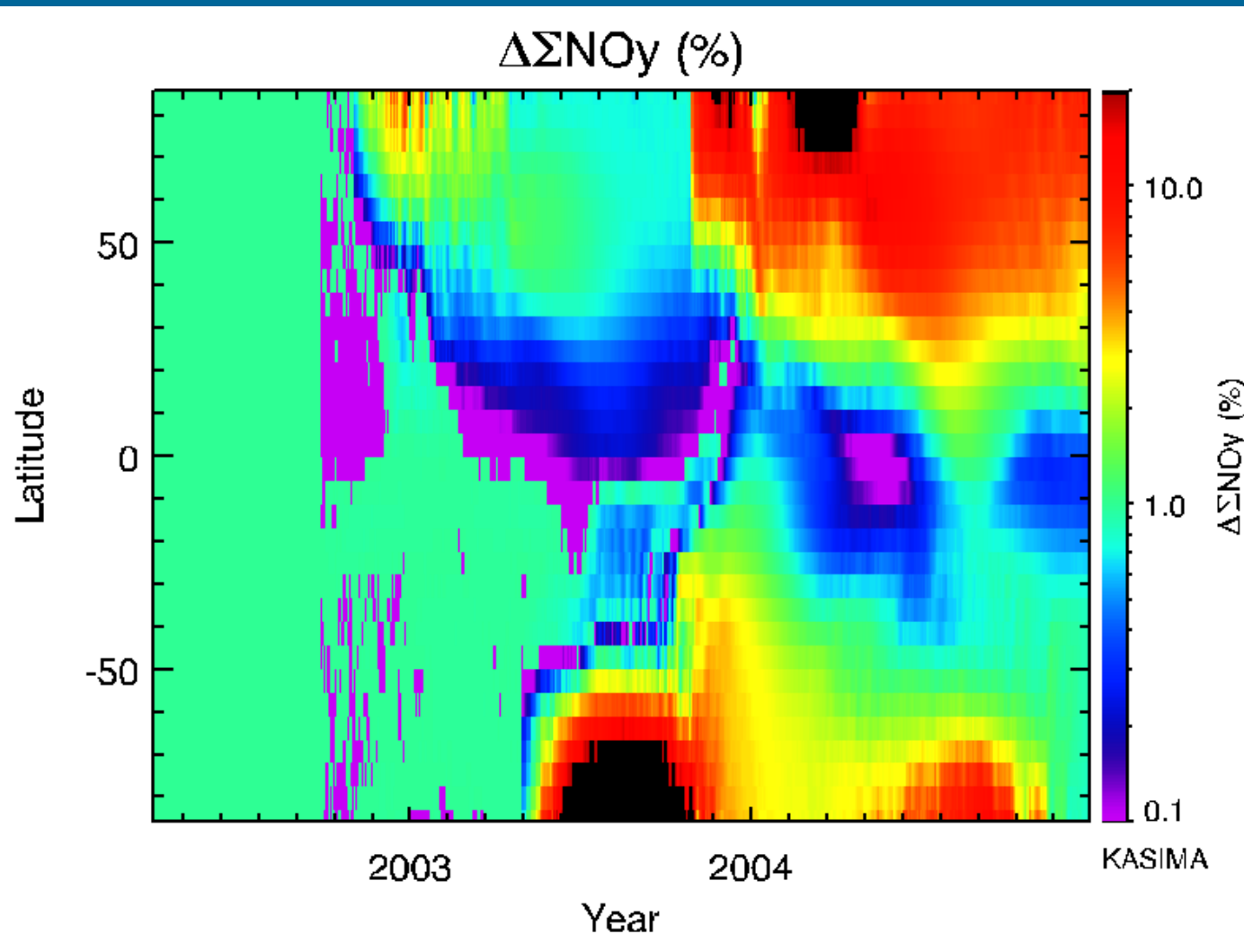


FIMK

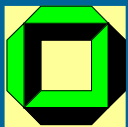
MANOXUVA



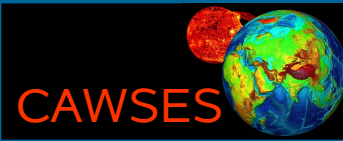
Effect on total NO_y



- NO_y enhancements lasting more than one winter
- Mostly confined to polar latitudes
- Strong in Arctic winter 2003/04

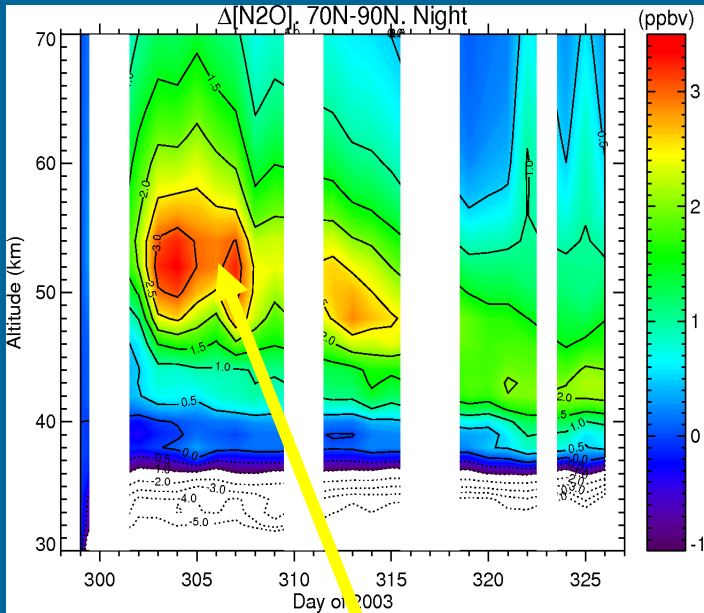


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NOx as a source for N2O

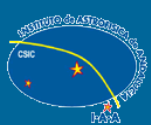
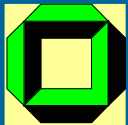
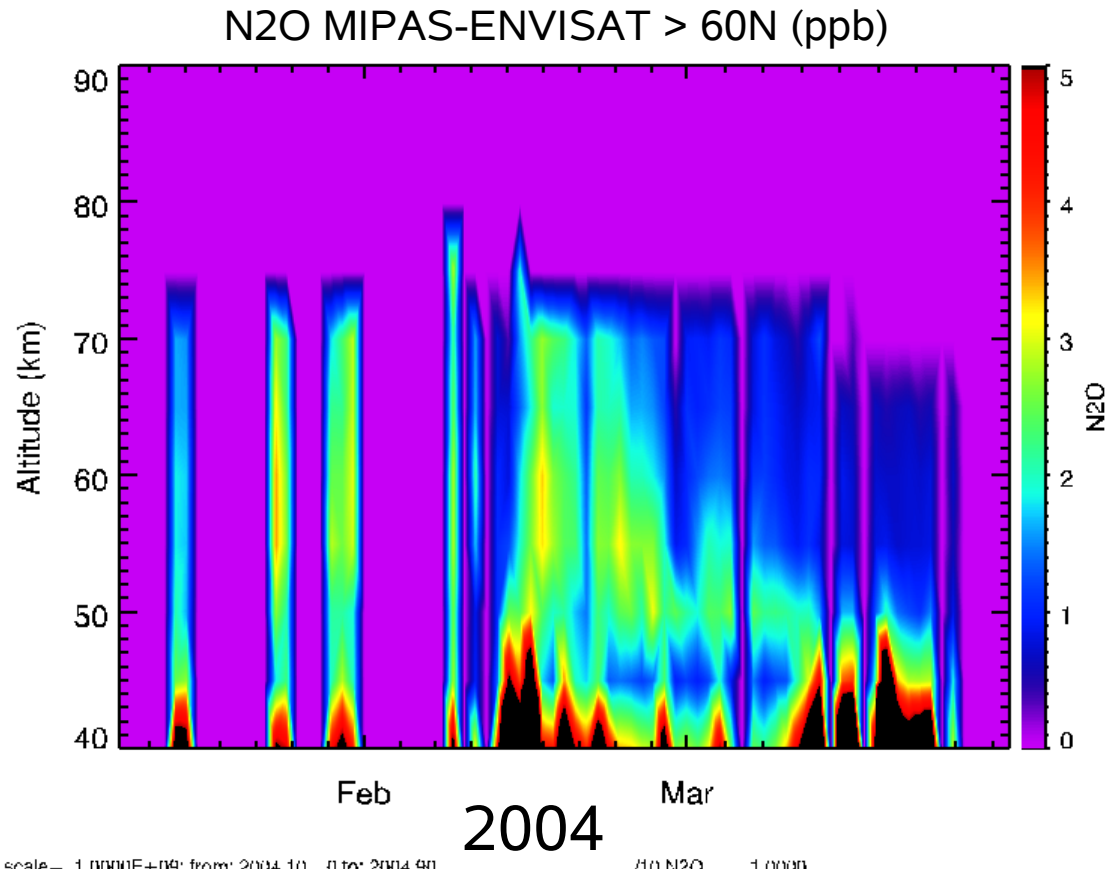
- Quasi In-situ produced N during aurora events?
- Effect for tracer-tracer correlation studies?



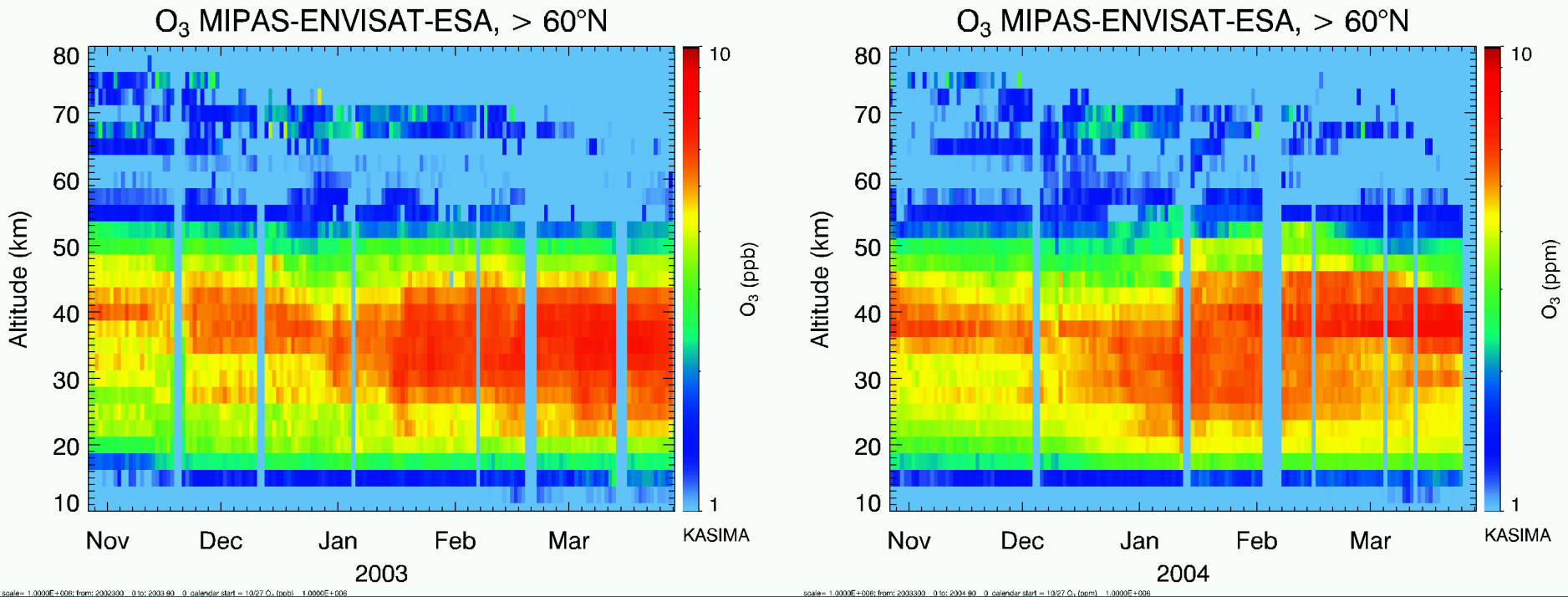
N2O from

$\text{N} + \text{NO}_2 \rightarrow \text{N}_2\text{O} + \text{O}$

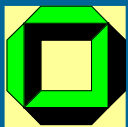
In-situ production of N
during SPE



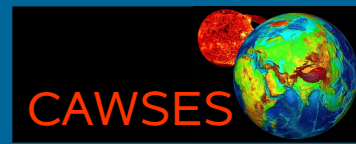
Ozone



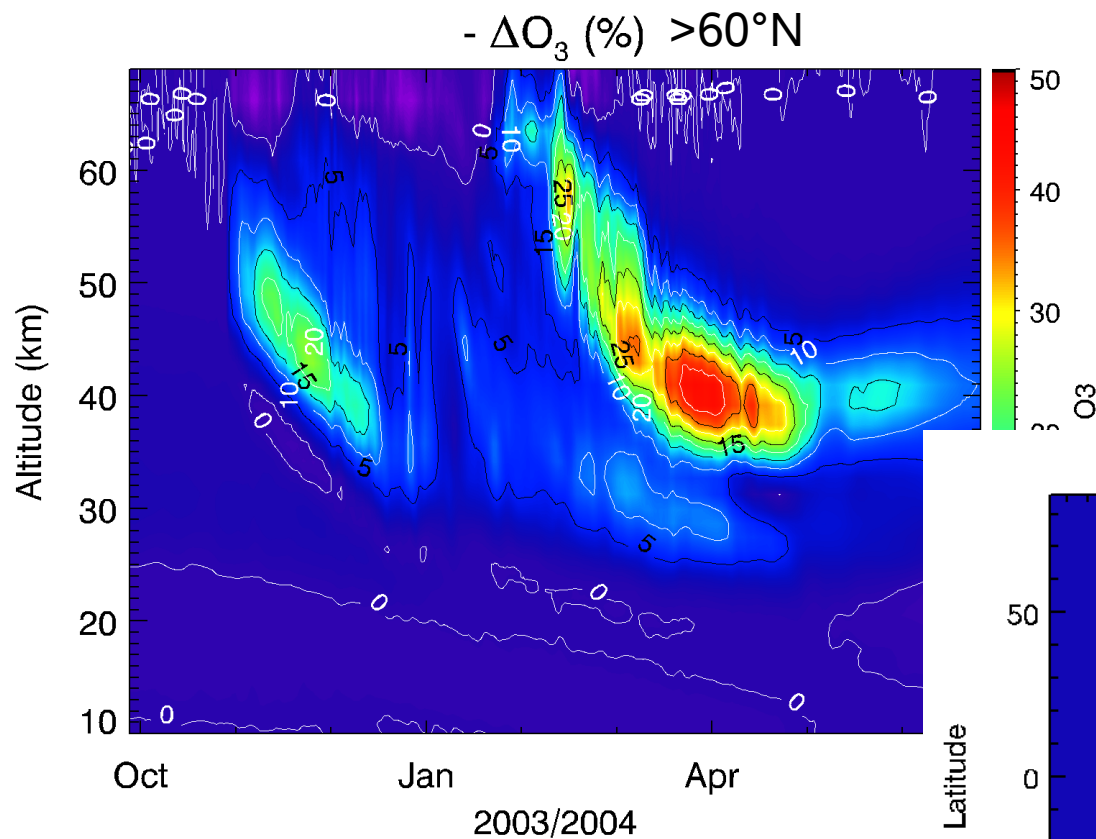
Effect on ozone is not very obvious on the long term



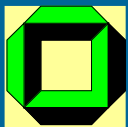
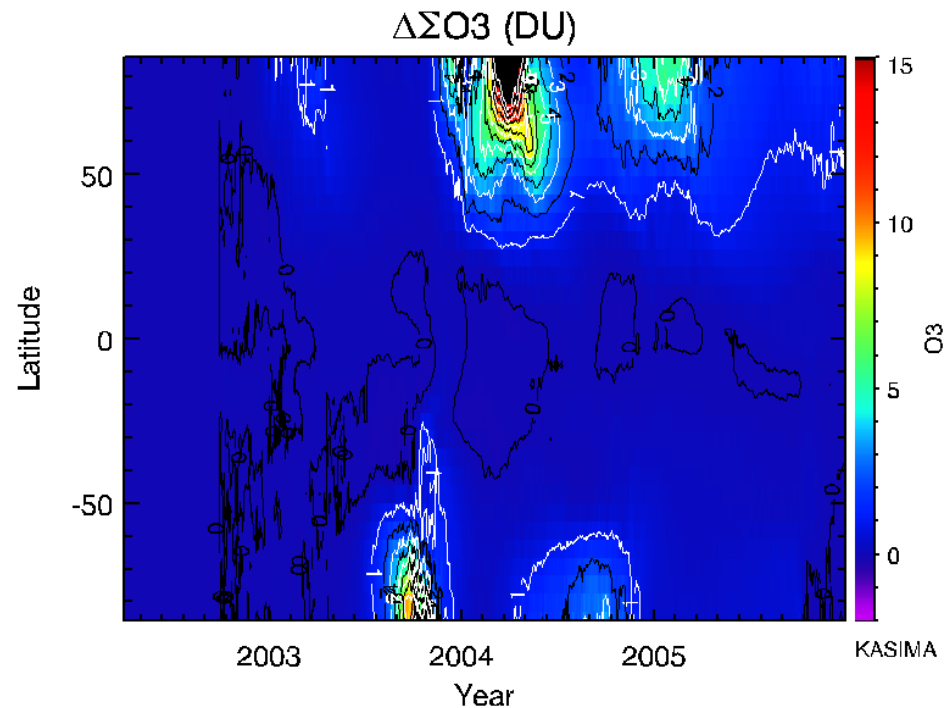
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Additional ozone loss

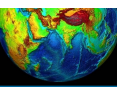


- Use reference run to estimate additional ozone loss as difference between the two runs



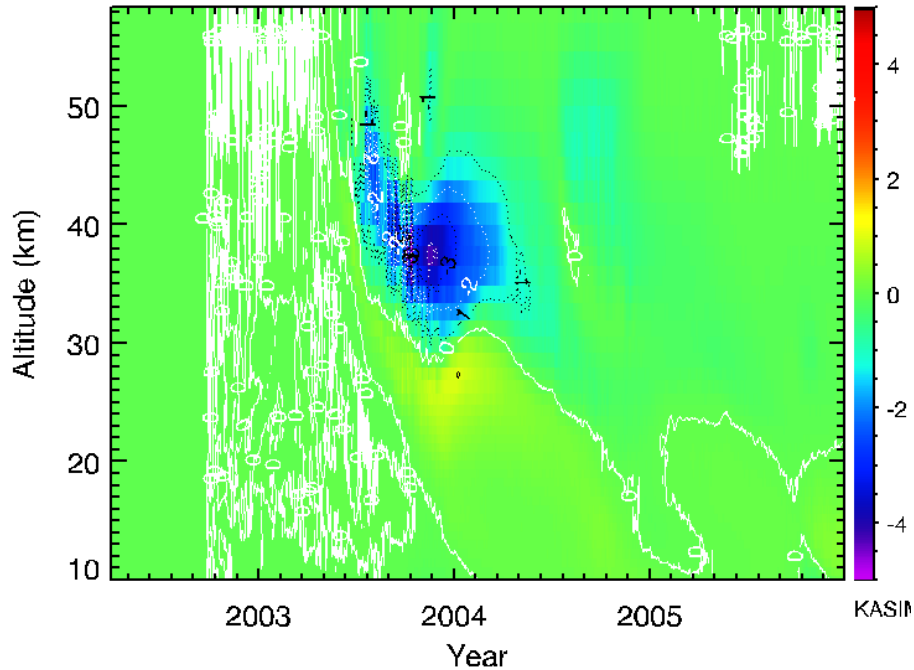
MANOXUVA

CAWSES

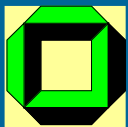
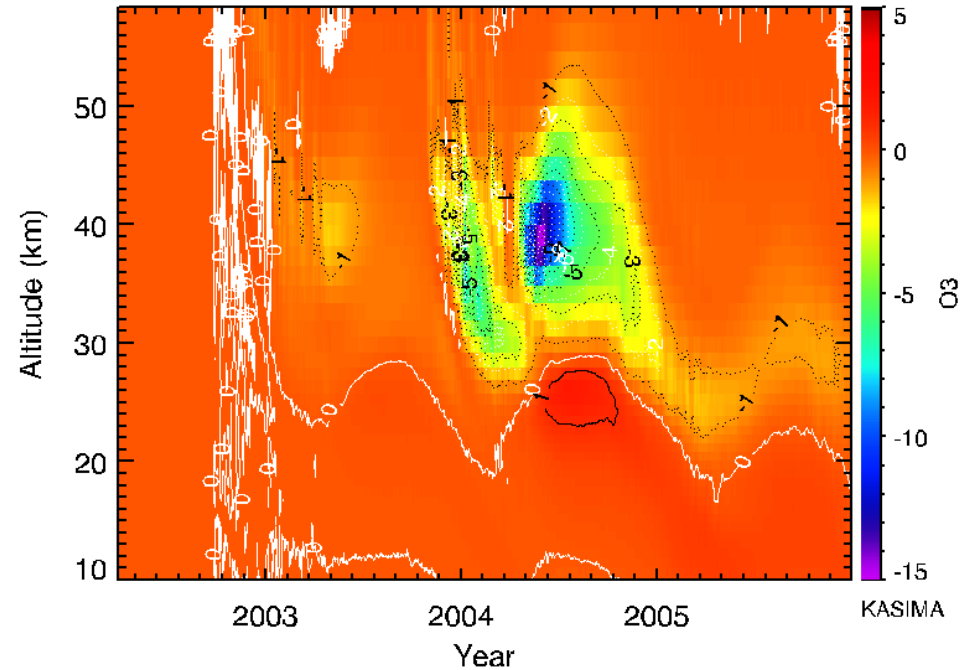


Midlat effects in ozone

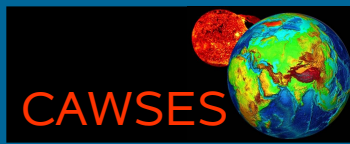
ΔO_3 (%) 45°S



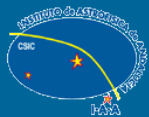
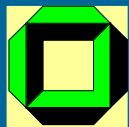
ΔO_3 (%) 45°N



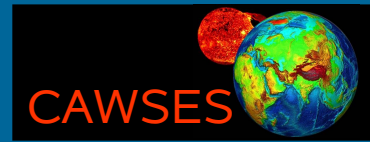
MANOXUVA



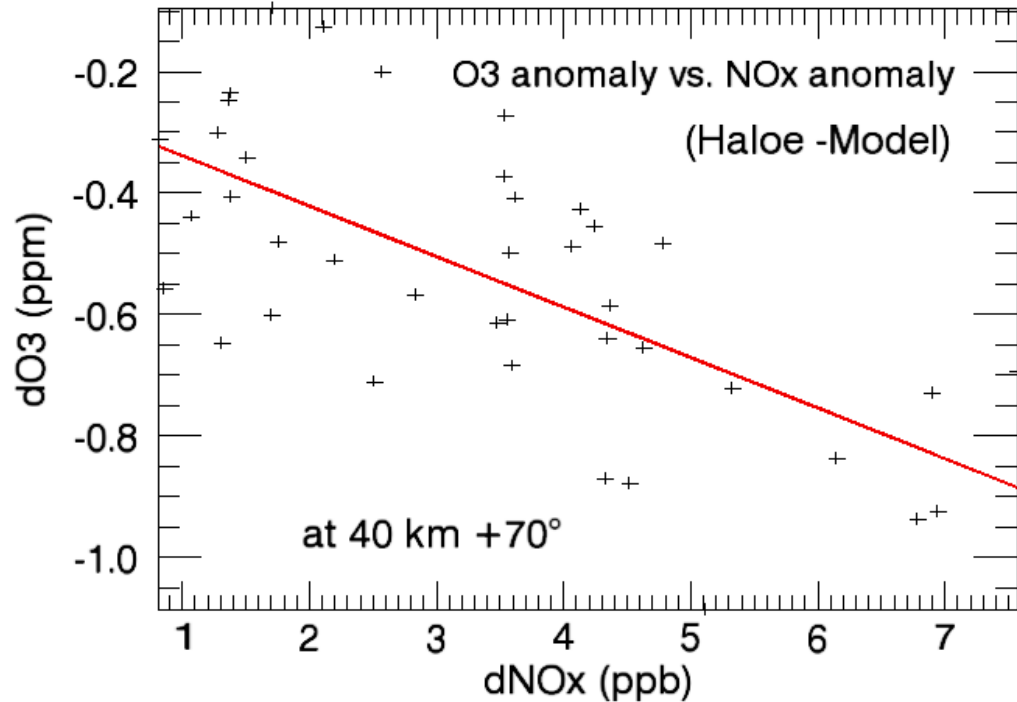
Multidecadal simulations and search for solar signal



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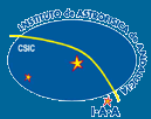
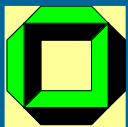
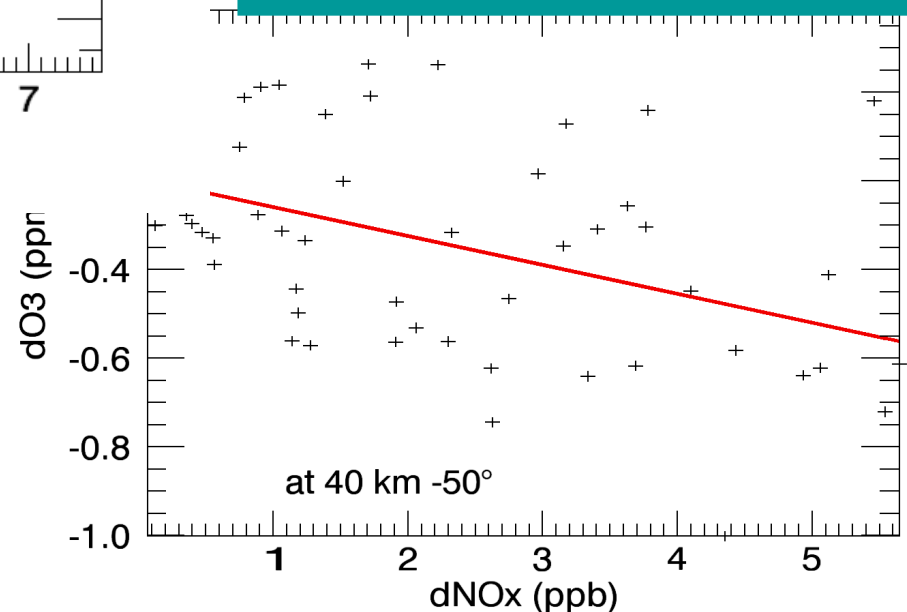


Long-term variation of NOx and effects on O3

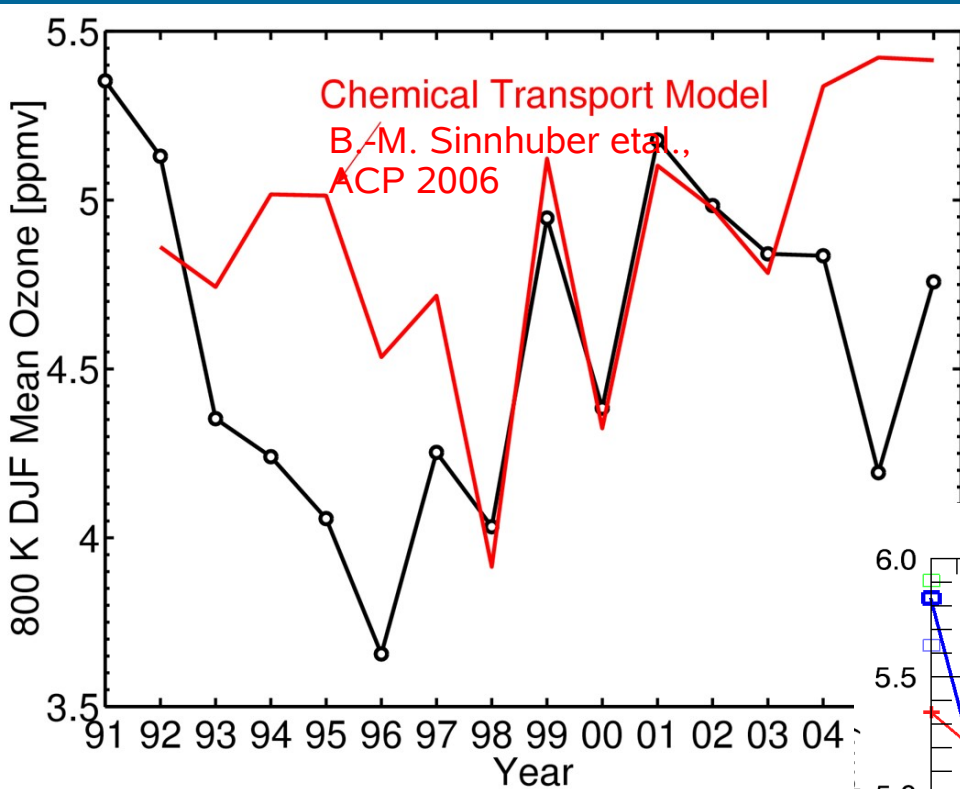


- Use multidecadal run to search for solar related effects
- As no solar signal is included in the run, look for anomalies when comparing with data
- Use HALOE O3 and NOx observations in altitude and latitude bands

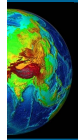
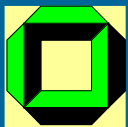
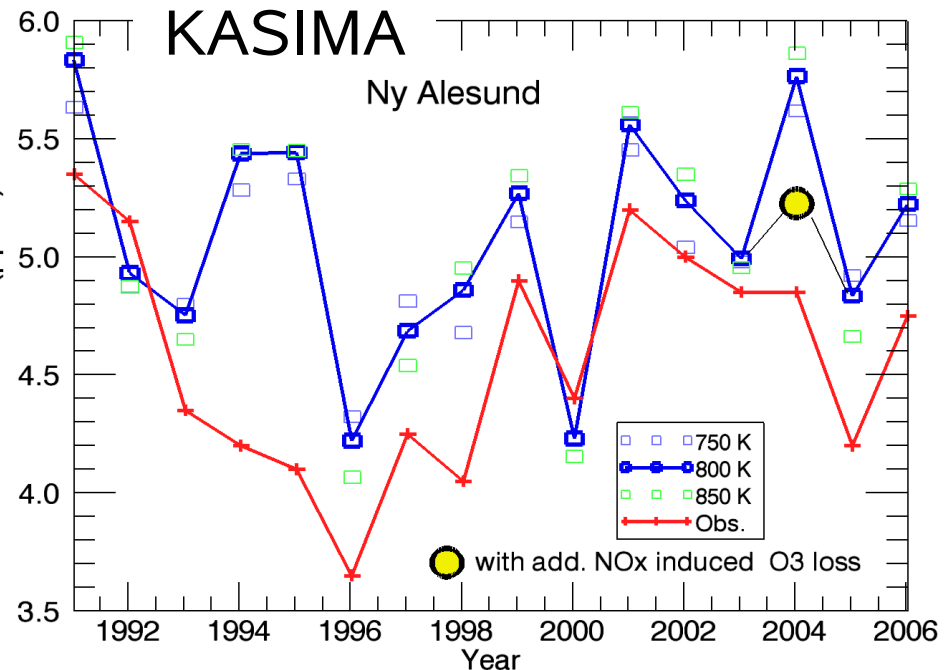
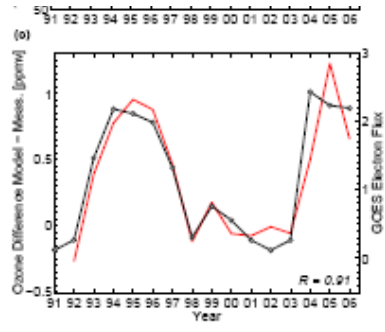
NOx related O3 signal!



Another mystery

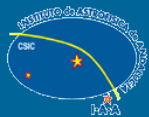
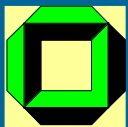
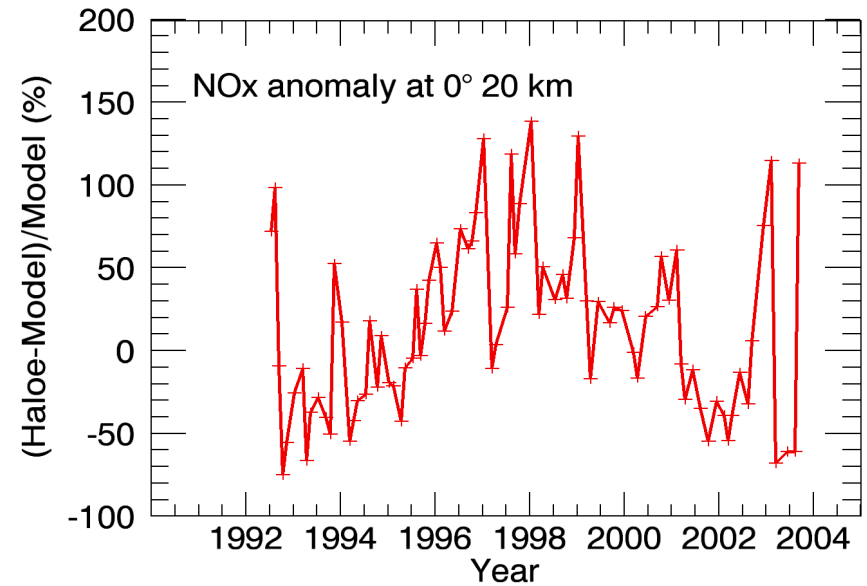
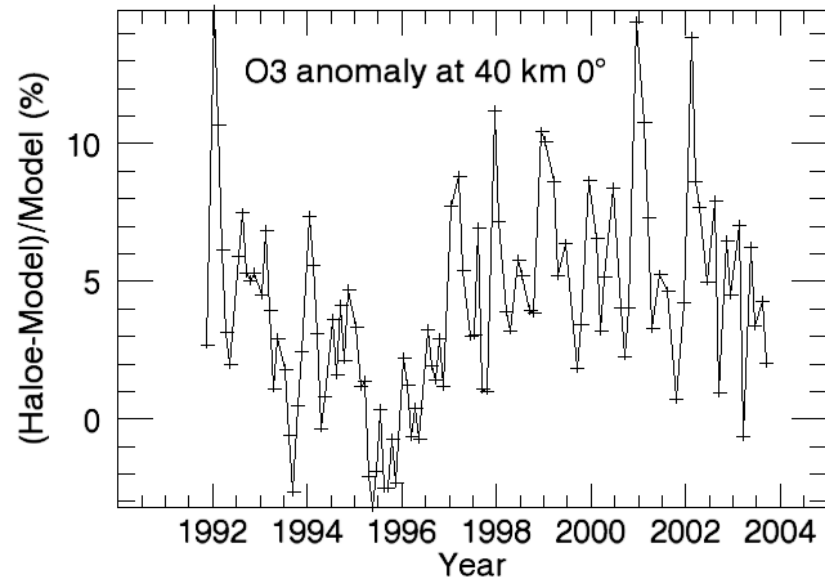


- Sinnhuber et al found anomaly in Spitzbergen O3 data
- Anomaly may be partly explained by NO_x intrusion
- Correlation with GOES lost

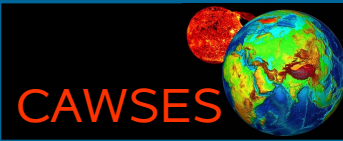


Signatures of UV variability?

HALOE - MODEL

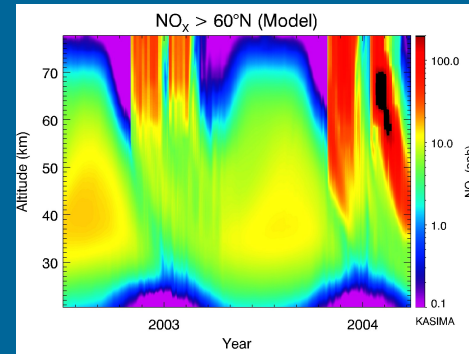


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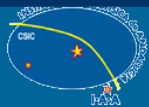
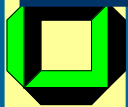
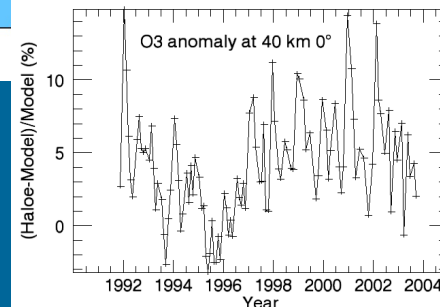
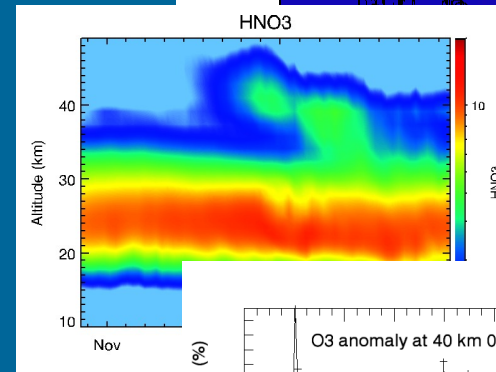
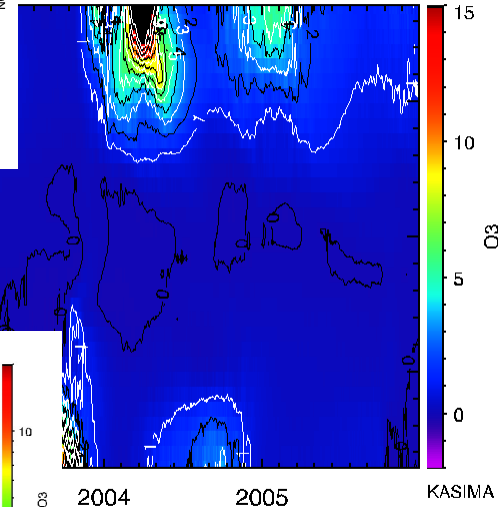


Summary:

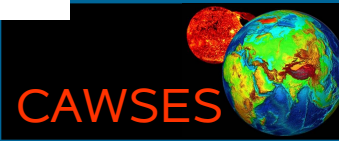
- MIPAS provides an extensive dataset for chemistry studies after SPEs
- MIPAS data can be used to prescribe NO_x in KASIMA's polar meso/stratosphere and to estimate influence of NO_x intrusions on chemistry
- Important effects on ozone restricted to polar latitudes. Trend analysis for high latitudes may be affected!
- Observed HNO₃ enhancements can be qualitatively reproduced when introducing HNO₃ buildup via H₂O cluster ions
- Open questions:
 - Interplay of in-situ production and dynamics in MLT for aurora effects
 - Downward transport during winter
 - Understand buildup of HNO₃ quantitatively
 - The buildup of N₂O
 - Relative contribution of EPP and UV for ΔO₃ on longer timescales



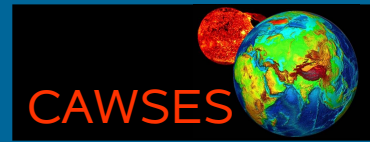
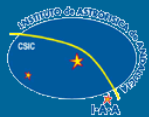
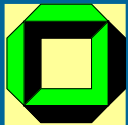
ΔΣO₃ (DU)

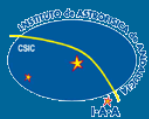
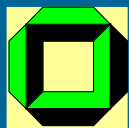


MANOXUVA



Thank you !





MANOXUVA



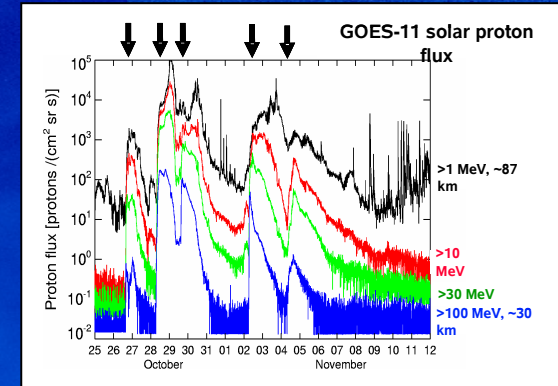
Particle precipitation, ionization and radical production

- Energy loss of energetic particles in the atmosphere by ionization, N_2^+ , and subsequent production of NO^+
- Recombination of N_2^+ or NO^+ or charge exchange reaction yield $N(2D, 4S)$
- $N(2D) + O_2 \rightarrow NO + O$



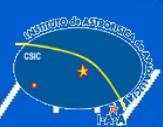
- Net effect is the production of NO_x in the lower thermosphere/mesosphere
- Production of HO_x in the meso-/stratosphere via water cluster ions

fast
slow



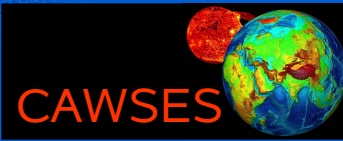
See Brasseur
Solomon 2005

Result: NO_x and HO_x enhancements in the middle atmosphere disturbing regular chemistry



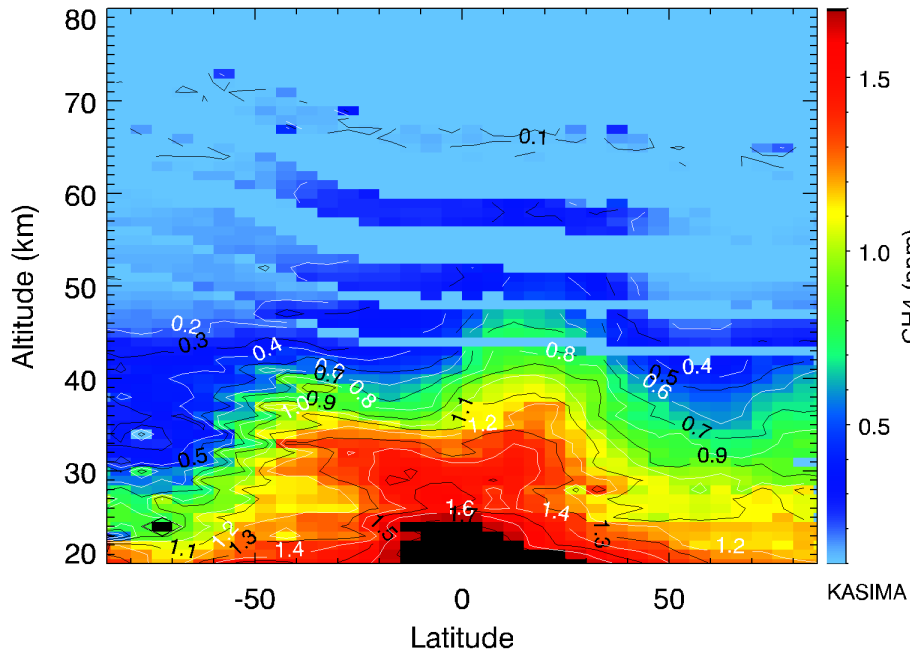
18

MANOXUVA

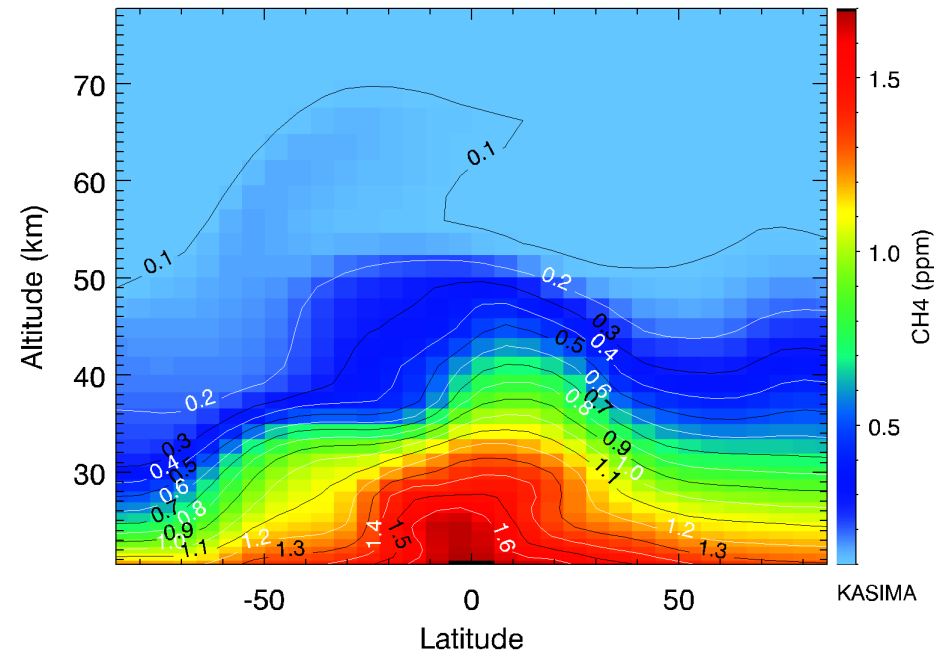


General transport properties of the model

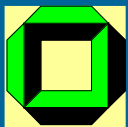
CH4 MIPAS-ENVISAT-ESA



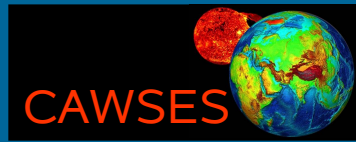
CH4 MIPAS-ENVISAT-ESA



Zonal mean CH4 (ppm) for period 2003 DOY 180-200

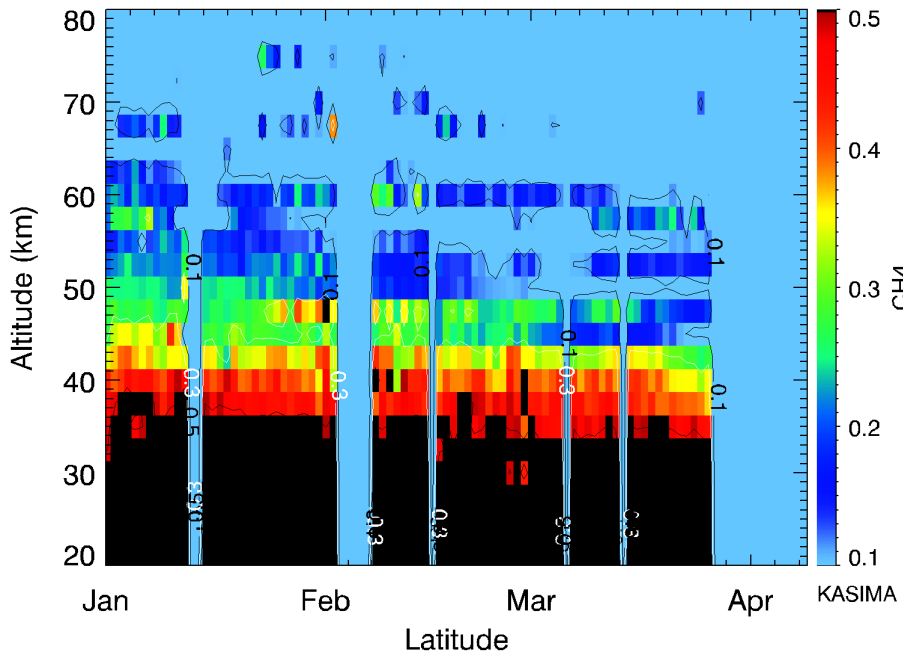


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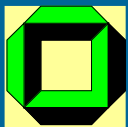
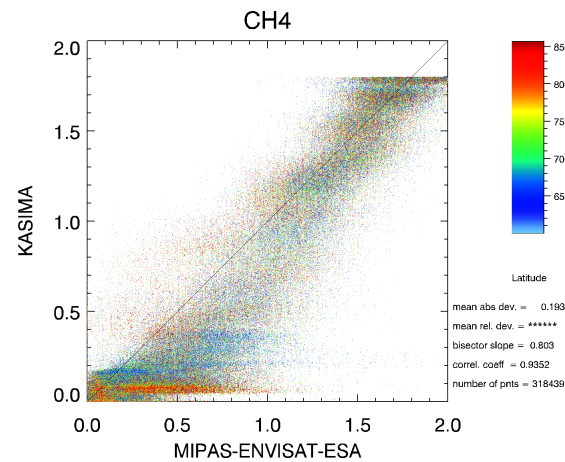
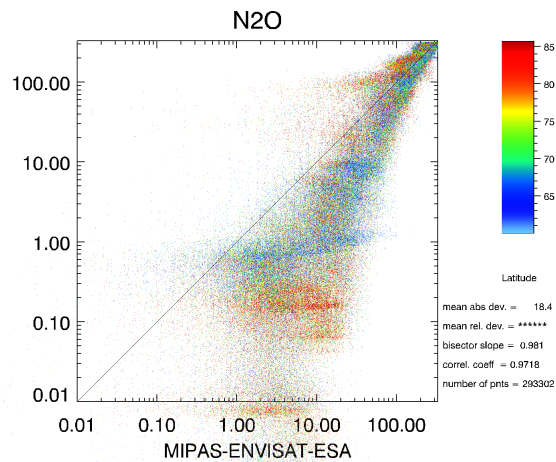
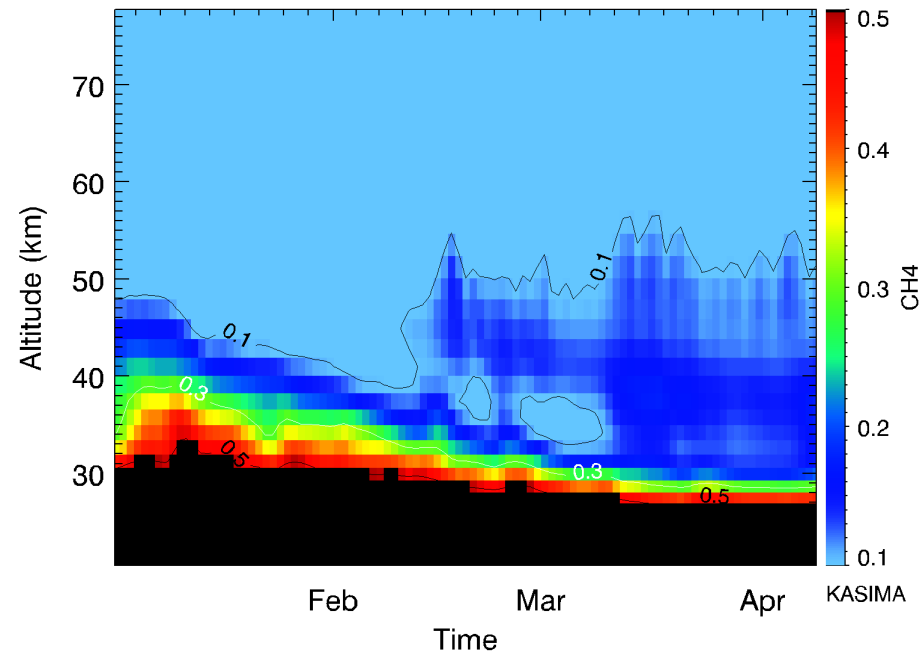


Descent in Arctic vortex winter 2003/2004(> 60N)

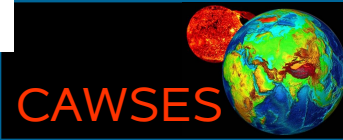
CH4 MIPAS-ENVISAT-ESA



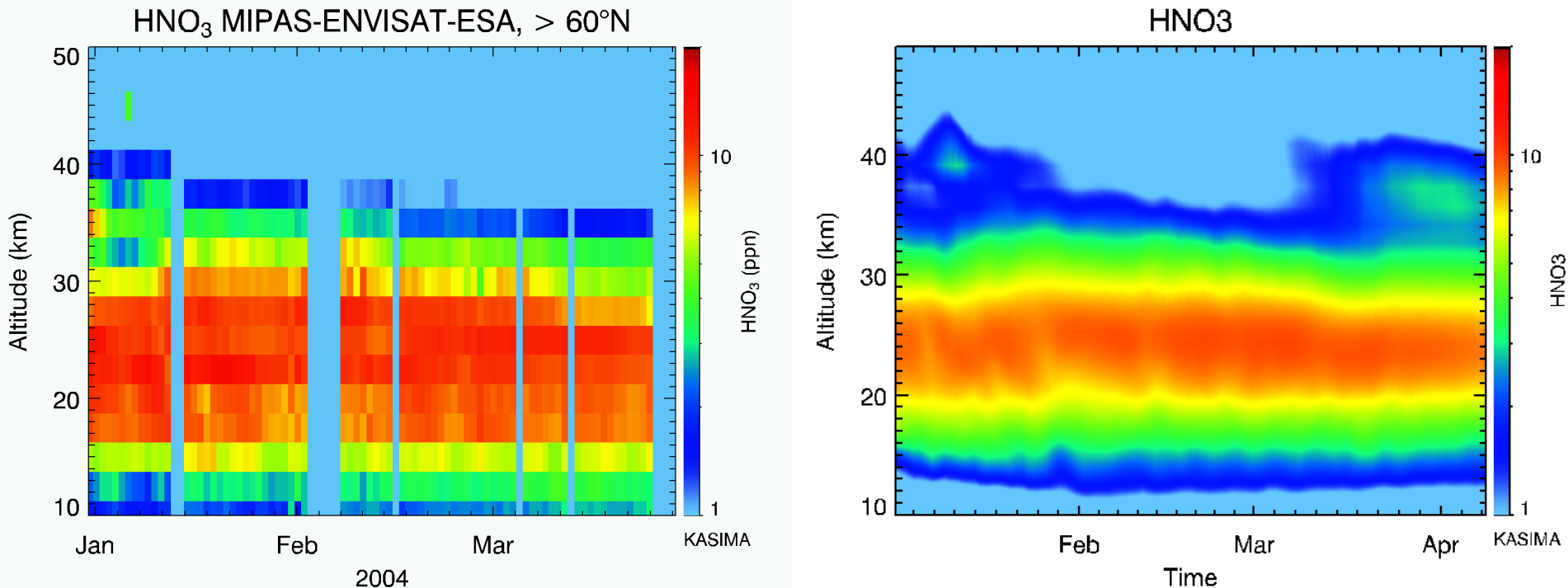
CH4



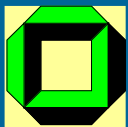
MANOXUVA



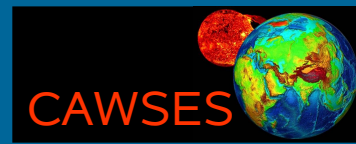
HNO₃ in the model and in observations



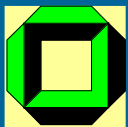
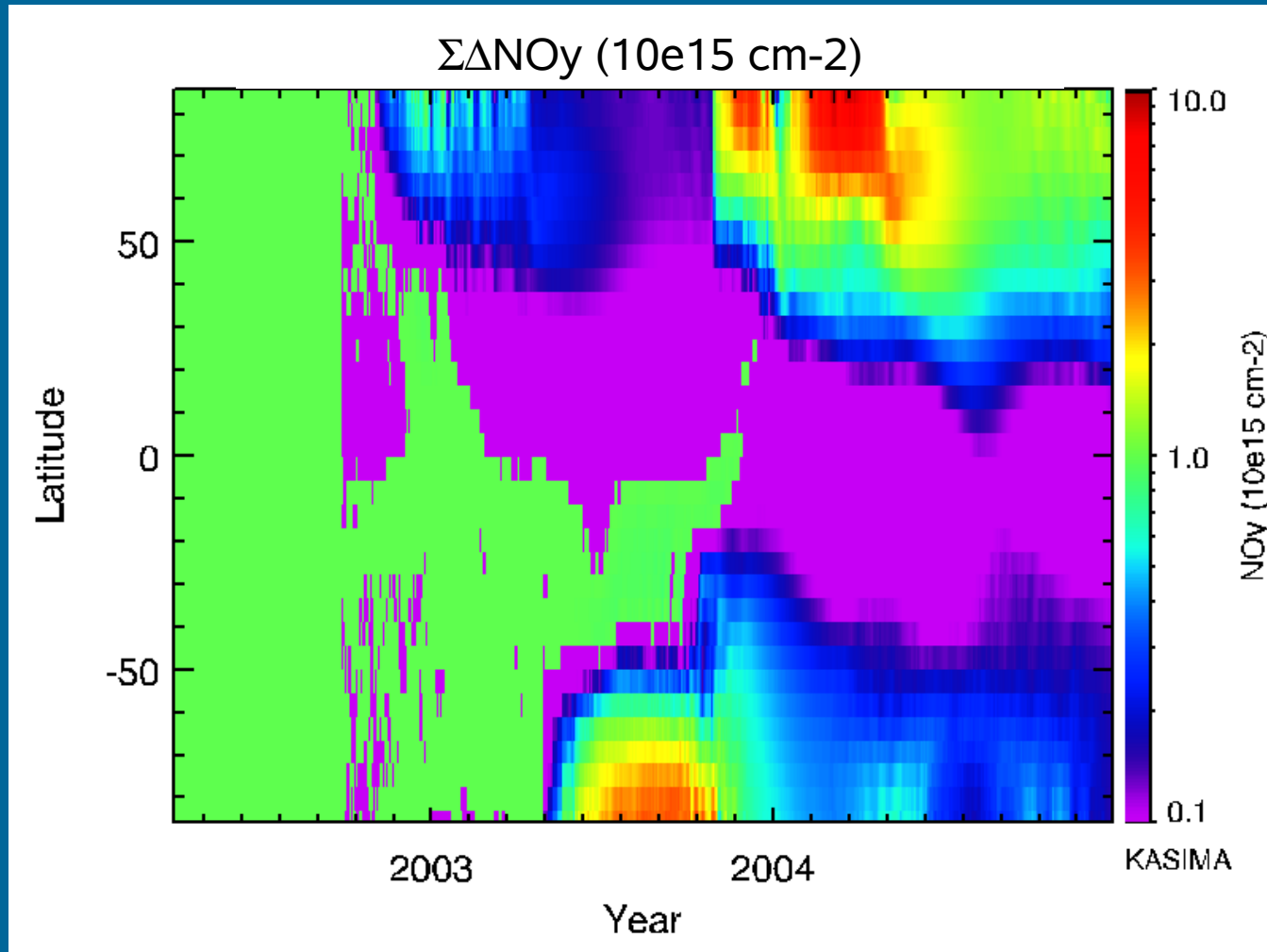
Late Arctic winter 2003/2004



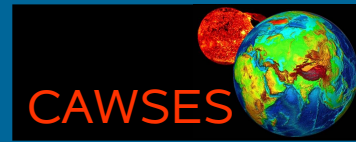
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Effect of NOx intrusions on total NOy

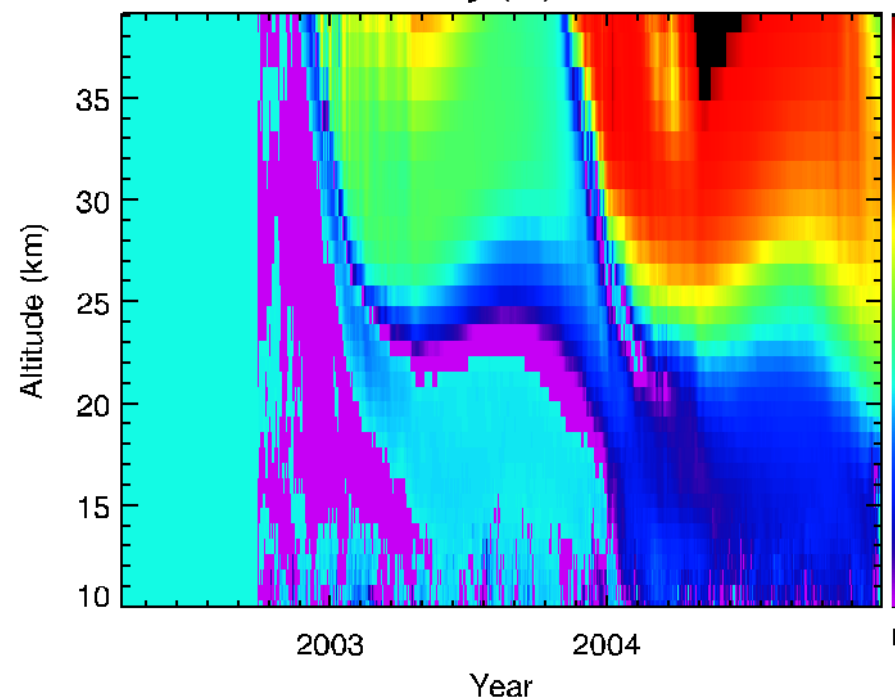


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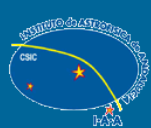
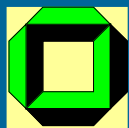
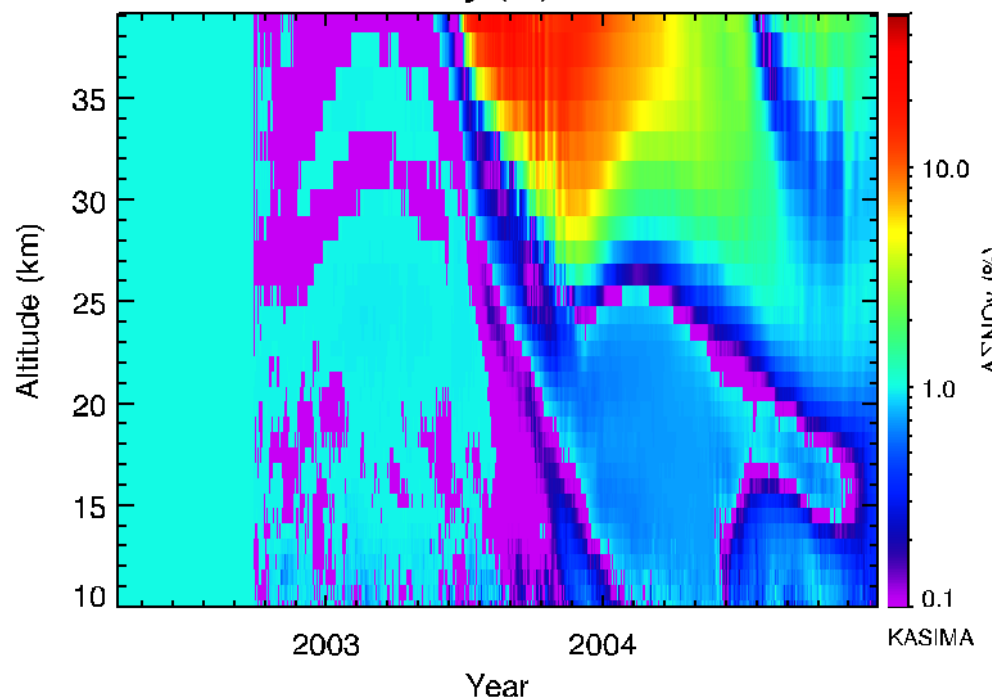


NO_y

$\Delta\Sigma\text{NO}_y$ (%) 45°N



$\Delta\Sigma\text{NO}_y$ (%) 45°S

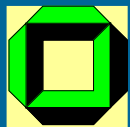
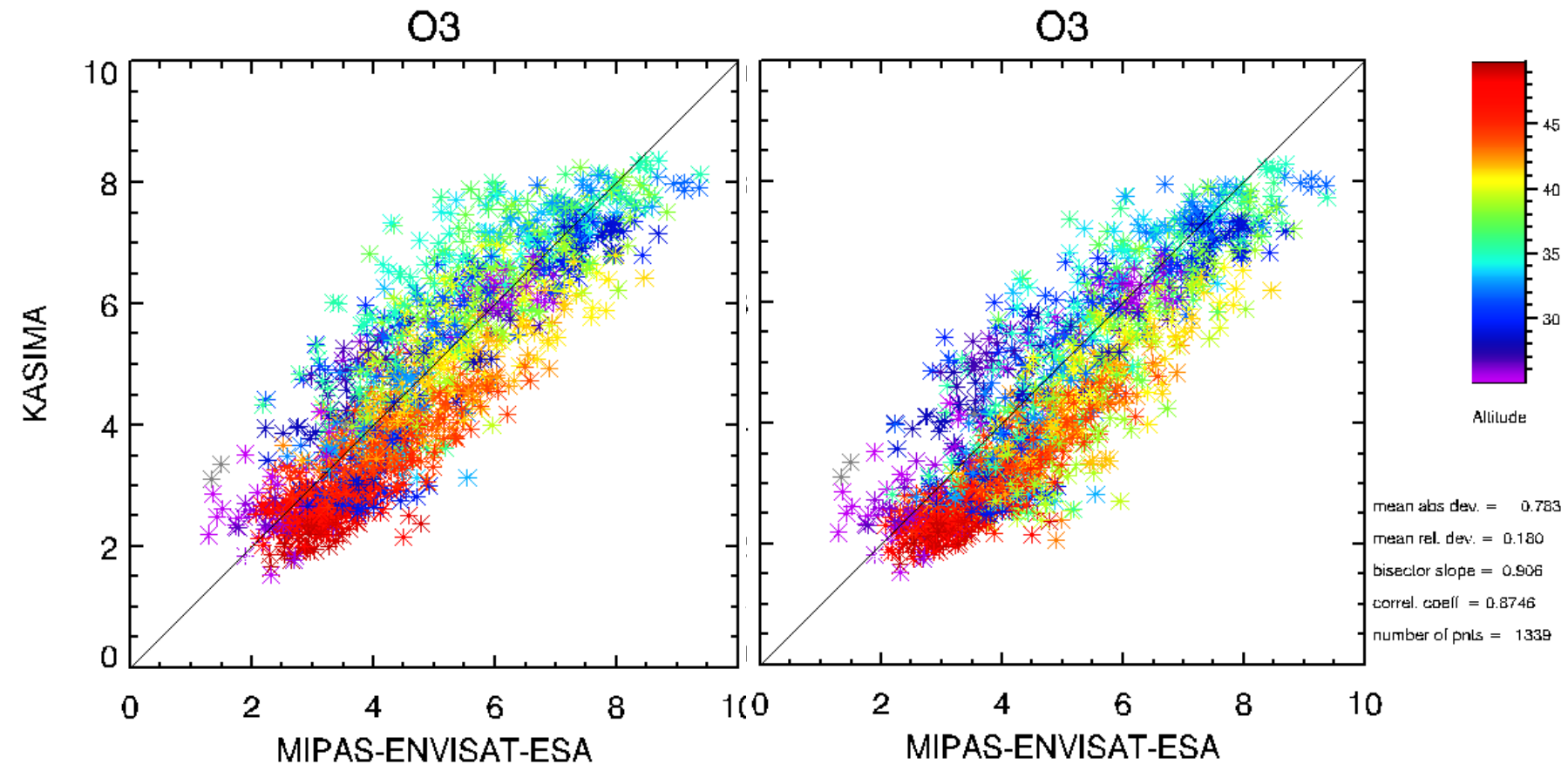


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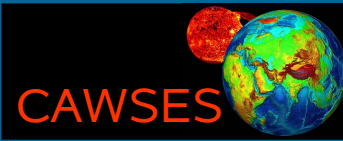


Ozone: Simulations vs. Observations

example antarctic winter, DOY 240 - 270



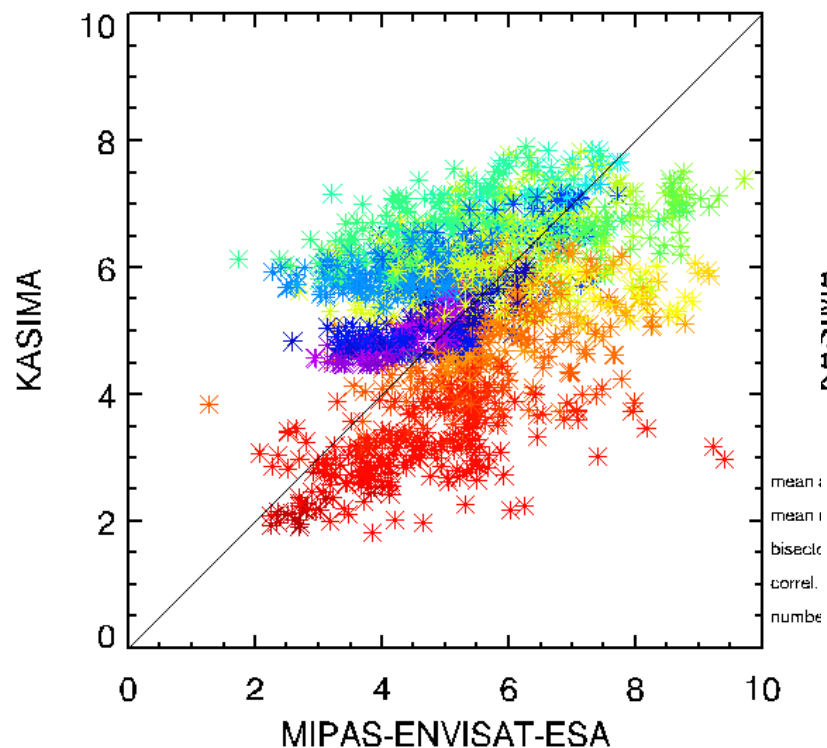
MANOXUVA



Correlation late winter 2003/2004 > 60°N DOY 40 - 80

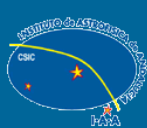
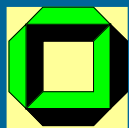
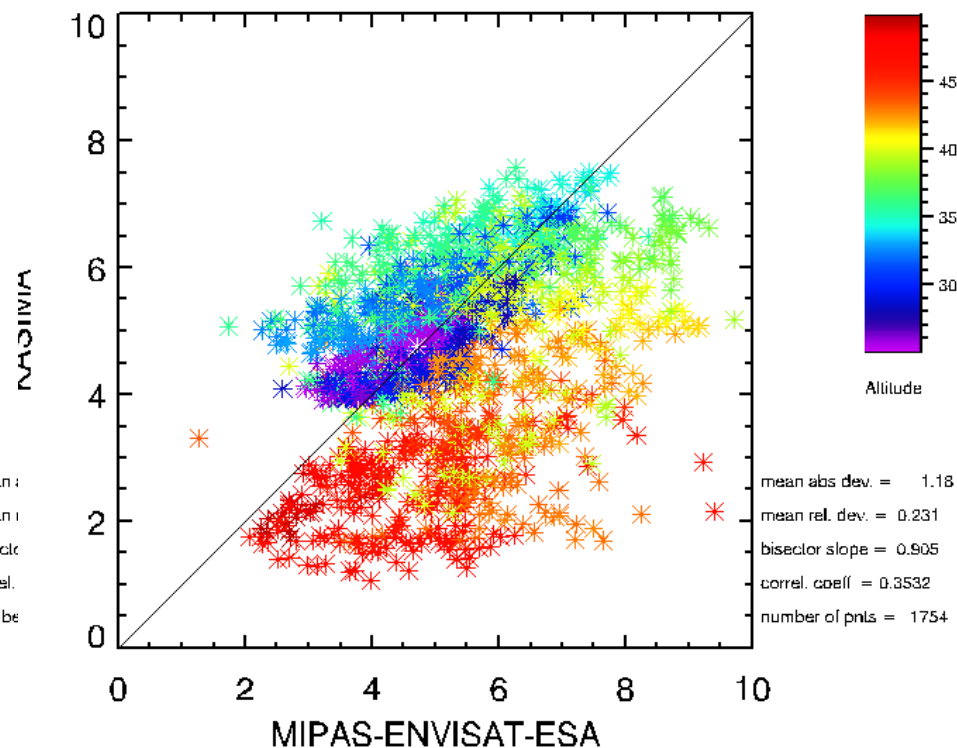
undisturbed

O3



disturbed

O3



MANOXUVA

