# **Global stratospheric aerosol extinction profile retrievals** from SCIAMACHY limb radiance: 7th Atmospheric Limb Conference algorithm description and validation



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 $S_v$  = noise covarinace matrix

This presentation participates

in the student award contest

## Introduction

Stratospheric aerosols significantly influence the radiative balance of the Earth and interact in stratospheric chemistry. Scattering of sunlight and influencing the cloud formation in the upper troposphere lead to a cooling effect of different magnitude under undisturbed background conditions and periods of volcanically enhanced aerosol load. Furthermore, stratospheric aerosols play an important role in stratospheric chemistry, especially ozone depletion.

Continuous global observations of stratospheric aerosols are essential for monitoring stratospheric aerosols. These observations can only be provided by satellite measurements. Limb observations of the scattered solar radiation combine high vertical resolution with near-global coverage, but require sophisticated radiative transfer modelling to perform the retrieval. The results can be validated with more accurate, but sparser occultation data sets. SCIAMACHY, the Scanning Imaging Absorption SpectroMeter for Atmospheric Chartography on ESA's Envisat spacecraft is one such instrument employing the limb-scatter observation geometry. It provided broadband limb radiance measurements from 2002-2012 covering a period of undisturbed background stratospheric aerosol conditions in the first years of its lifetime and several volcanic eruptions with stratospheric impact from 2005 onwards. Thus, its comprehensive data set offers the rare possibility to investigate stratospheric aerosols under changing aerosol loads. Goal of this work is to employ limb-scatter measurements with the SCIAMACHY instrument to retrieve stratospheric aerosol extinction profiles that form a valuable data set with global coverage on a daily basis.

For more information incl. a climatological interpretation of the complete SCIAMACHY stratospheric aerosol extinction data set see talk given by Lena A. Brinkhoff: Retrieval and variability of stratospheric aerosols from SCIAMACHY limb-scatter observations, Wednesday 9:20!

## Retrieval scheme

Non-linear inverse problem: Obtaining stratospheric aerosol extinction coefficients out of SCIAMACHY limb radiance profiles

Construction of the measurement vector [Bourassa et al., 2007]; Starting point: measured or simulated SCIAMACHY radiance profiles at 2 wavelengths, 470 and 750 nm (weak atmospheric absorption)

Step 1. Tangent height normalization. 2 2 2  $TH_{ref}$  = 35 km (aerosols below, staylight above)

# Comparison with co-located SAGE II data

### SAGE II:

- Sun occultation instrument, 1984 2005
- Measures aerosol extinction at 4 wavelengths, we are comparing to 525 nm.
- The data set is considered to be one of the stratospheric aerosol data sets with the highest accuracy.

#### **Comparison SCIAMACHY – SAGE II at collocated measuring points:**

- Collocation criteria: spatial difference < 500 km and temporal difference < 6h - data with SZA@TP > 87° filtered out
- overlap between SCIAMACHY and SAGE II: 01/01/2003 08/17/2005 - approx. 3500 collocations

The retrieved SCIAMACHY aerosol extinction is interpolated to 525 nm with a

$$I_N^{\lambda}(TH) = I^{\lambda}(TH)/I^{\lambda}(TH_{ref})$$

Assumption: fraction of ground-reflected sunlight in the limb radiance is similar at all tangent heights (including TH<sub>ref</sub>)

Advantages: reduction in albedo dependence, absolute calibration not required

Step 2: Wavelength pairing:
$$y(TH) = \ln \left( \frac{I_N^{\lambda_i}(TH)}{I_N^{\lambda_s}(TH)} \right)$$
 $\lambda_i = 750 \text{ nm}, \lambda_s = 470 \text{ nm}$ Rayleigh scattering ~  $\lambda^4$ , Mie scattering (spherical aerosols) >  $\lambda^4 \rightarrow$  Color-index-approach  $\lambda/\lambda_s$  amplifies the aerosol signal  
Improves the linearity of the inverse problem  $\rightarrow$  smaller linearization errors.Modified Optimal Estimation Method (starting with [Rodgers, 2000]:Basic equation: $\mathbf{y} = \mathbf{f}(\mathbf{x}) + \epsilon$  $\mathbf{y} = \text{normalized and paired SCIAMACHY limb radiances ("measurement vector")}$   
 $\mathbf{x} = \text{desired aerosol extinction profile ("state vector")}$  ( $j$  components)Non-linear basic equation: $\hat{\mathbf{y}} = \mathbf{K}\hat{\mathbf{x}} + \epsilon$ with $\hat{\mathbf{y}} = \mathbf{y} - \mathbf{y}_a$  $\hat{x}_j = (x_j - x_{a_j})/x_{a_j}$   
 $\mathbf{y}_a = \mathbf{K}\mathbf{x}_a$ Non-linear basic equation: $\hat{\mathbf{y}} = \mathbf{K}\hat{\mathbf{x}} + \epsilon$ with $\hat{\mathbf{y}} = \mathbf{y} - \mathbf{y}_a$  $\hat{x}_j = (x_j - x_{a_j})/x_{a_j}$  $\mathbf{x}_a = a \text{ priori aerosol extinction } \mathbf{y}_a = \mathbf{K}\mathbf{x}_a$ Iterative solution: using  $x_{n-1}$  instead of  $x_a$  in each iterative step (important difference to OEM!)K = weighting function matrix  
 $\mathbf{S}_a = a \text{ priori covariance matrix}$ 

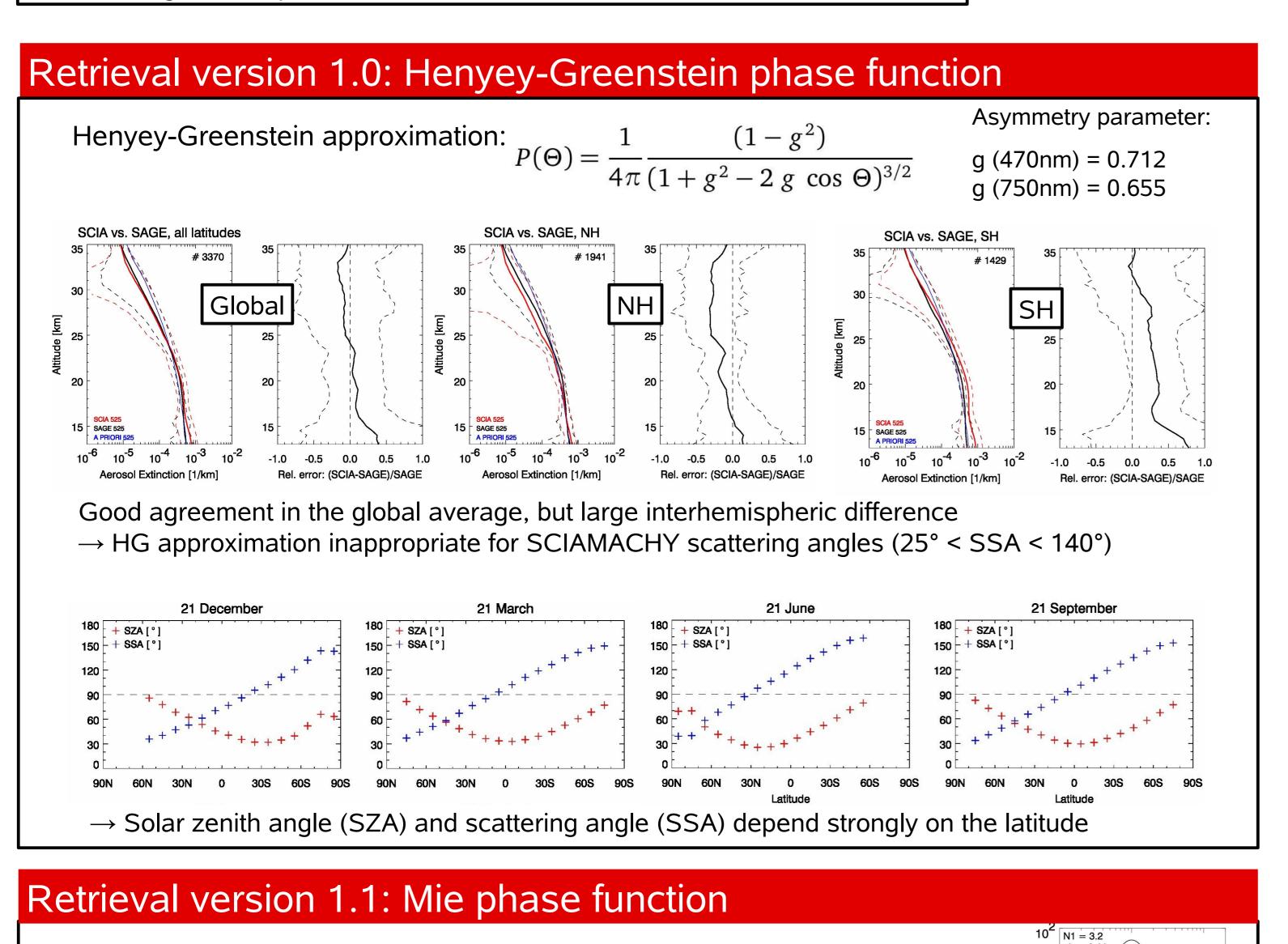
$$\mathbf{x}^* = (\mathbf{K}_i^{\mathrm{T}} \mathbf{S}_y^{-1} \mathbf{K}_i + \mathbf{S}_a^{-1})^{-1} \mathbf{K}_i^{\mathrm{T}} \mathbf{S}_y^{-1} (\mathbf{y} - \mathbf{y}_i) \quad \text{with} \quad x_j^* = (x_{i+1_j} - x_{i_j})/x_{i_j}$$

Advantage: more robust against bad a priori vs. disadvantage: information from a priori strongly reduced

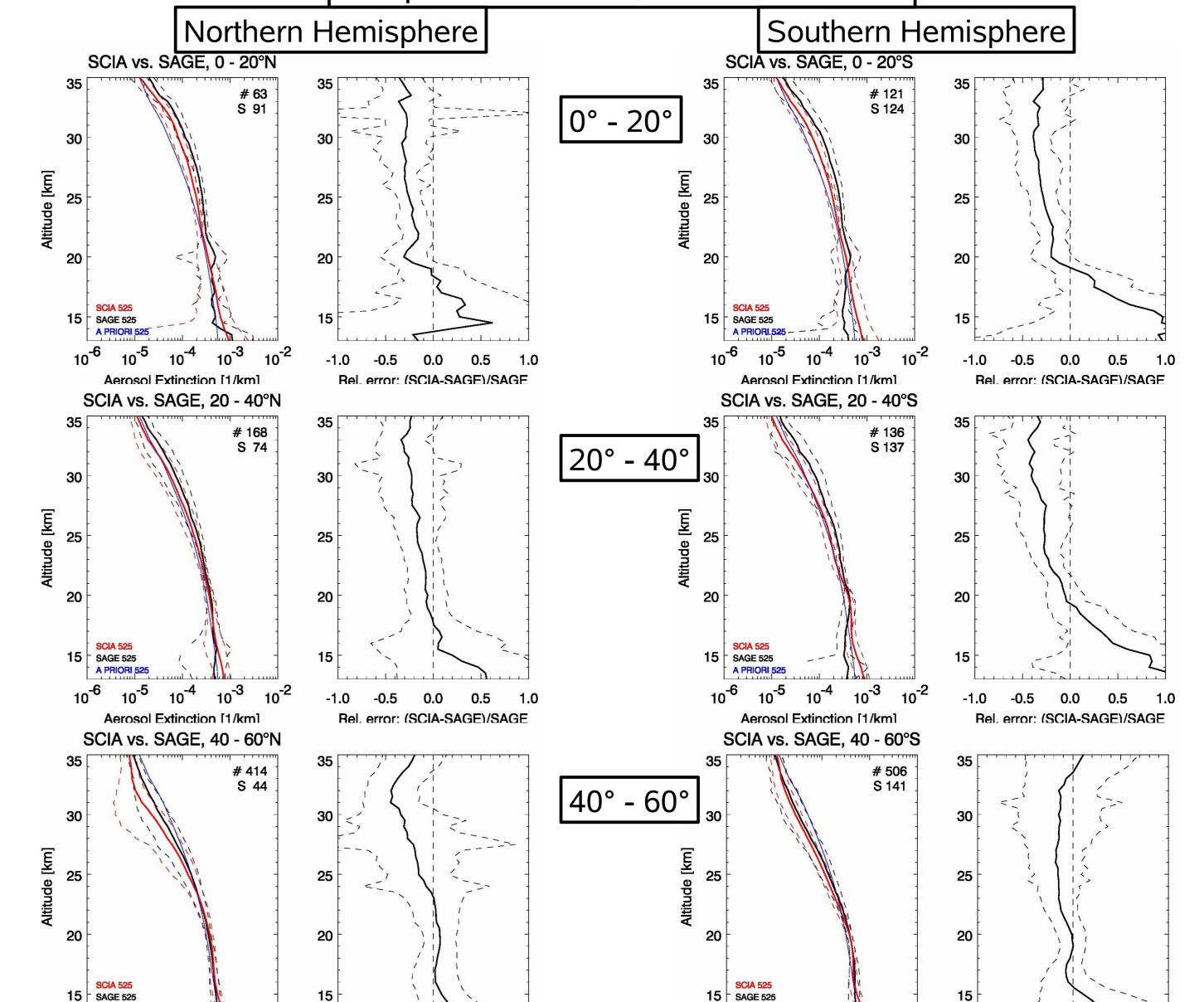
Radiative transfer model and retrieval code SCIATRAN 3.1 [Rozanov et al., 2013] is used to apply the scheme to the SCIAMACHY data

|                     | Mie phase function, 20° latitude bins |  |  |          |            |
|---------------------|---------------------------------------|--|--|----------|------------|
| lartharn Hamicahara |                                       |  |  | Couthorn | Homicnhoro |

#### constant Ångstrøm exponent of 1.43.



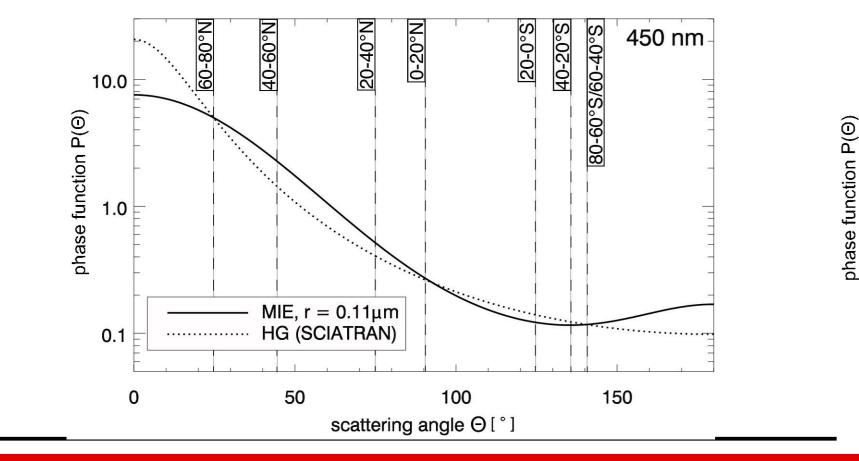




Particle size distribution above Laramie, Wyoming, USA (41°N); in situ measurements by [Deshler, 2003] for background aerosols, 06.05.2006: Bimodal lognormal distribution • we use the small mode as monomodal distribution:

Median radius r = 0.11  $\mu$ m, distribution width  $\sigma$  = 1.37

 $\rightarrow$  used to calculate Mie phases function for two wavelengths:



## Summary

**Retrieval version 1.0**: good agreement with SAGE II data in the global average

20° latitude bins: systematic interhemispheric differences with values of in general around 20 % to up to 50 % underestimation in the northern hemisphere (~20%), overestimation in the southern hemisphere (~30%, max. 60%).

10.0

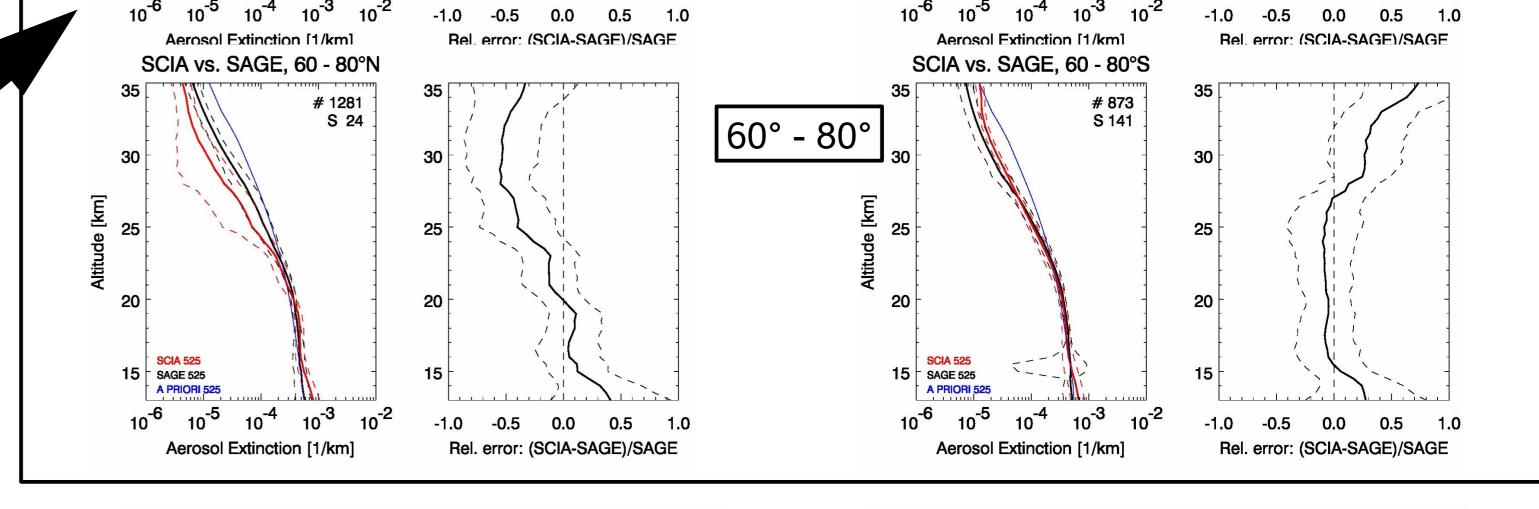
1.0

0.1

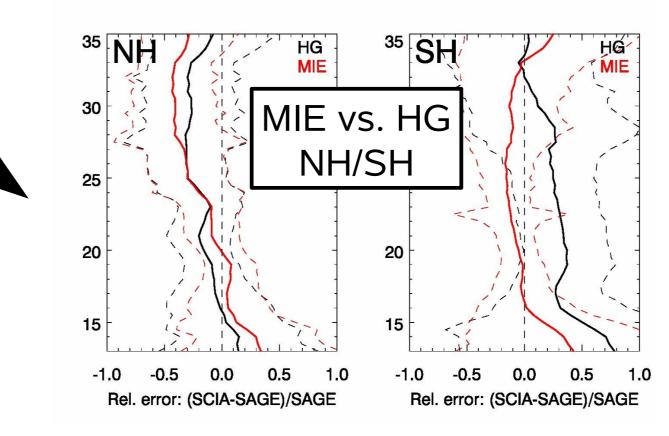
Retrieval version 1.1: significant reduction of the interhemispheric difference.

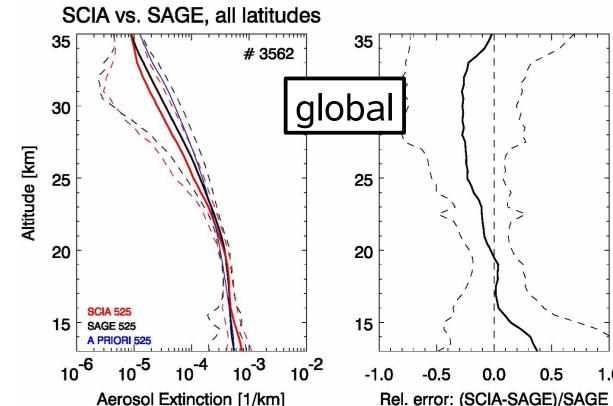
SH: difference in the southern hemisphere was reduced to values of typically 0–15 % (15-33 km)

NH: bias remained in the northern hemisphere in comparison with SAGE II measurements.



A PRIORI 525





#### Selected references:

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σ1 = 1.37

20-0°S

100

scattering angle  $\Theta[°]$ 

20-40°

MIE,  $r = 0.11 \mu m$ 

HG (SCIATRAN)

50

40-20°S

N2 = 0.01r2 = 0.38

150

σ2 = 1.07 \

800 nm

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