

Stirring and mixing of stratospheric filaments around the subtropical jet as observed by the CRISTA-NF infrared limb sounder

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CRISTA-NF

An instrument realises its potential

CRISTA-NF Instrument

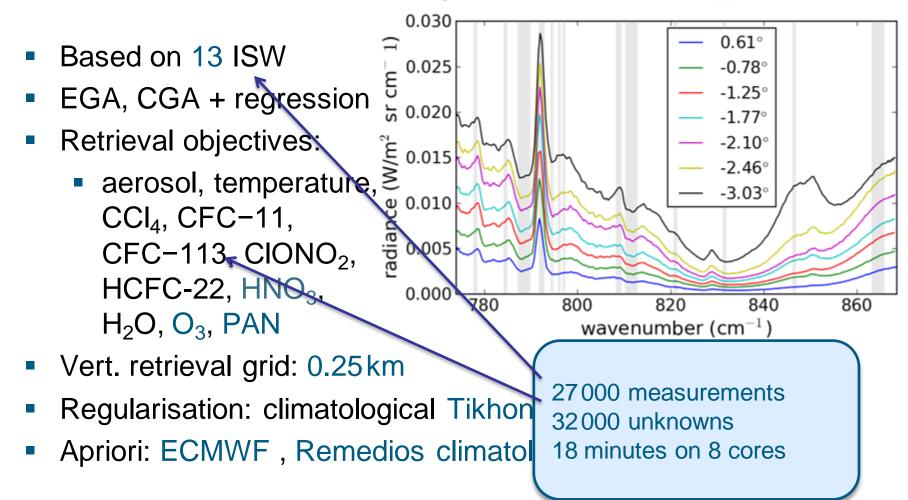
- CRISTA-NF: Cryogenic Infrared Spectrometers and Telescope for the Atmosphere – New Frontiers
- Platform: M55-Geophysica (max. flight altitude: ~21 km)
- Method: Passive mid-infrared limb sounding (4 15µm)
- Profile acquisition time: ~70s (60·1.2s)
- Vert. Sampling: ~0.25km
- Hor. Sampling: ~15km







CRISTA-NF Cross-section retrieval setup

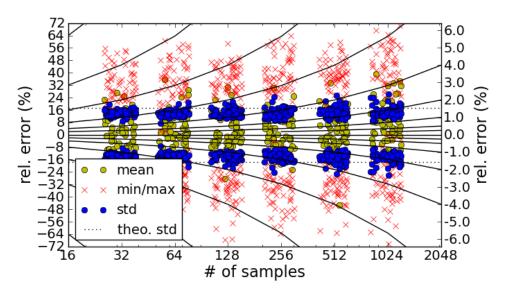


Retrieval for large-scale inverse problems Linearised diagnostics

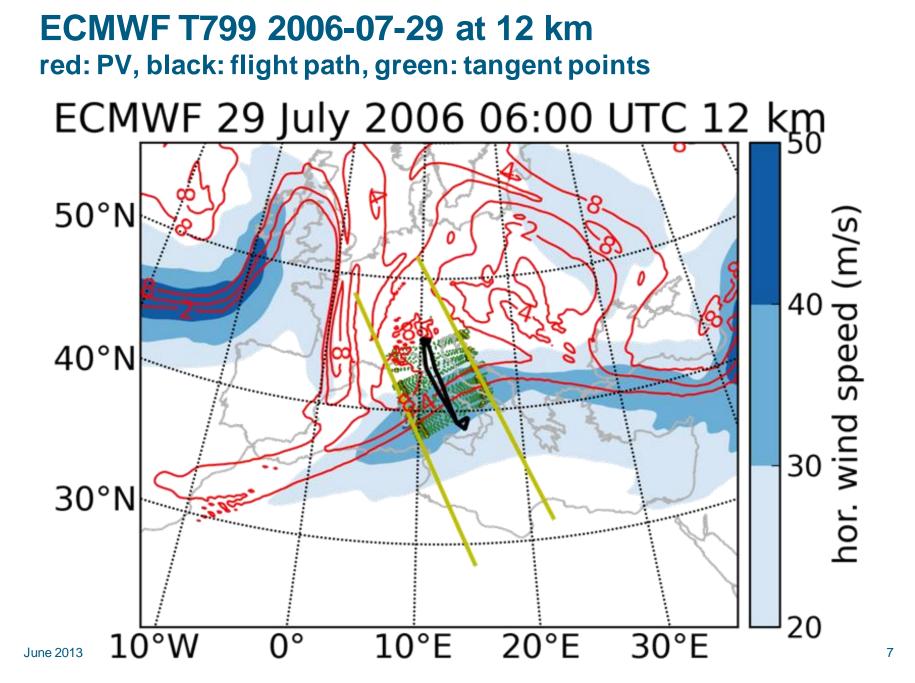
$$\boldsymbol{x}_{i}^{\mathrm{mc}} = \boldsymbol{x}_{\mathrm{f}} - \left(\mathbf{S}_{\mathrm{a}}^{-1} + \mathbf{F}'(\boldsymbol{x}_{\mathrm{f}})^{T} \mathbf{S}_{\epsilon}^{-1} \mathbf{F}'(\boldsymbol{x}_{\mathrm{f}})\right)^{-1} \left(\mathbf{F}'(\boldsymbol{x}_{\mathrm{f}})^{T} \mathbf{S}_{\epsilon}^{-1} \boldsymbol{\epsilon}_{i}^{\mathrm{mc}}\right).$$

"Monte carlo" diagnostics: 1) Determine Error vectors according to known or assumed Gaussian distributions.

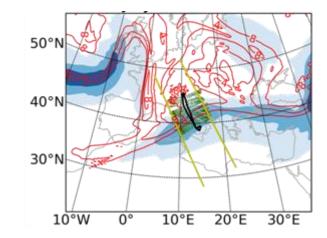
- 2) Derive perturbed solutions.
- 3) Determine effect on result by calculating the variance of the perturbed solutions.
 4) Repeat for every important error source.

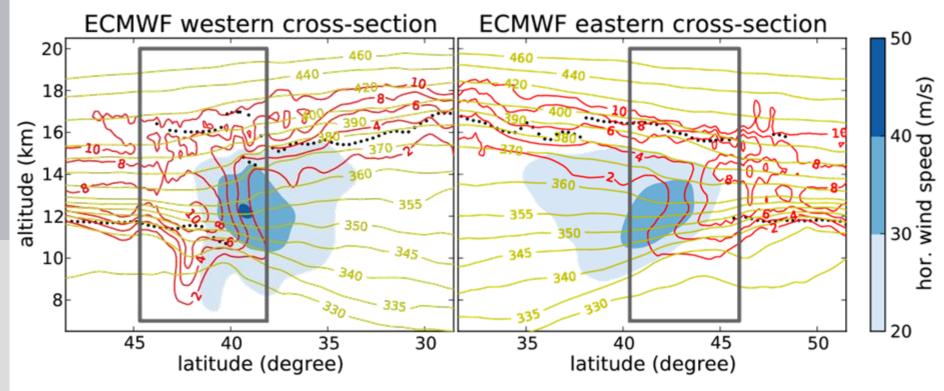


METEOROLOGICAL SETTING



ECMWF T799 2006-07-29 Vertical cross-sections



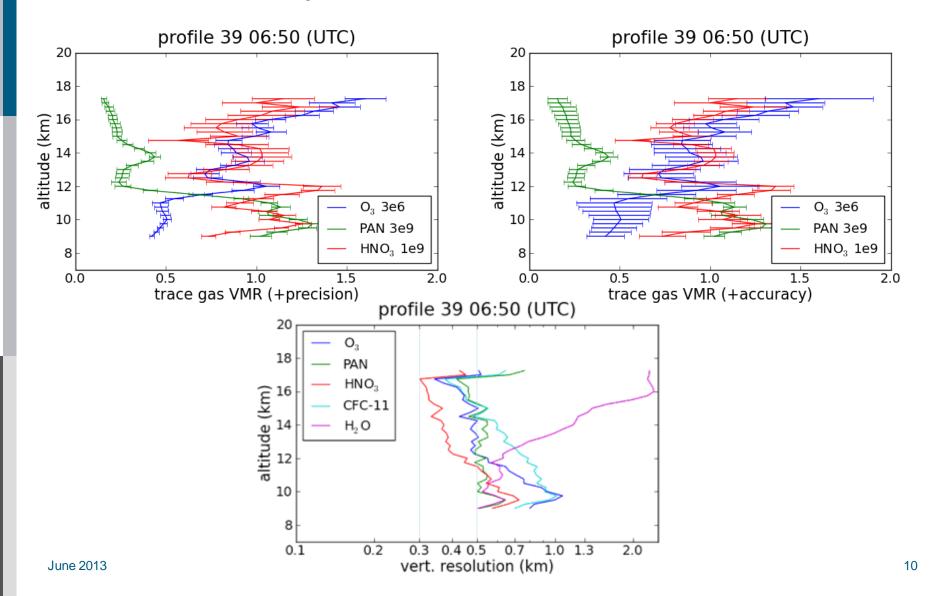


red: PV (PVU), yellow: Theta (K), blue: hor. wind speed (m/s), black dots: thermal lapse-rate tropopause, gray square: measured area

VALIDATION

Quality of retrieved trace gases

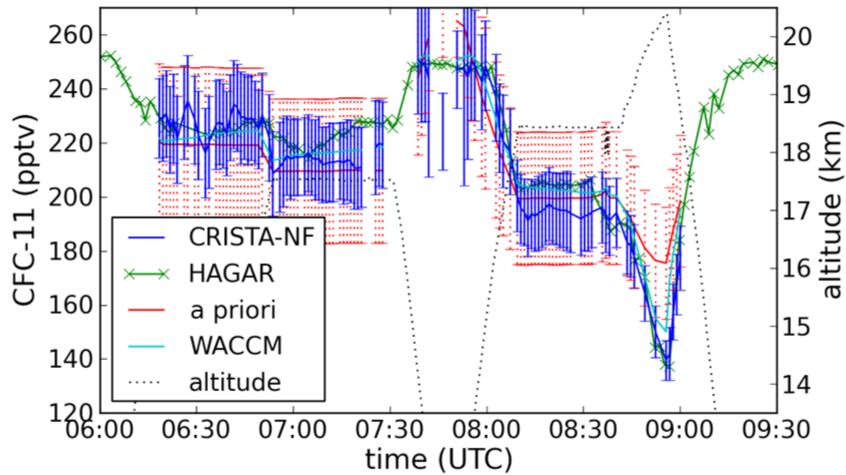
Retrieval quality Precision, accuracy, and vertical resolution



Retrieval quality

"Validation"

comparison at flight level

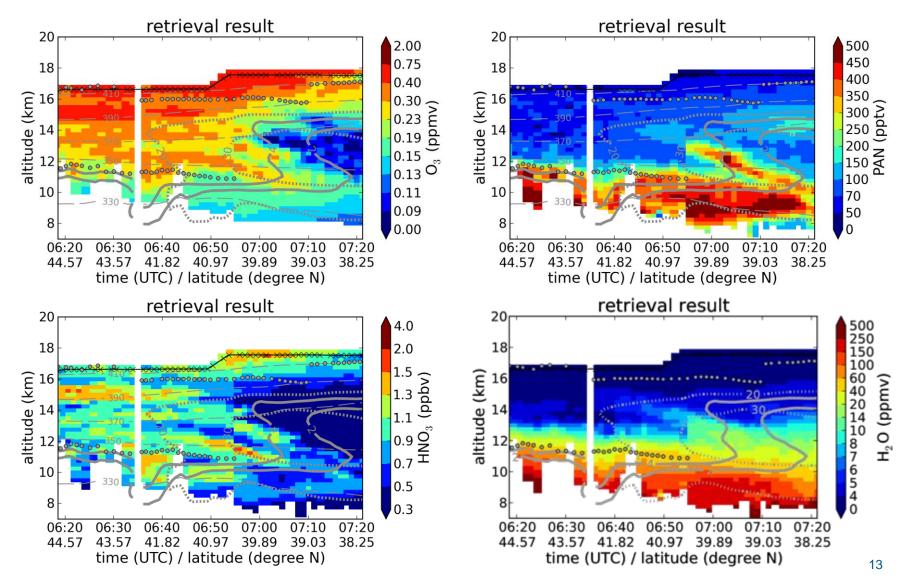


HAGAR CFC-11 data provided by M. Volk. WACCM data provided by D. Kinnison. ¹¹

SMALL SCALE FILAMENTS

Exploiting the resolution of CRISTA-NF

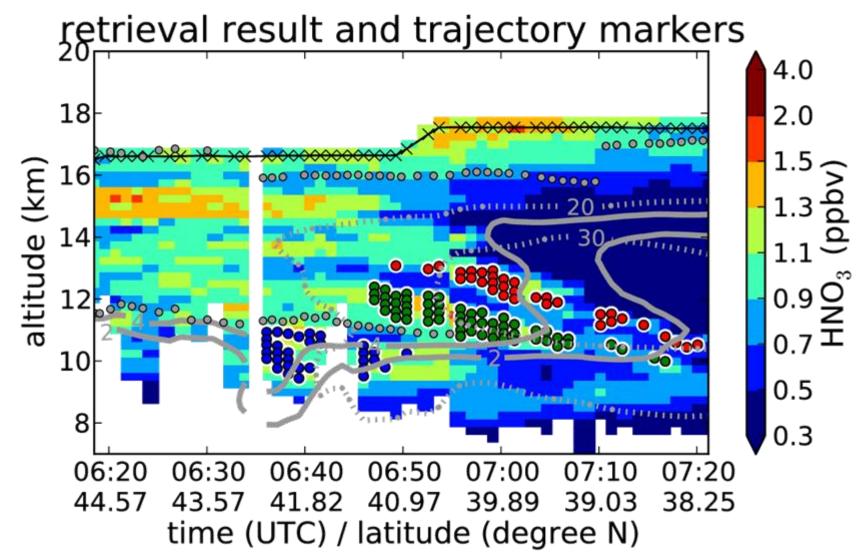
AMMA/CRISTA-NF retrieval results 2006-07-29 Western cross-sections



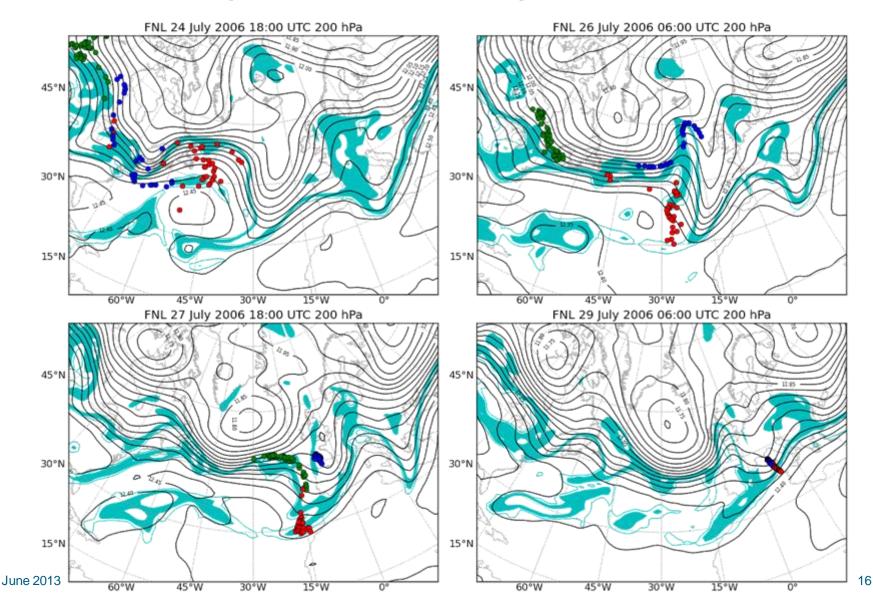
STIRRING

How one breaking Rossby-wave produces the imaged situation

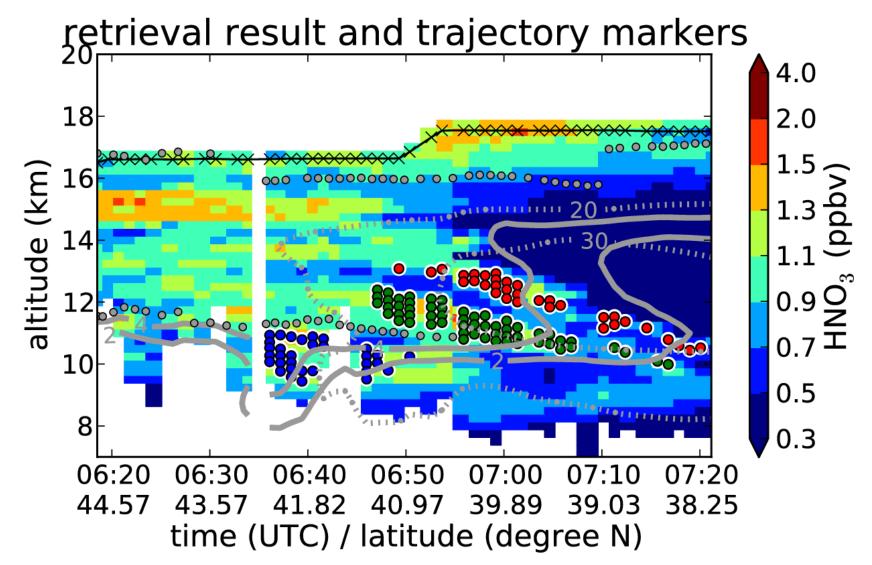
AMMA/CRISTA-NF retrieval result 2006-07-29 red, green, and blue dots mark stratospheric filaments



FNL 2006-07-24 to 2006-07-29 at 200 hPa black: GPH, turquois line: 1.5 PVU, turquois area: 2 to 4 PVU



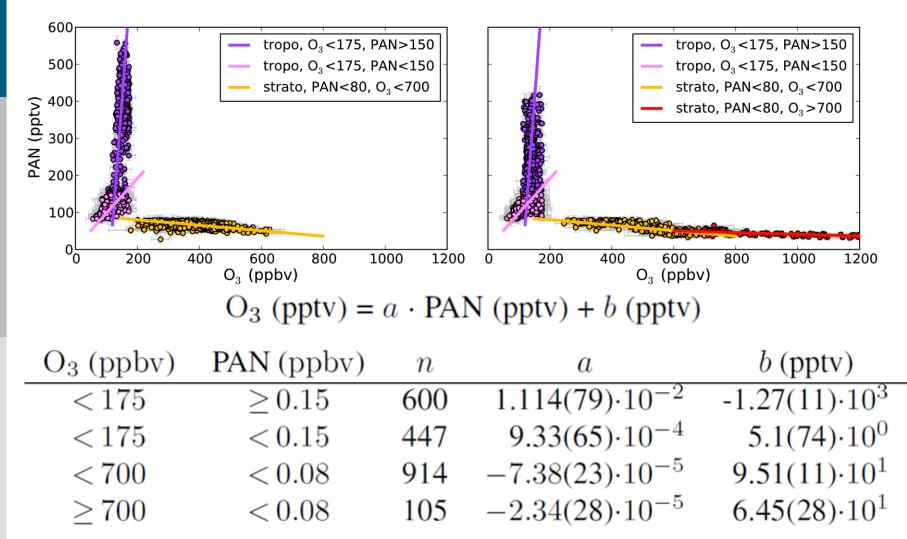
AMMA/CRISTA-NF retrieval result 2006-07-29 red, green, and blue dots mark stratospheric filaments



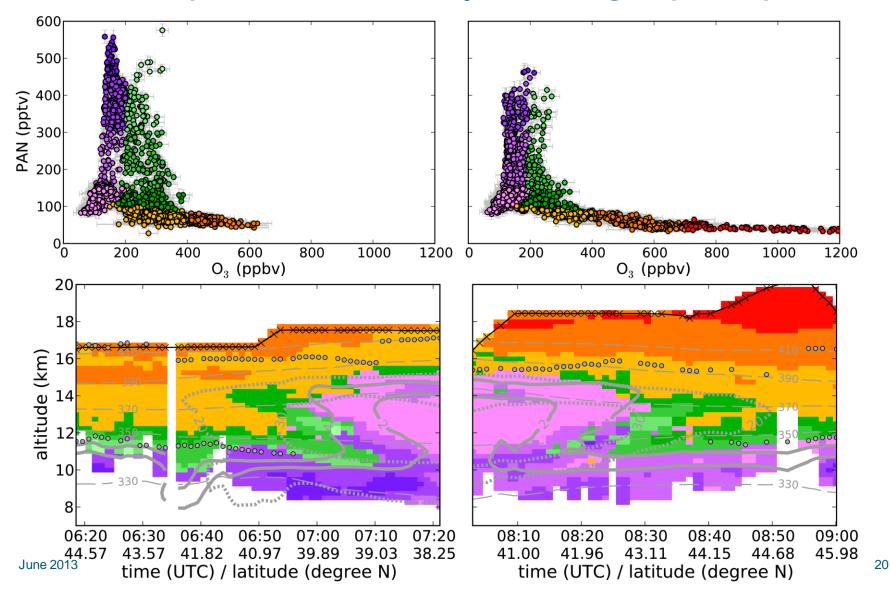
MIXING

Connecting tracer-tracer space and geophysical space

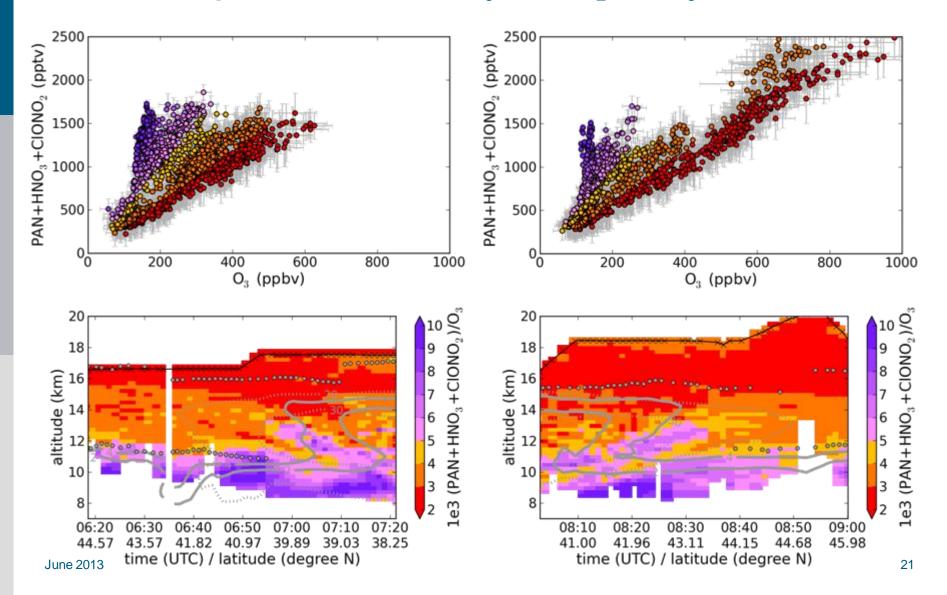
AMMA/CRISTA-NF retrieval result 2006-07-29 Relationship between PAN and O₃ in tracer / geospatial space



AMMA/CRISTA-NF retrieval result 2006-07-29 Relationship between PAN and O₃ in tracer / geospatial space



AMMA/CRISTA-NF retrieval result 2006-07-29 Relationship between PAN+HNO₃+CIONO₂ and O₃ and ratio



Stirring and mixing in the UTLS Conclusions

- Fine-scale filaments with a vertical extent of less than 0.8km could be observed. The originating breaking wave could be identified using trajectory calculations
- Vertical resolution is sufficient but also necessary to map the mixing zone between UT and LS
- Available trace gases (nearly) allow a derivation of NO_y and a classification of air as pristine or polluted
- Analysis lacks in several (in principle retrievable!) trace gases and also temporal and spatial coverage:
 - GLORIA offers much higher spectral resolution at better spatial resolution
 - PREMIER or a similar instrument would provide global coverage at comparable quality