

UV/vis/near-IR Limb Measurements Onboard The NASA Global Hawk In The Tropical Tropopause Layer (TTL)

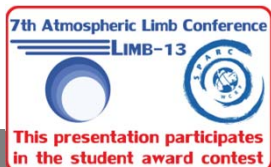
Bodo Werner (1)



- 7th Limb Conference and Workshop, Bremen -

Collaborators:

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Bodo Werner, 19.6.2013

UCLA AOS



NASA ATTREX Objectives and DOAS Contribution

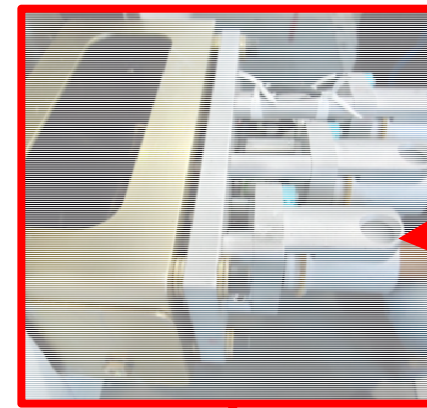
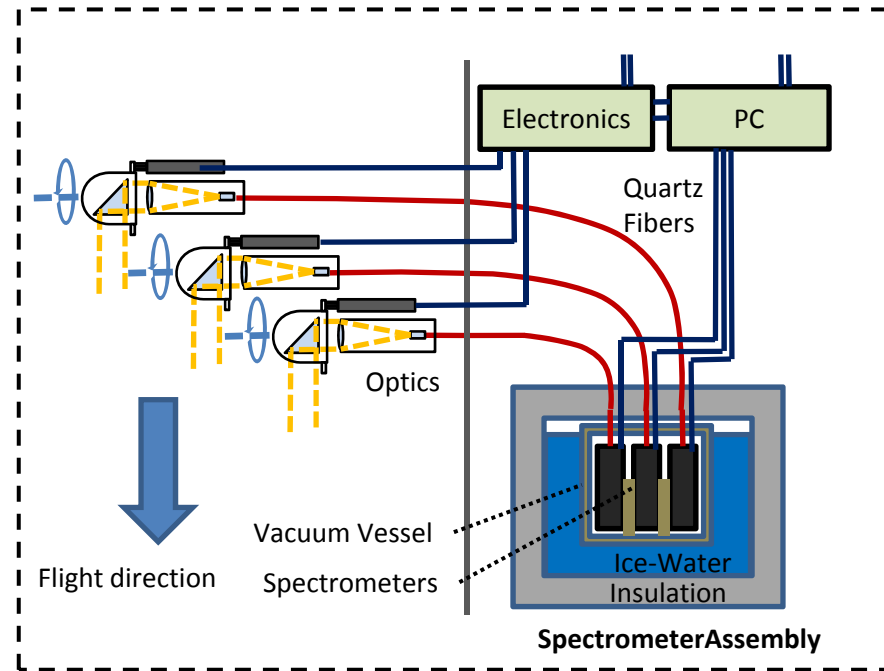
NASA ATTREX overarching goals and mini-DOAS contributions

1. **Across TTL transport** due to deep convection, slow large-scale ascent waves, et cetera
2. **Microphysics**
 - Mini-DOAS: Detection of water in the TTL and of ice particles (sub-visible cirrus clouds) and quantification of some of their optical properties
3. **TTL Photochemistry**
 - Mini-DOAS: The oxidation capacity in the TTL by observation of some relevant nitrogen (NO_x) and halogen compounds (BrO, and IO)
 - Mini-DOAS: The budget and photo-chemistry of chlorine, bromine and iodine in the TTL and lowermost stratosphere
4. **Radiative heating:**
 - Mini-DOAS: Validation of short-wave radiation schemes

Campaigns:

- Test campaign at EAFB (Ca) in Oct/Nov. 2011 with 3 deployments ✓
- 1st scientific campaign at EAFB (Ca) in Jan./Feb. 2013 with 6 deployments ✓
- 2nd scientific campaign at Guam in Jan./Feb. 2014
- 3rd scientific campaign at Darwin in June/July 2014

The GH mini-DOAS Instrument



Telescopes

Diffuser for direct solar reference



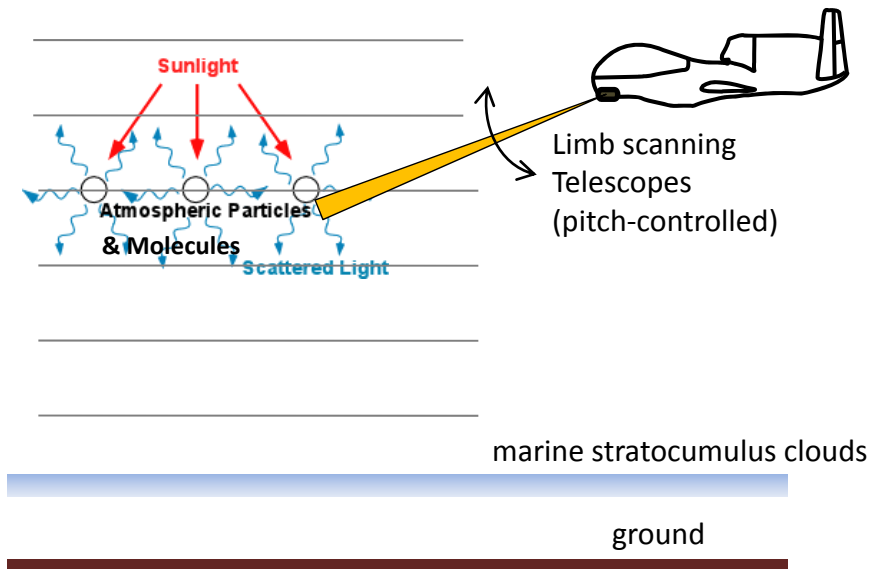
Spectrometers

Quartz fibers



Telescopes with cover

Major Measurement Features



- Limb viewing geometry
- Pointing into flight direction (-> sensitive to the pitch of the aircraft)
- Telescope elevation angles: $+1^\circ$, 0° , -0.5° , -1° , -1.5° , -2° , -2.5° , -3° , -4° , -7° , -15°
- During ascents/descents: fixed limb (-0.5° , which yields the highest sensitivity at flight altitude)

Major optical properties and measured species of the GH mini-DOAS instrument

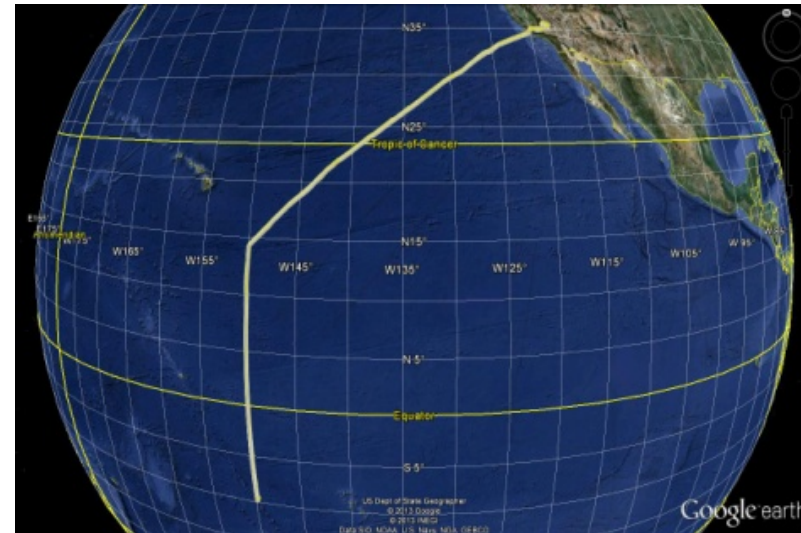
Channel	Wavelengths	Resolution	Vertical FOV	Measured Species
UV	300 - 380 nm	0.8 nm	0.19°	BrO, O ₃ , O ₄ (361nm)
vis	410 - 530 nm	0.9 nm	0.24°	NO ₂ , O ₃ , O ₄ (477nm)
near-IR	900 - 1700 nm	20 nm	0.19°	H ₂ O (vapor, liquid, ice)

2013 NASA Dryden Deployment - Flights

Flights DRFC deployment 2013

Date	Duration
Jan 19	6.5 hrs (test flight)
Feb 5-6	24.5 hrs
Feb 9-10	24.3 hrs
Feb 14-15	24.5 hrs
Feb 11-22	24.6 hrs
Feb 26-27	24.4 hrs
Mar 1-2	24.1 hrs

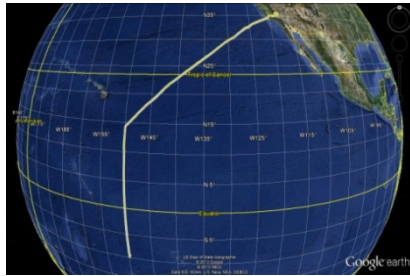
Flight path Feb 9-10



Southernmost point: 10°S

The GH mini-DOAS instrument worked well during all conducted flights
→ 150 hrs of data (from which approx. 80 hrs at sunlight)

Science Flight 2 2013 – (diff.) Slant Column Densities



Radiance
@450nm [a.u.]

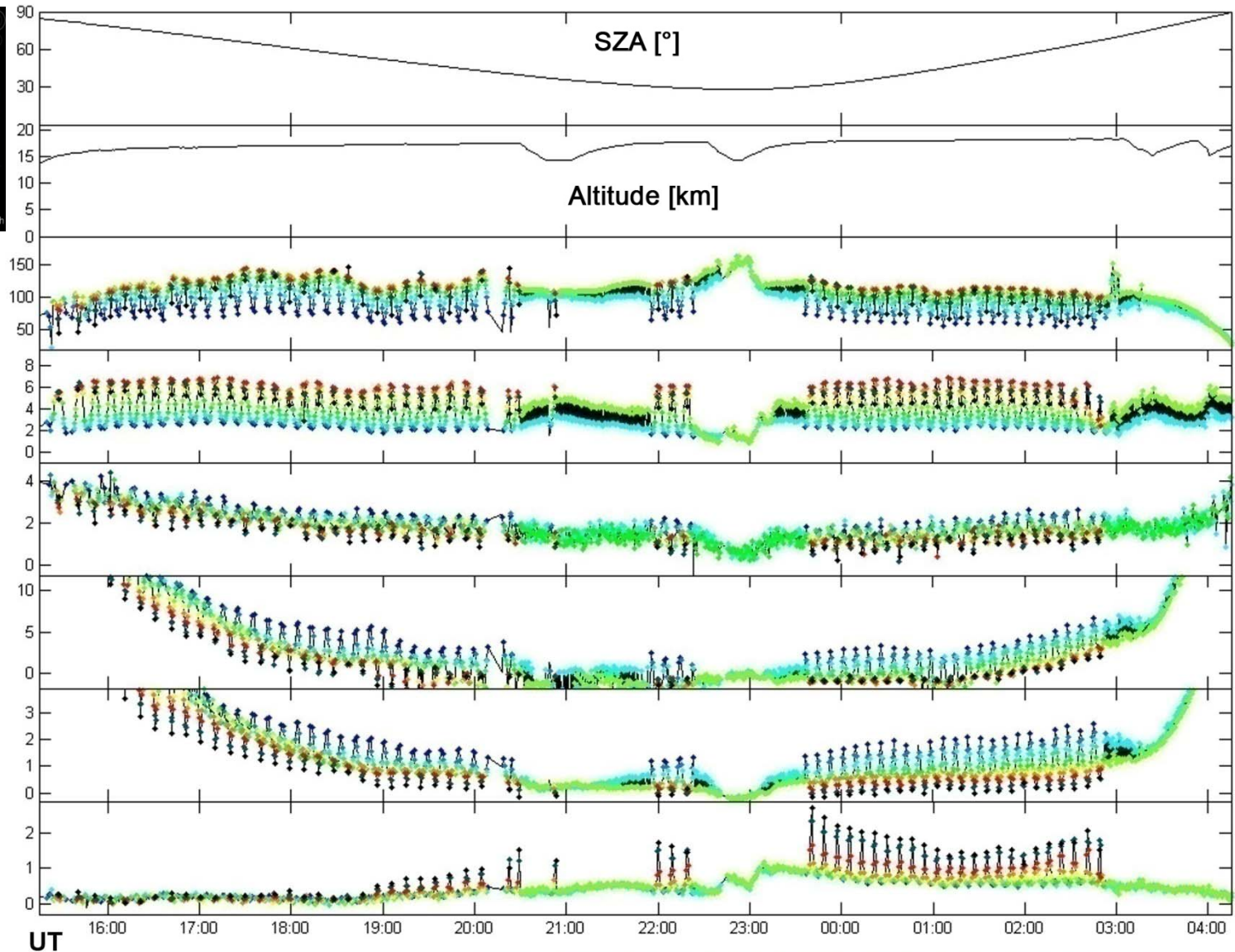
O₄ dSCD
[10⁴³ molec²*cm⁻⁵]

BrO dSCD
[10¹⁴ molec*cm⁻²]

NO₂ dSCD
[10¹⁵ molec*cm⁻²]

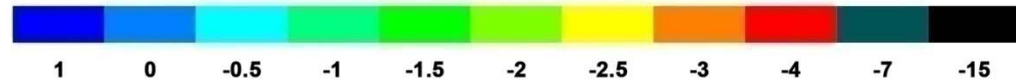
O₃ dSCD
[10¹⁹ molec*cm⁻²]

H₂O dSCD
[10²³ molec*cm⁻²]

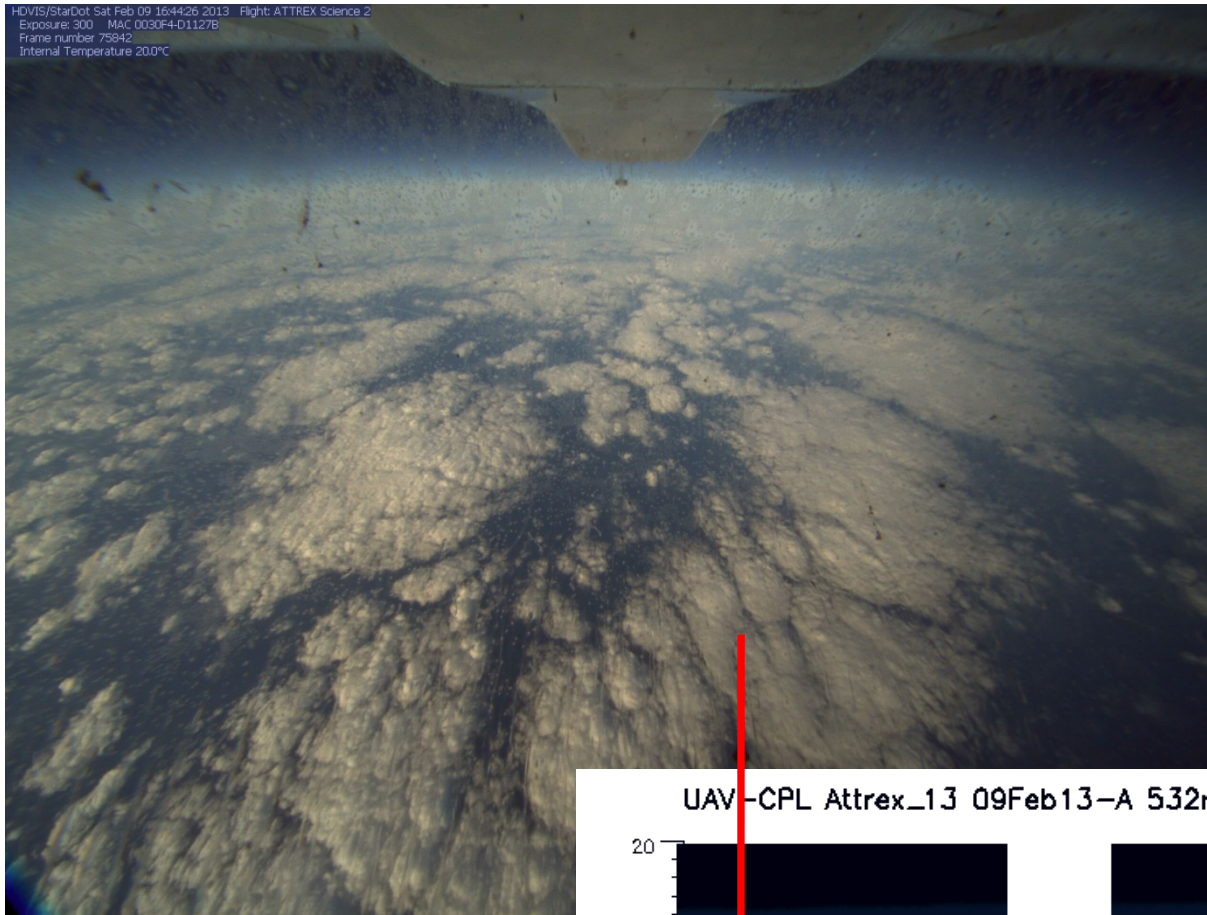


UT

Telescope Elevation:

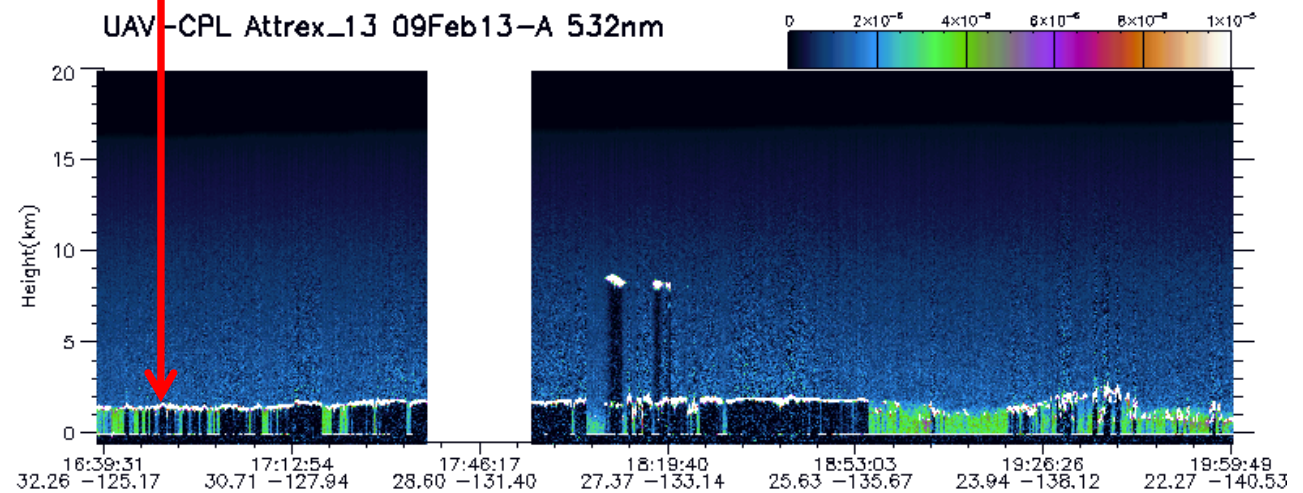


Typical Marine StratoCumulus (MSC) Clouds



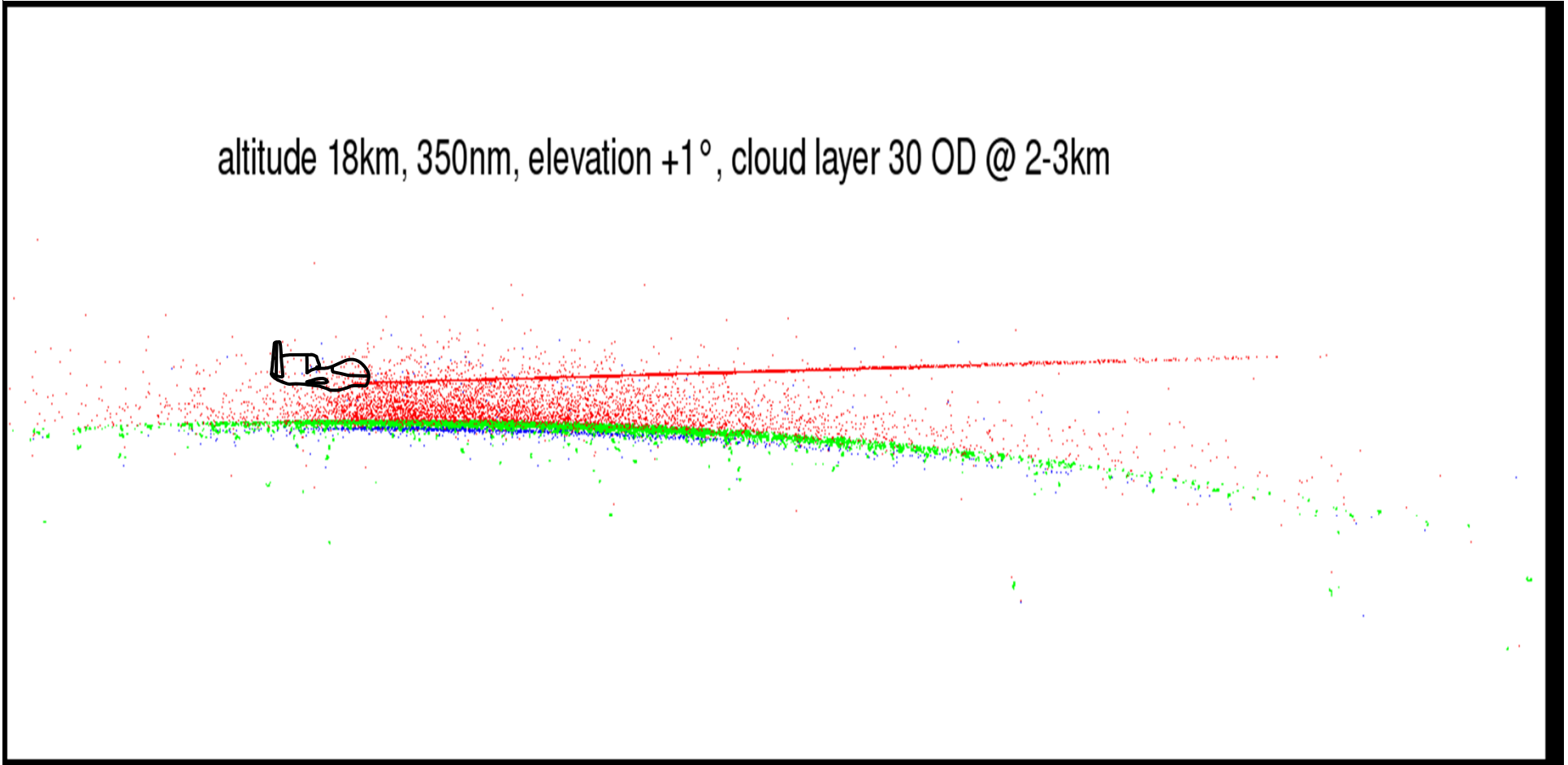
GH belly cam picture

Cloud Physics LIDAR (CPL)



Scattering Events - Visualisation

altitude 18km, 350nm, elevation +1°, cloud layer 30 OD @ 2-3km



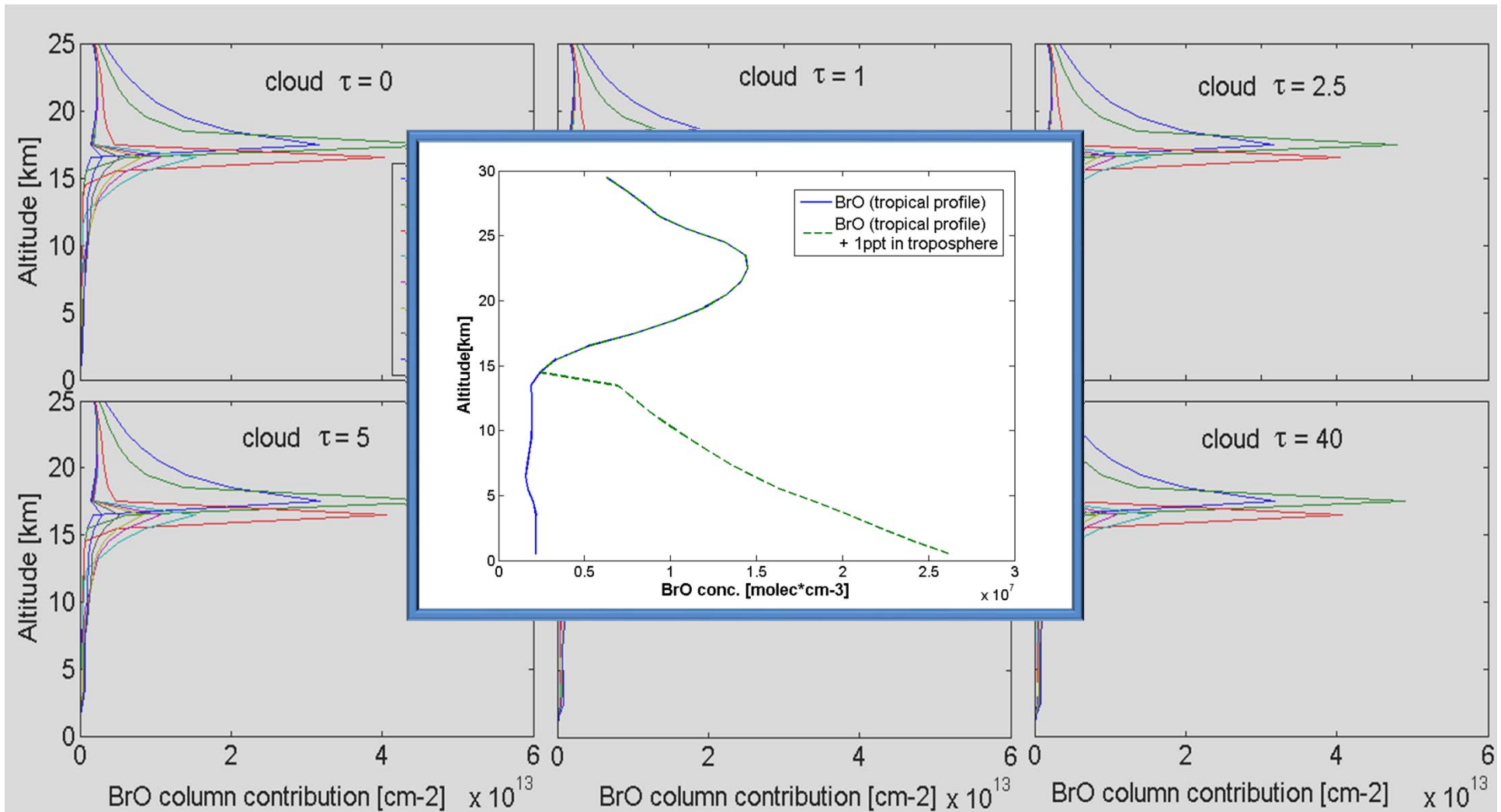
[Tim Deutschmann, PhD thesis, 2013]

Red: Rayleigh scattering events

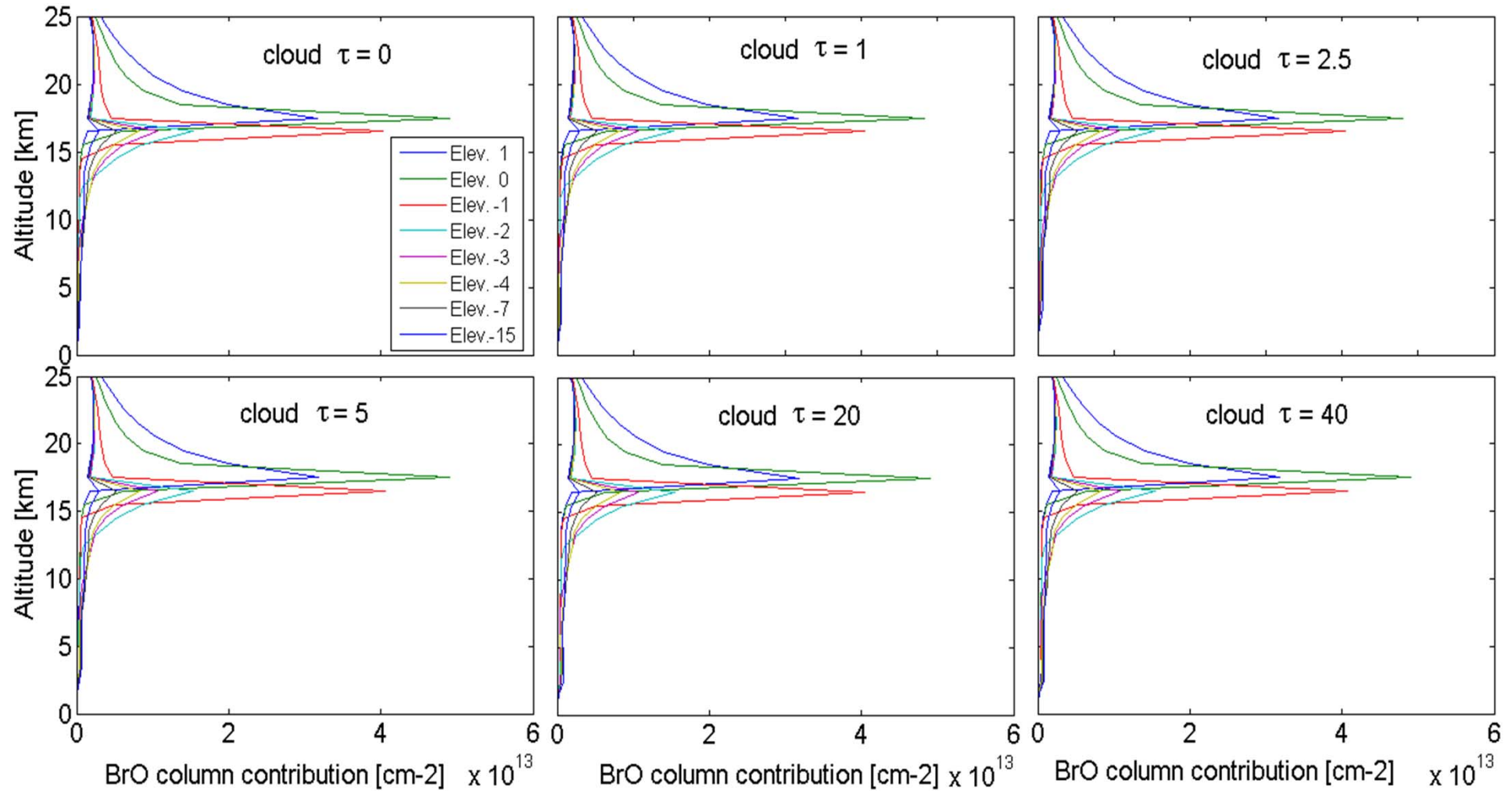
Green: Mie scattering events

Blue: Ground reflection

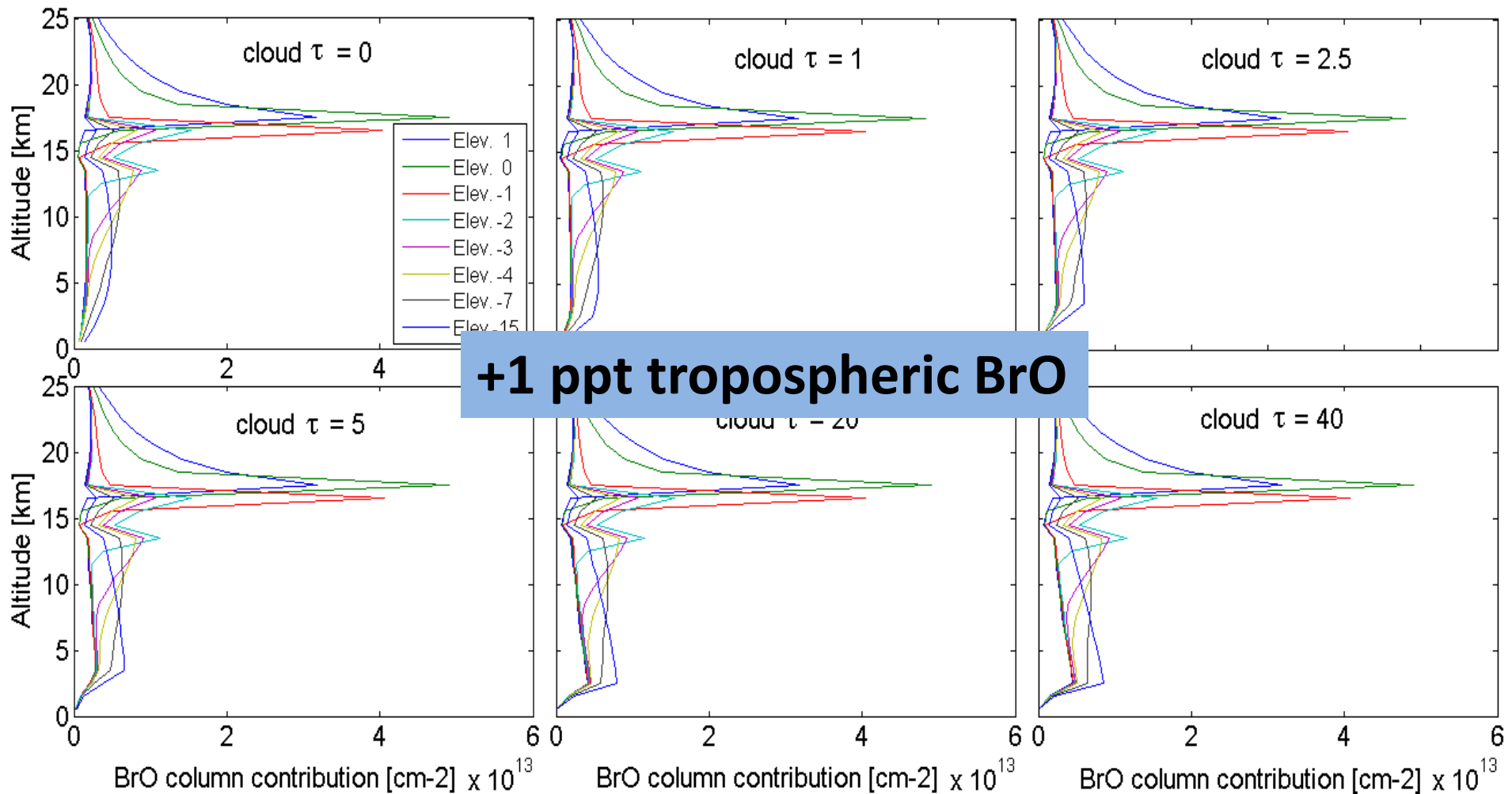
Influence of MSC Clouds on BrO Box Slant Columns



Influence of MSC Clouds on BrO Box Slant Columns

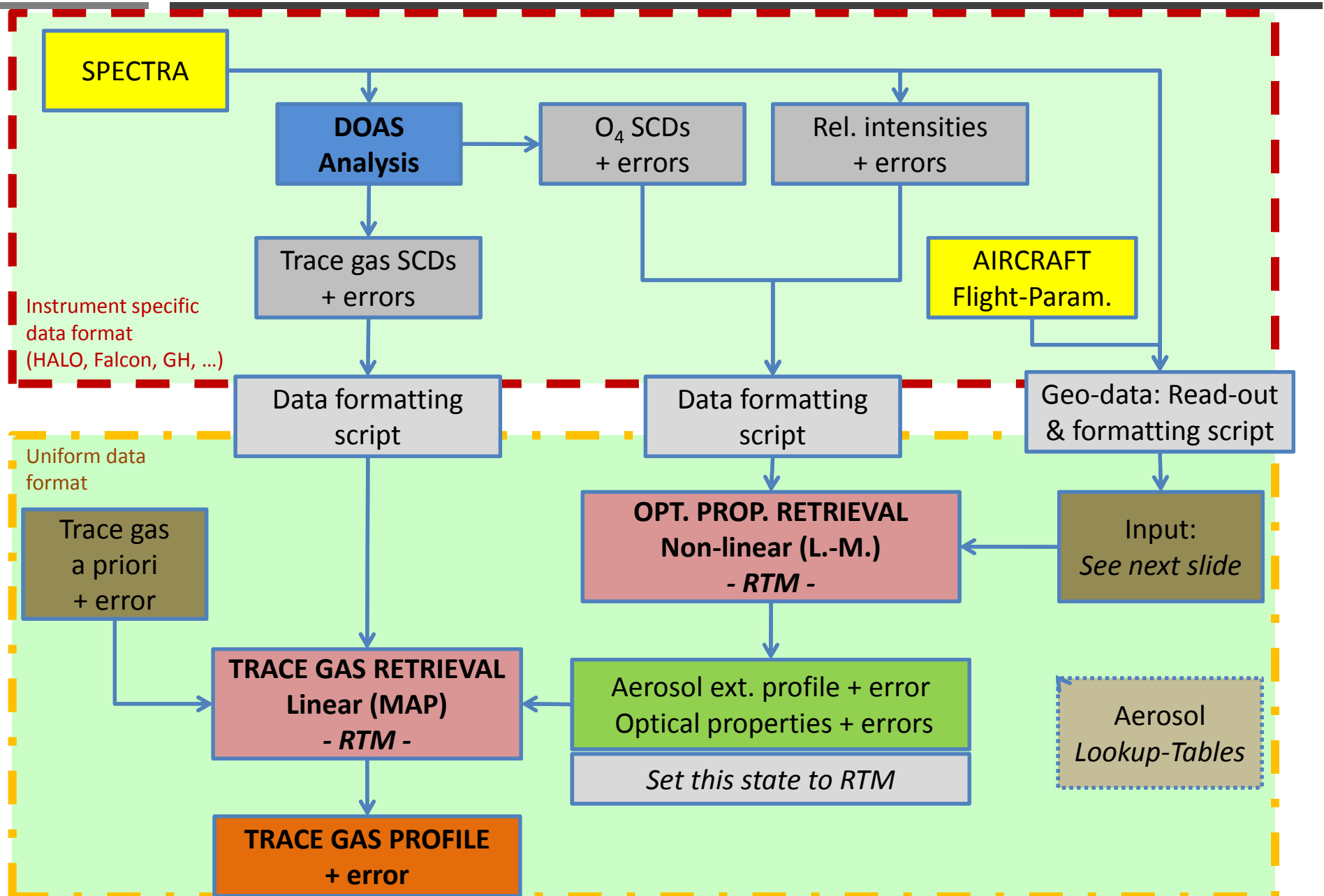


Influence of MSC Clouds on BrO Box Slant Columns

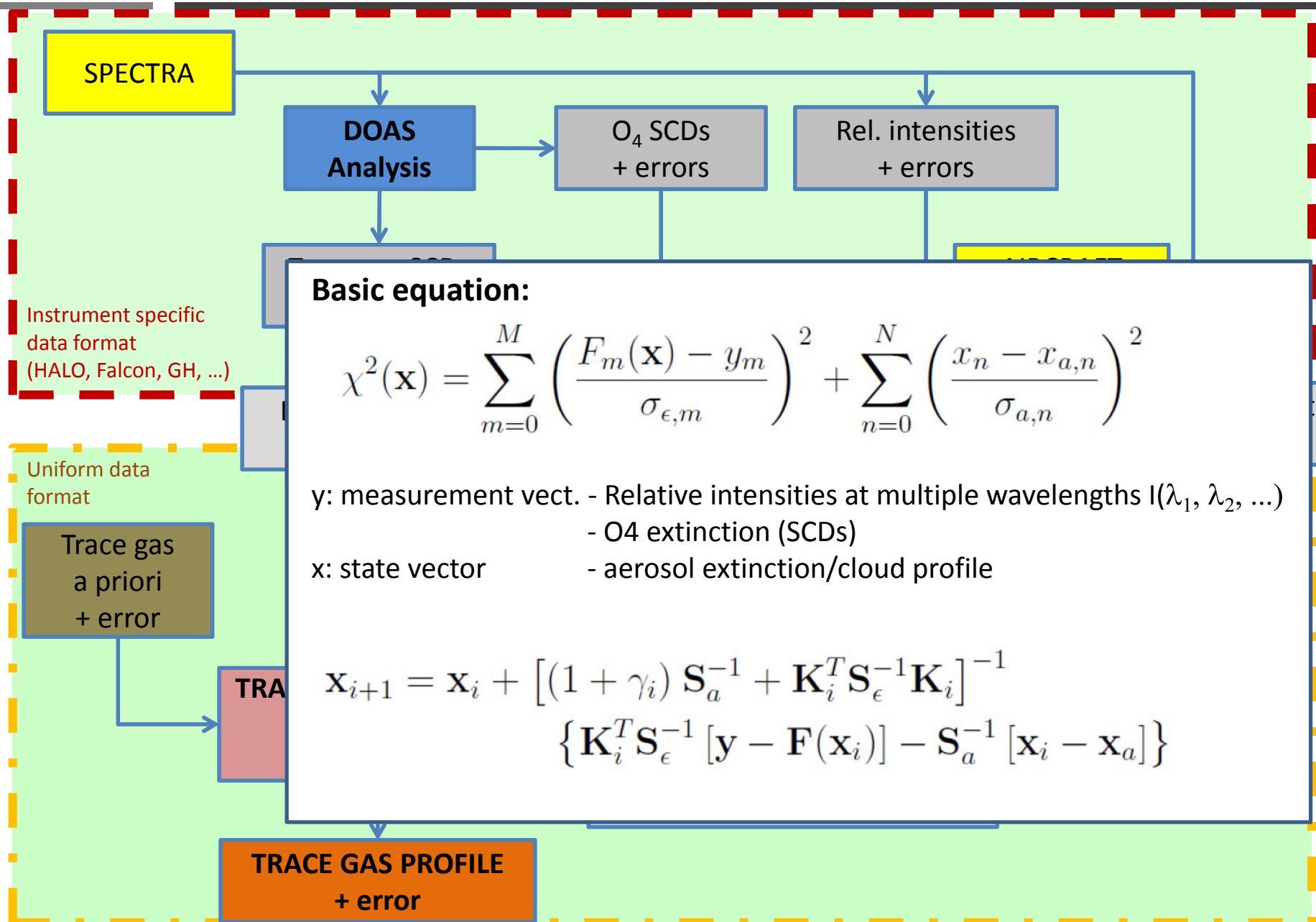


→ The error propagation of the tropospheric cloud retrieval into the BrO retrieval can eventually be significant

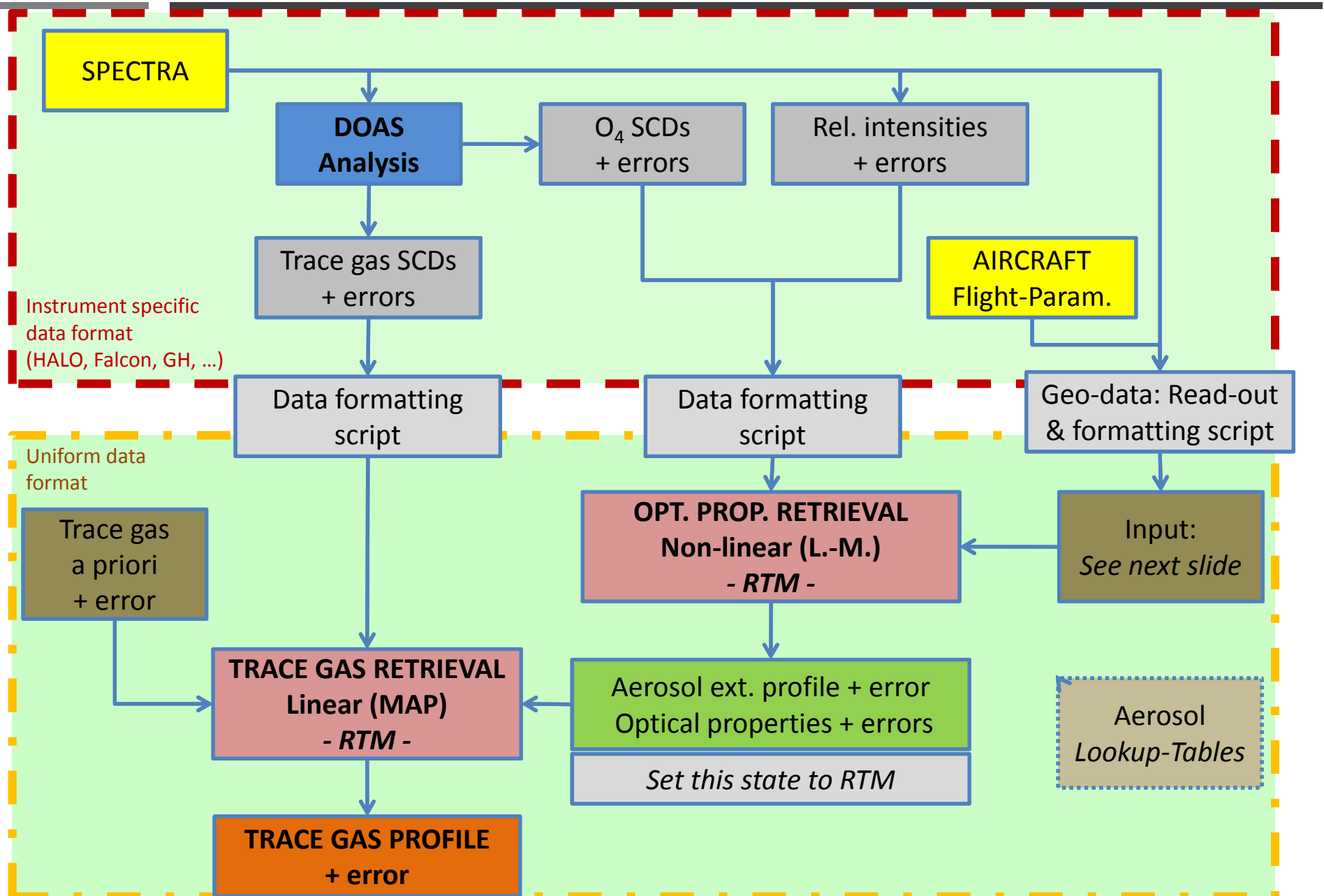
Retrieval Flow Chart



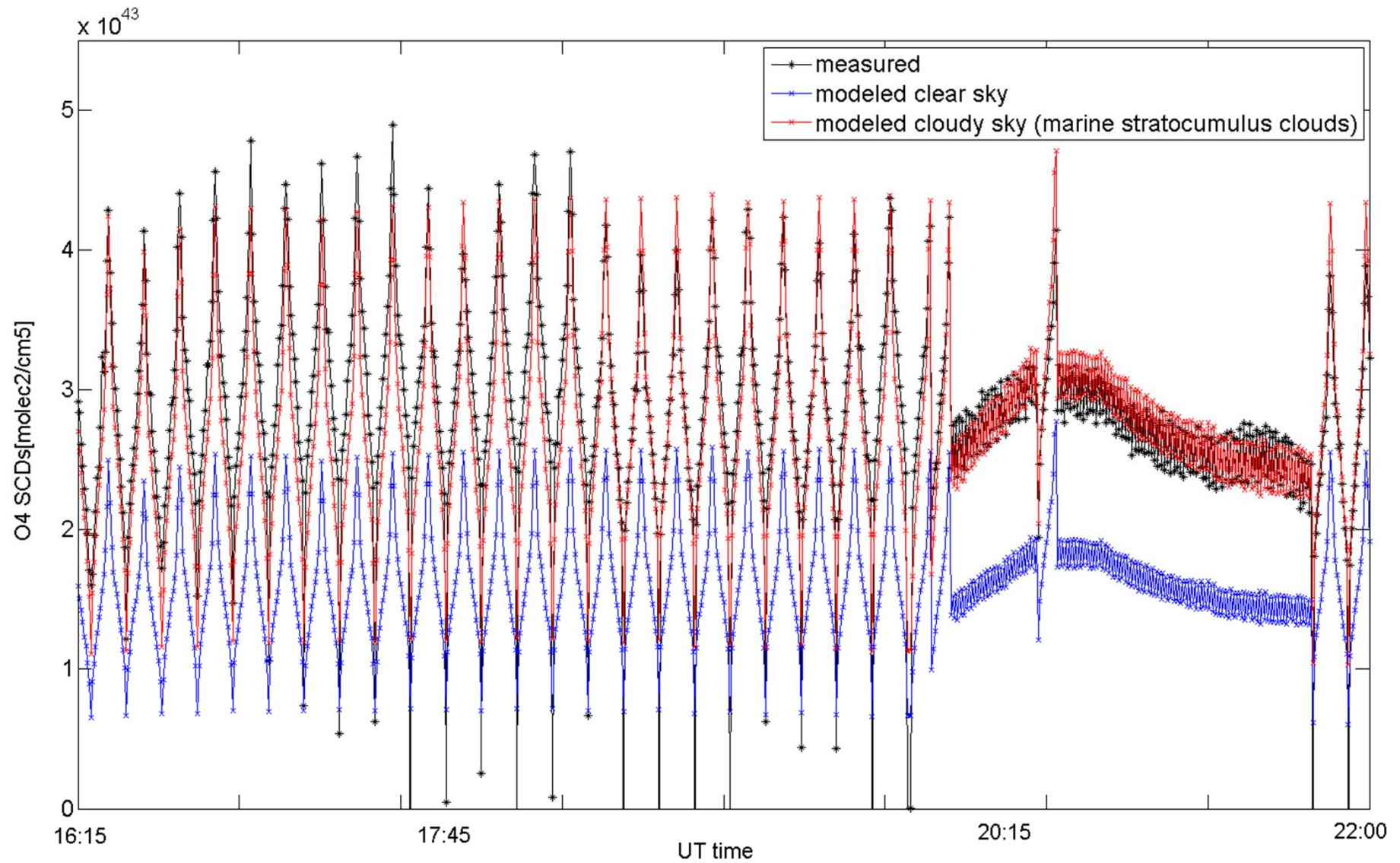
Retrieval Flow Chart



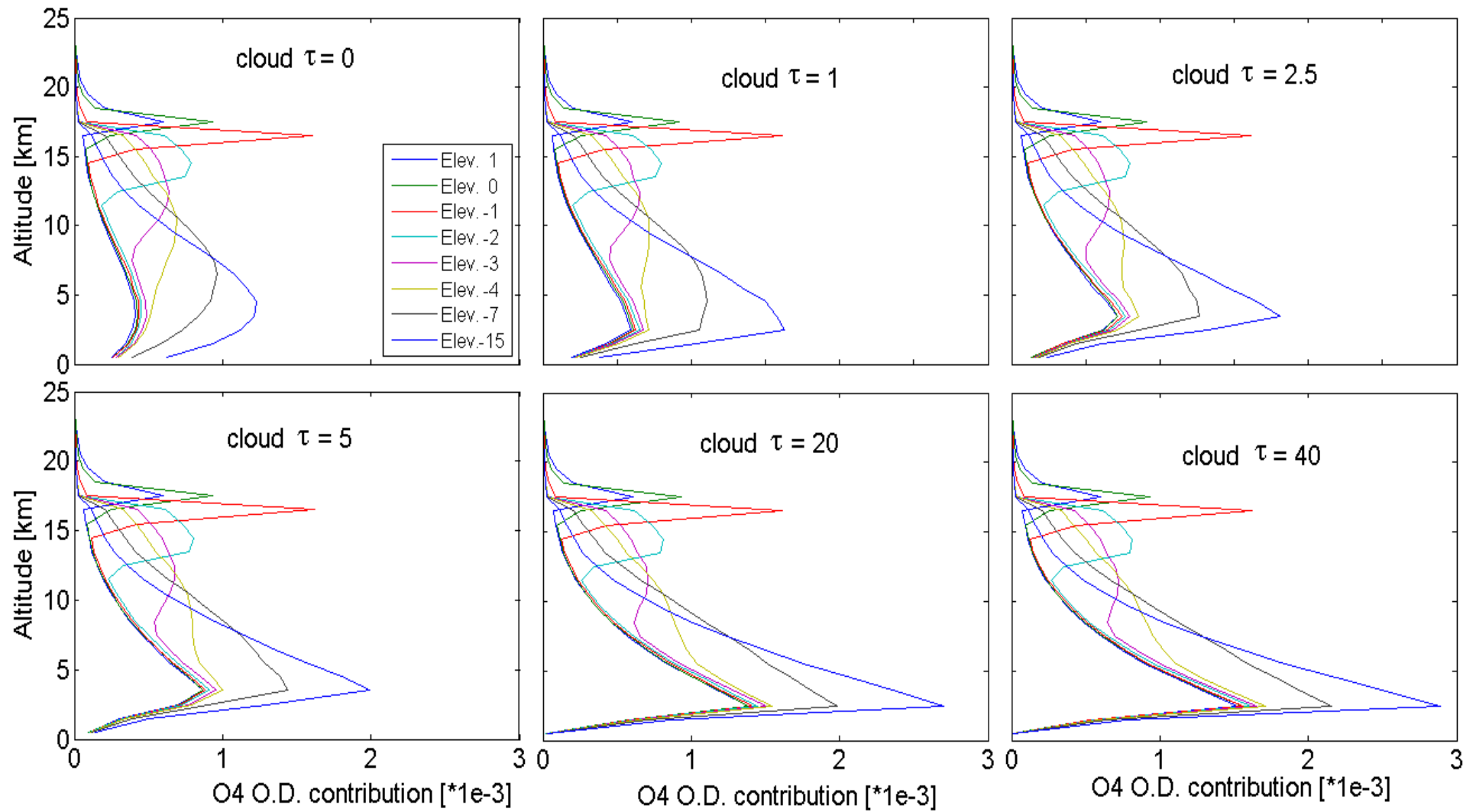
Retrieval Flow Chart



O4 361nm Extinctions – Clear and Cloudy Sky (MSC)

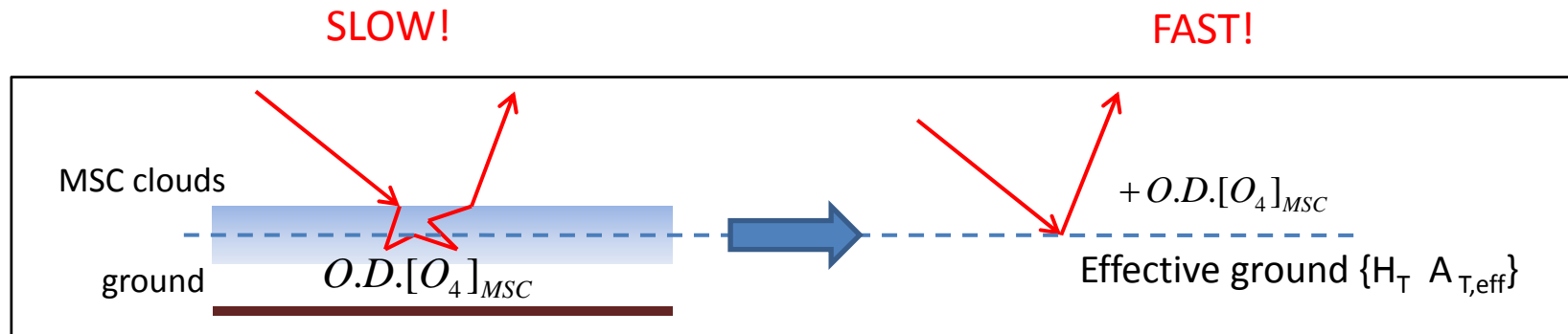


O4 361nm Extinctions, MSC Clouds at 2km



•Long computing time

Effective Model for Low Clouds



$$O.D.[O_4]_{MSC} = k_{eq} \cdot \sigma \cdot [O_2]^2 \cdot \tau \cdot d_{MSC} \cdot l_{Mie} (1 - g)$$

→ Retrieve effective albedo $A_{T,eff}$ at effective cloud top height H_T

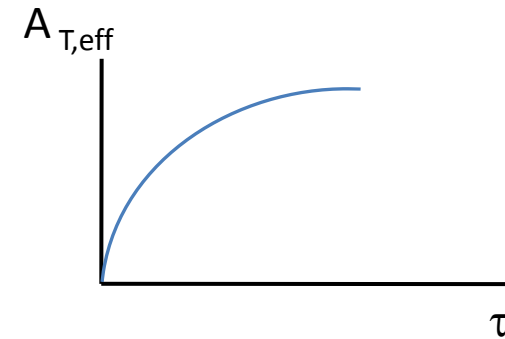
Needed derivatives:

(from RTM)

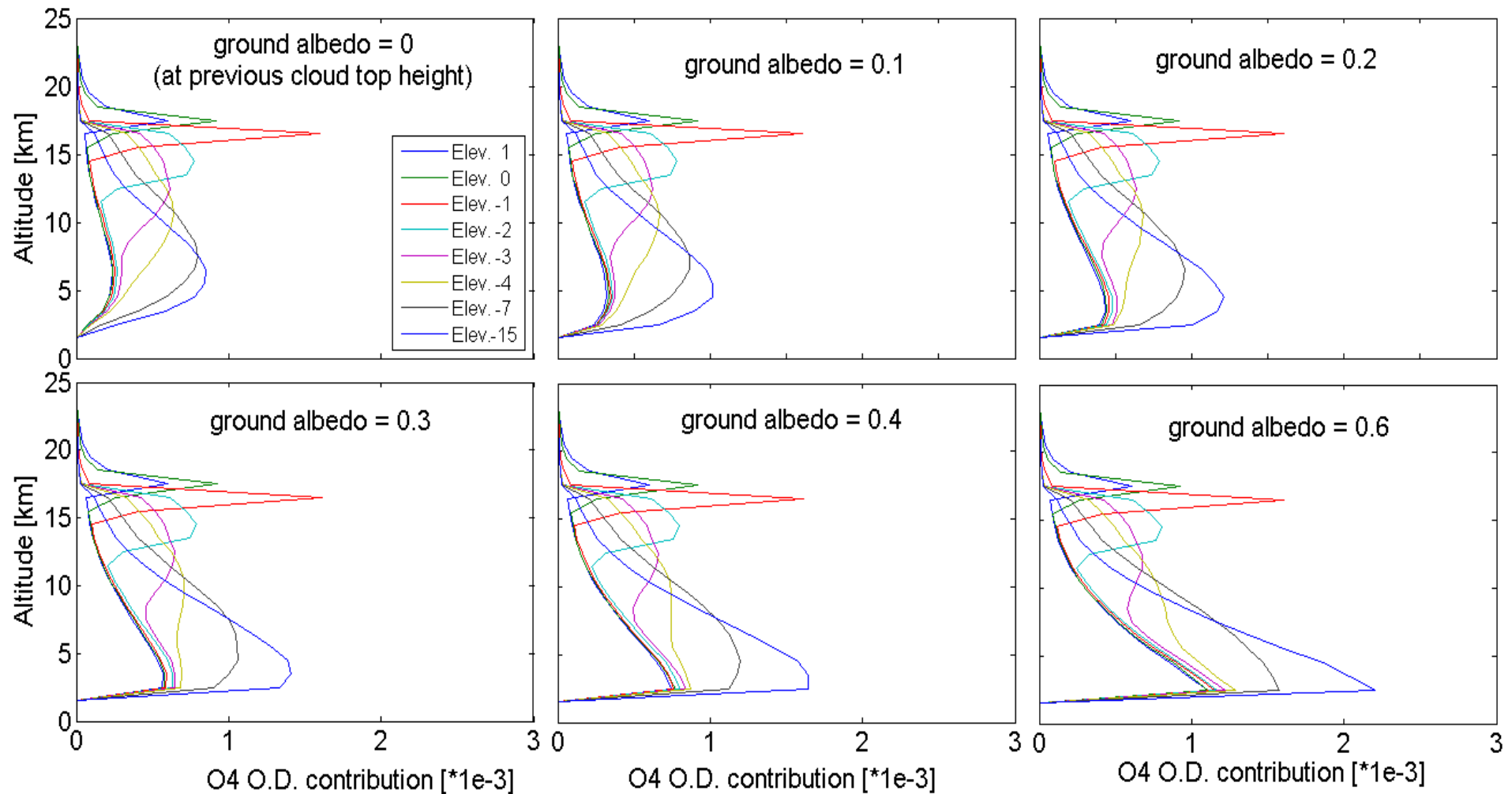
$$\frac{\partial\{O_4, I_{rel}(\lambda_i)\}}{\partial\{A_T, H_T\}_{eff}} = \frac{\partial\{O_4, I_{rel}(\lambda_i)\}}{\partial\tau} \cdot \frac{\partial\tau}{\partial\{A_T, H_T\}_{eff}}$$

$$\{A_T, H_T\}_{eff}(\tau) \Big|_{\{O_4, I_{rel}(\lambda_i)\}} \rightarrow \tau(\{A_T, H_T\}) \Big|_{\{O_4, I_{rel}(\lambda_i)\}}$$

↑
derivative



O4 Extinctions, Ground at 2km with Effective Albedo



•Parametrized differently -> very short computing time

Conclusion & Outlook

1. Long horizontal light paths (AMFs) supports sensitive detection species of extremely low (some ppt) concentrations (e.g., BrO, IO, NO₂, CH₂O, ...) in the UT/LS, TTL
2. Interference of the measurements comes with
 - Aerosol and cloud cover (cirrus) around flight level
 - Cloud cover (m.s.c.) / ground albedo
 - but these effects are minor/major on retrieval of gases with small/high concentration in the lower atmosphere i.e. NO₂, BrO, O₃, IO,... vs. O₄, H₂O
3. Major constraints for forward modelled RT are measured O₄, rel. intensities (and potentially the Ring effect)

Current / next steps

4. Set-up of a suitable forward model
 - Test stability of the cloud retrieval with synthetic scenarios
 - Quantify influence of UT/LS aerosols&clouds
5. Optimal estimation retrieval
6. Validate and compare results at flight altitude with in-situ data (e.g. AWAS)

Thanks for your attention!

Questions and comments are very welcome

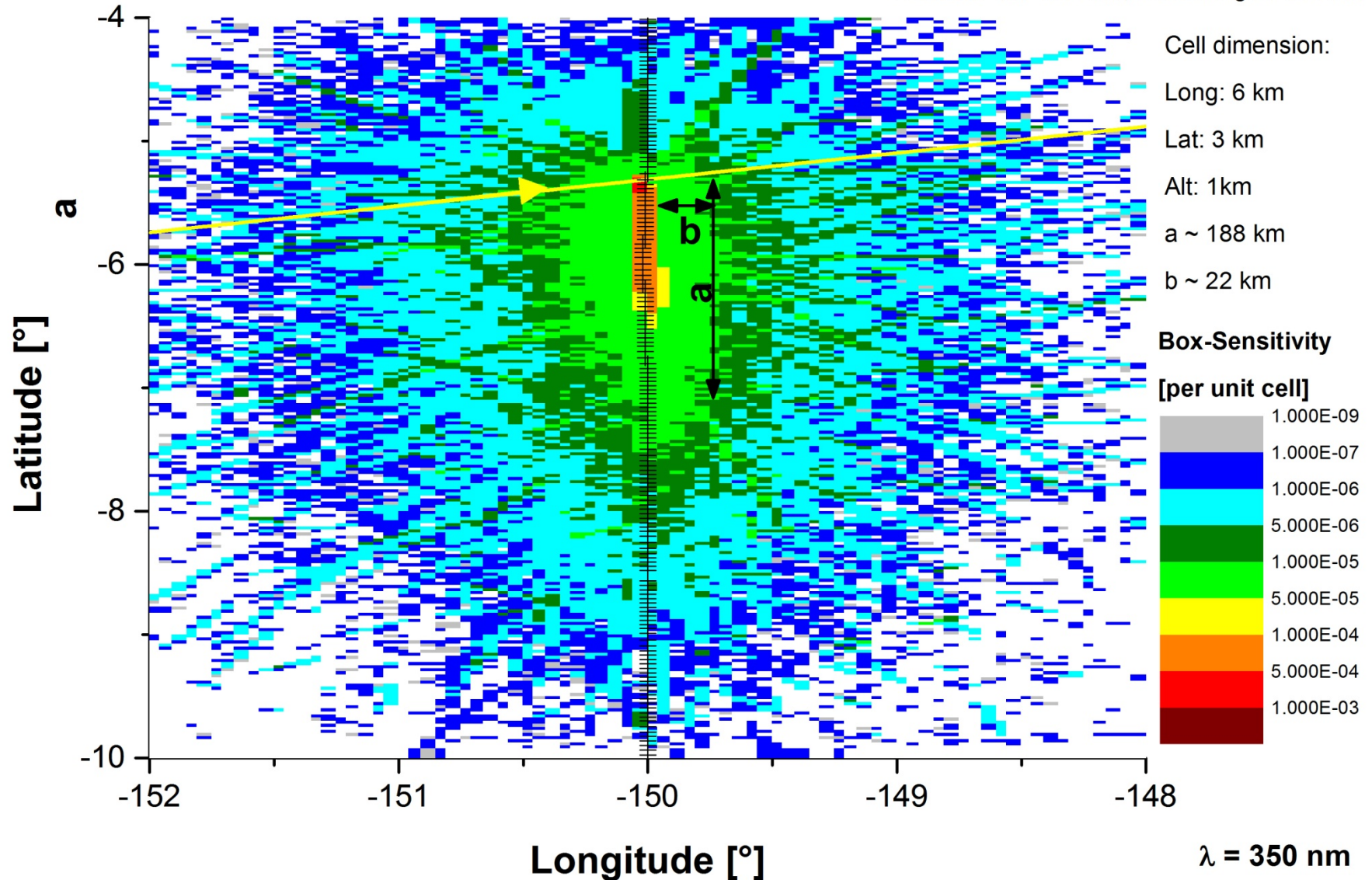
Additional slides

Retrieval – Horizontal Sensitivity

GH SF2 2013 - Horizontal Box Sensitivities

Displayed altitude layer: 18.0 ~ 19.0 km

Detector alt: 18.2 km, Wavelength: 350 nm



Retrieval – Relative Intensities and O4

Basic equation:

$$\chi^2(\mathbf{x}) = \sum_{m=0}^M \left(\frac{F_m(\mathbf{x}) - y_m}{\sigma_{\epsilon,m}} \right)^2 + \sum_{n=0}^N \left(\frac{x_n - x_{a,n}}{\sigma_{a,n}} \right)^2$$

y: measurement vect. - Relative intensities at multiple wavelengths $I(\lambda_1, \lambda_2, \dots)$

- O4 extinction (SCDs)

x: state vector

- aerosol extinction profile, SSA

Goal:

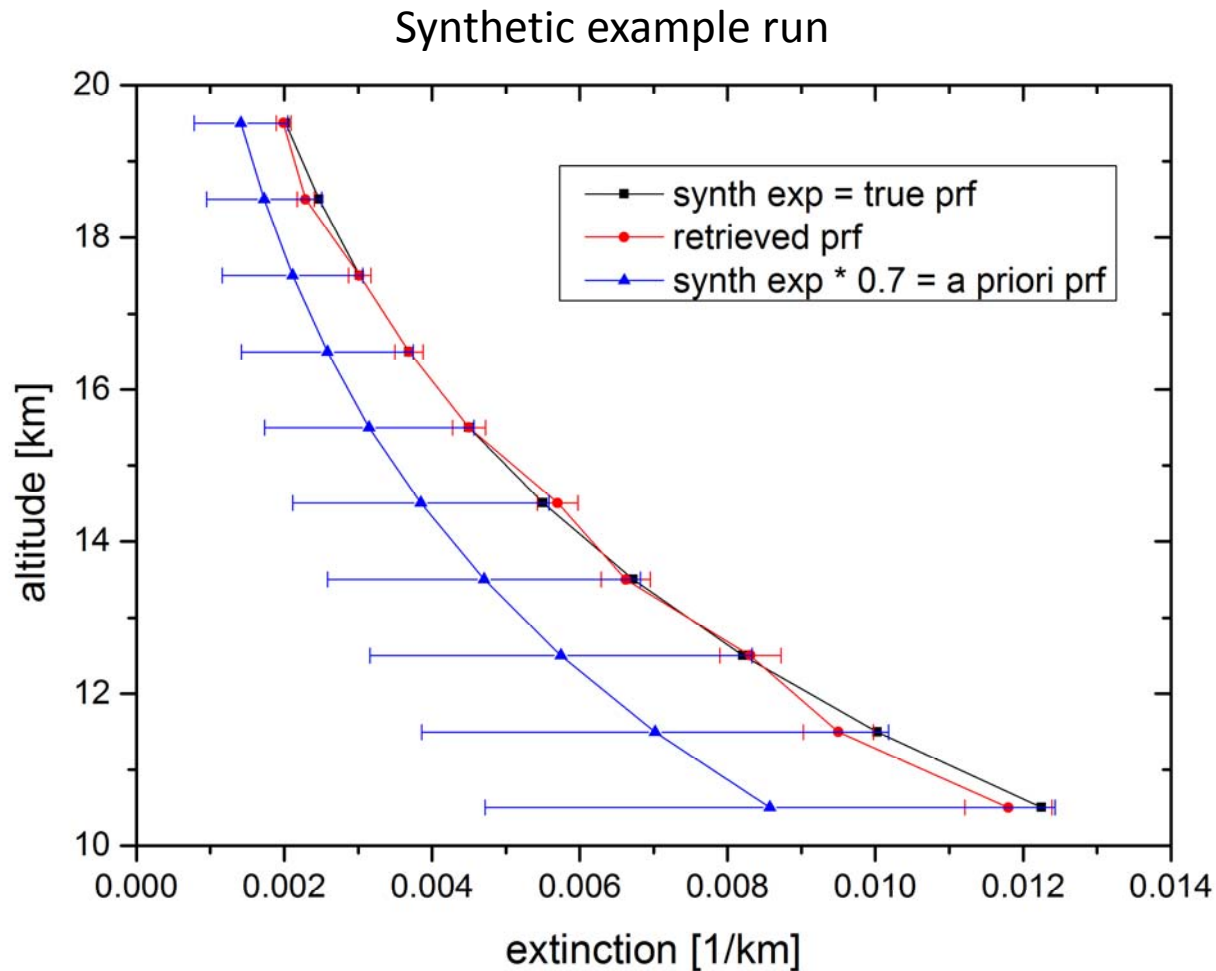
Minimize χ^2 which is the error-weighted difference between measurement and forward model output and the difference between the actual state and the a priori state

The forward model \mathbf{F} depends in a non-linear way on \mathbf{x}

→ χ^2 has to be minimized iteratively (Levenberg-Marquard-Algorithm)

$$\mathbf{x}_{i+1} = \mathbf{x}_i + \left[(1 + \gamma_i) \mathbf{S}_a^{-1} + \mathbf{K}_i^T \mathbf{S}_\epsilon^{-1} \mathbf{K}_i \right]^{-1} \left\{ \mathbf{K}_i^T \mathbf{S}_\epsilon^{-1} [\mathbf{y} - \mathbf{F}(\mathbf{x}_i)] - \mathbf{S}_a^{-1} [\mathbf{x}_i - \mathbf{x}_a] \right\}$$

Retrieval – Synthetic Test



y: relative Intensities at three wavelengths (320, 350, 380nm)

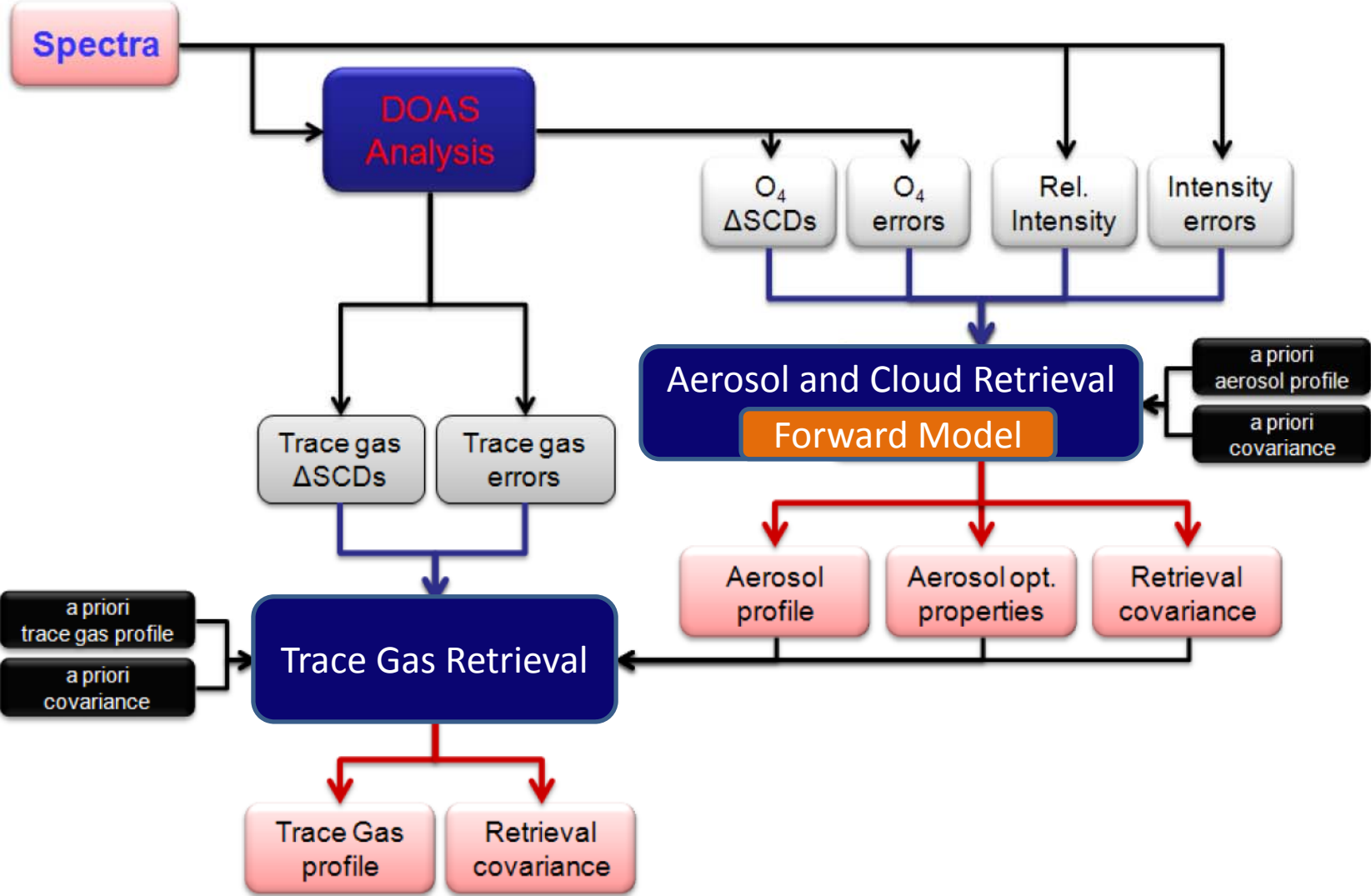
Information gained at different wavelengths are combined with an Angstrom model

$$\rightarrow \tau(\lambda_1) / \tau(\lambda_2) = (\lambda_1 / \lambda_2)^{-\alpha}$$

The retrieved profile agrees well with the true profile

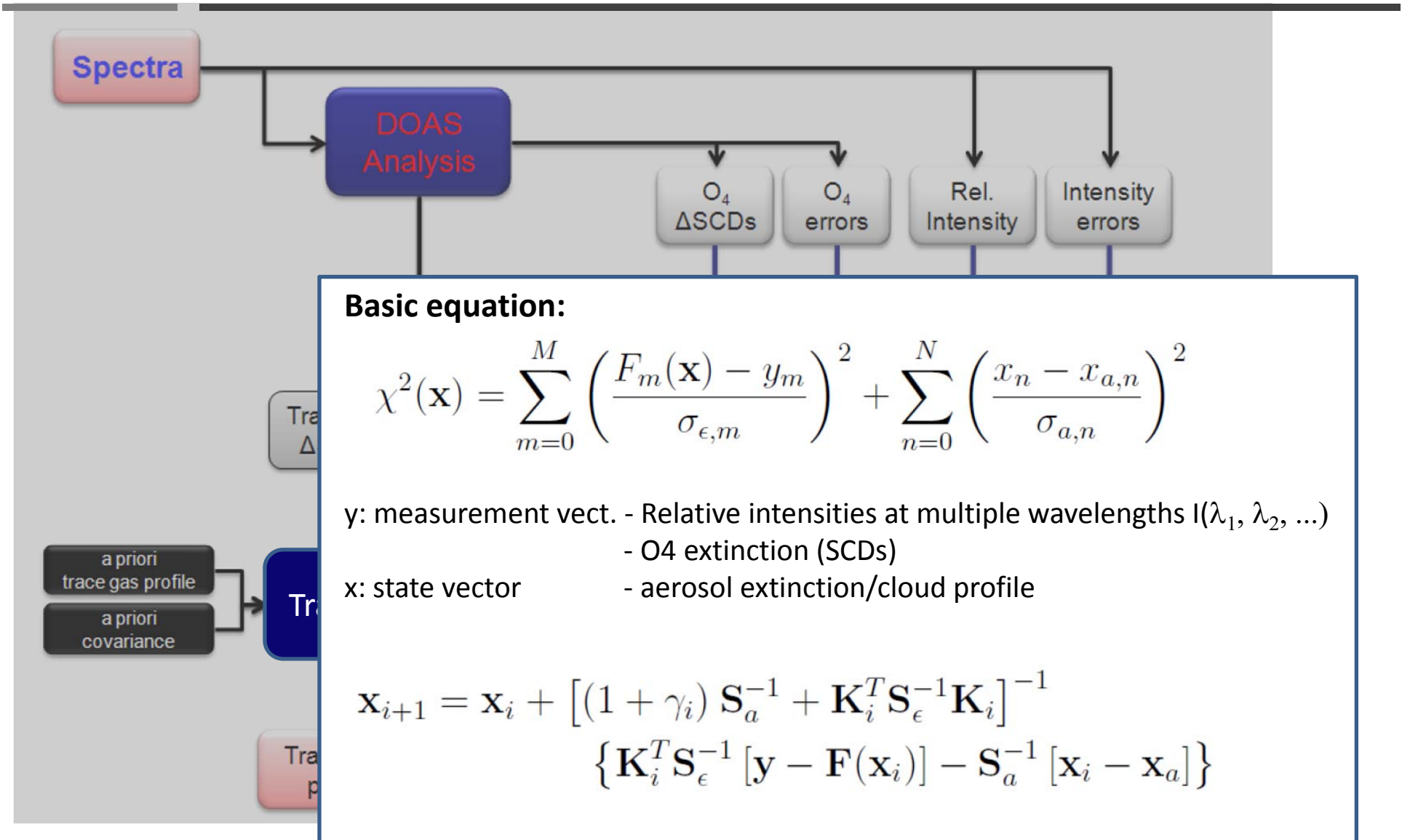
→ The technical implementation and the Jacobians are correct

Retrieval Flow Chart

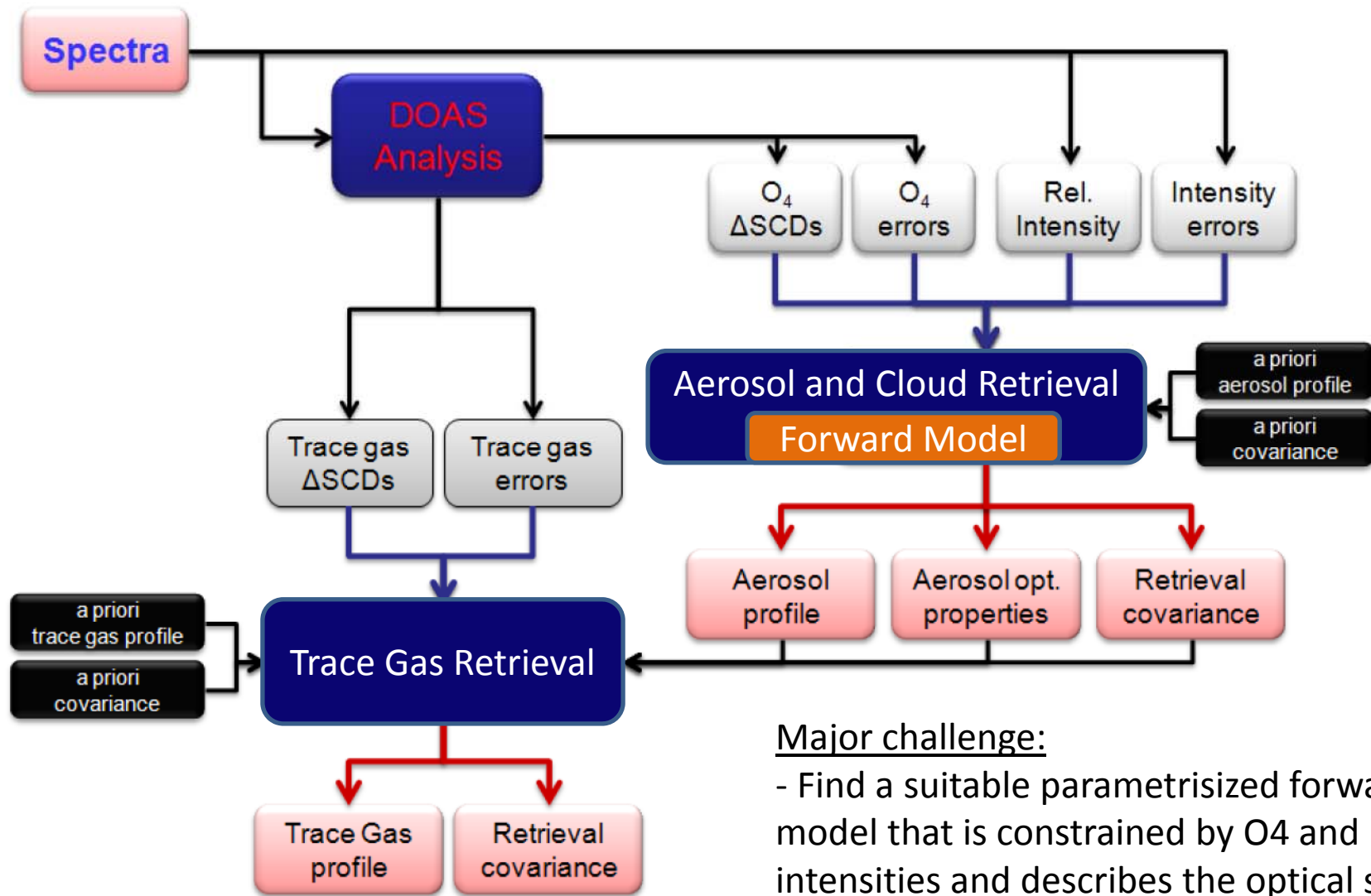


[Yilmaz, PhD thesis 2012]

Retrieval Flow Chart



Retrieval Flow Chart



Major challenge:

- Find a suitable parametrized forward model that is constrained by O₄ and relative intensities and describes the optical state of the atmosphere in a sufficient way to invert BrO concentration profiles